

Representation of the warm conveyor belt of the Stalactite Cyclone in convective- permitting simulation of Meso-NH

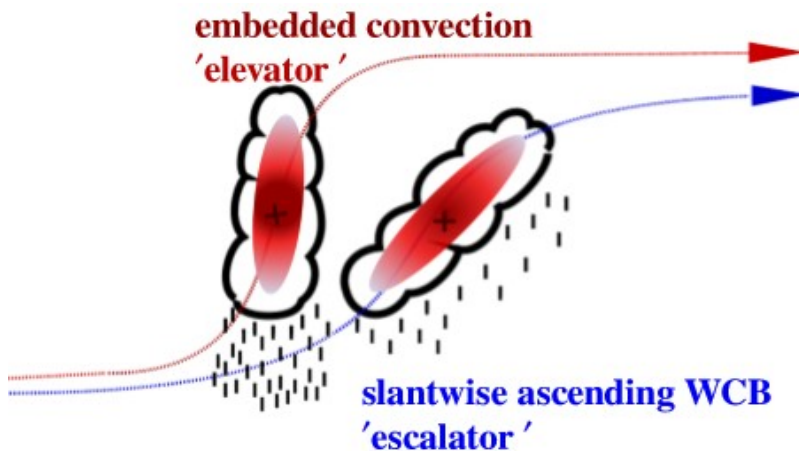
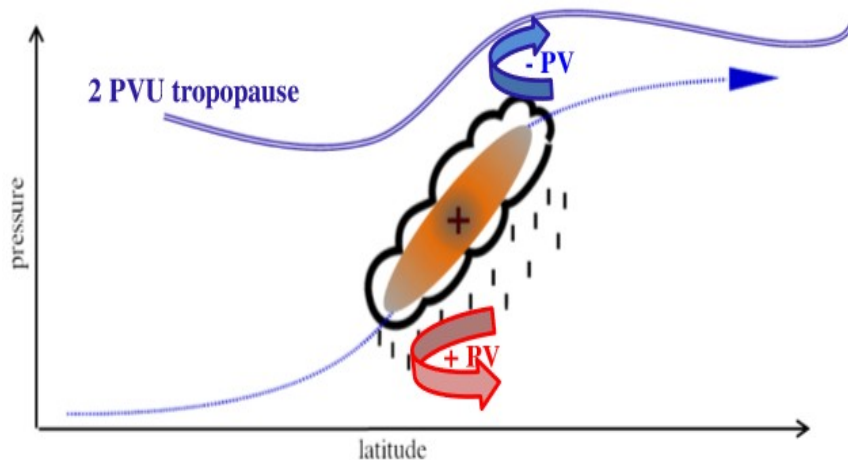
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Motivation



Conceptual model by Oertel et al. (2019)

- WCBs are continuously rising air masses which typically ascent 600 hPa in 48 h
(Browning, 1986; Wernli and Davies, 1997; Madonna et al., 2014)
- Cloud diabatic processes in WCBs can impact the cyclone dynamics and the downstream large-scale circulation
(Hoskins et al. 1985; Joos and Wernli, 2012; Binder et al. 2016, Joos and Forbes, 2016;)
- Recent studies with convection resolving models suggest the occurrence of rapid ascents embedded in some WCBs cases and can also be relevant for the large-scale dynamics
(Rasp et al., 2016; Oertel et al., 2019a; Oertel et al., 2019b)
- Cloud processes and convection are parameterized in NWP models: suspected to be sources of forecast error growth over the Atlantic
(Davies and Didone, 2013; Selz and Craig, 2015)

1) What are the characteristics of the ascents in the WCB of the Stalactite cyclone?

2) What are the processes leading to fast ascents?

Simulation and observations



simulations:

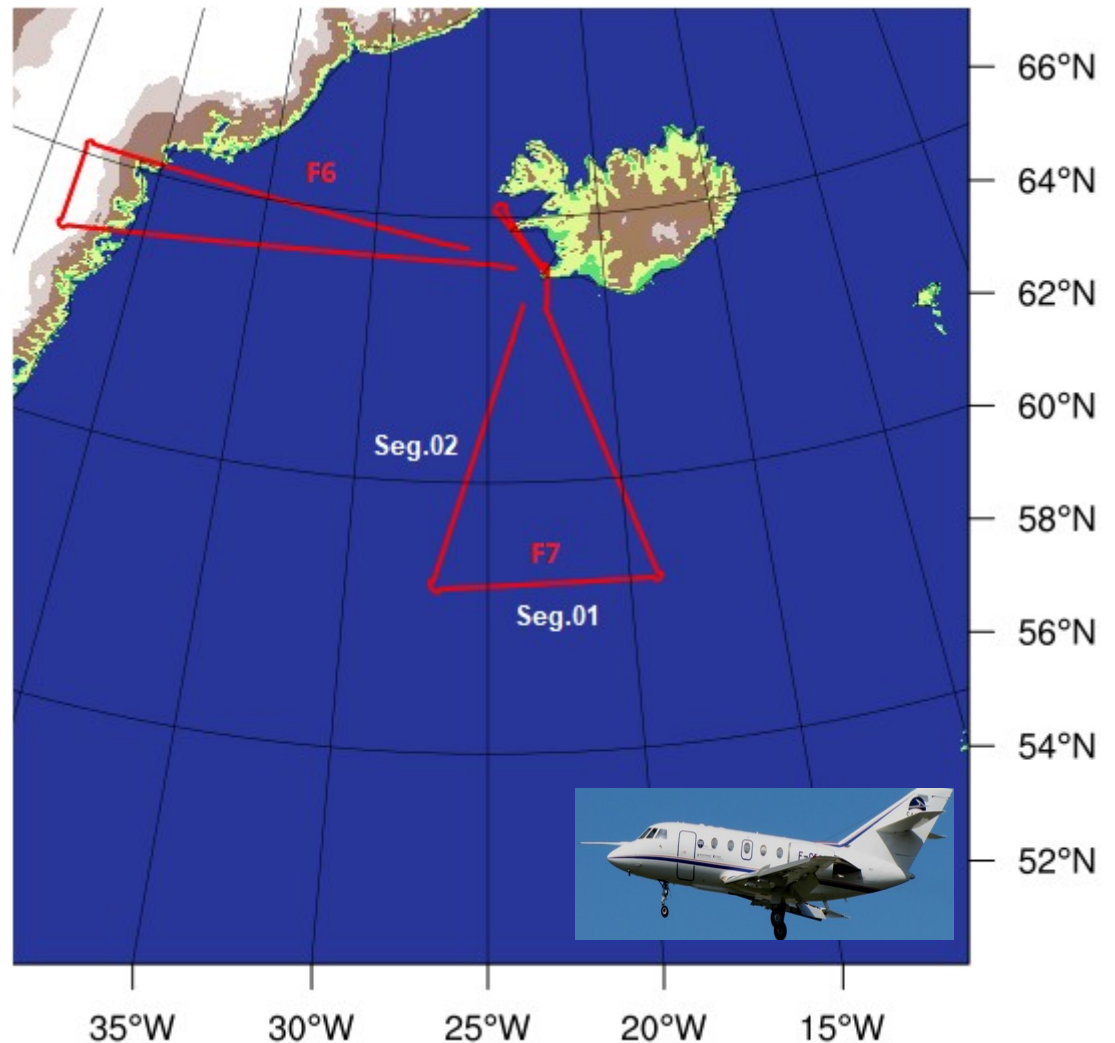
- $\Delta x = 2.5$ km
→ convection permitting
- 800x800 grid points
- start at 00 UTC, 2 October 2016
ECMWF analyses; $t = 36$ h

WCB ascents identification in the reference simulation REF:

→ Lagrangian online trajectories calculated from 3D wind at every model time step ($\Delta t = 20$ s)

→ Clustering tool in Meso-NH

Dauhut et al., 2016



Observations:

- MSG satellite images ($10.8 \mu\text{m}$)
- RASTA Doppler radar (3 mm)

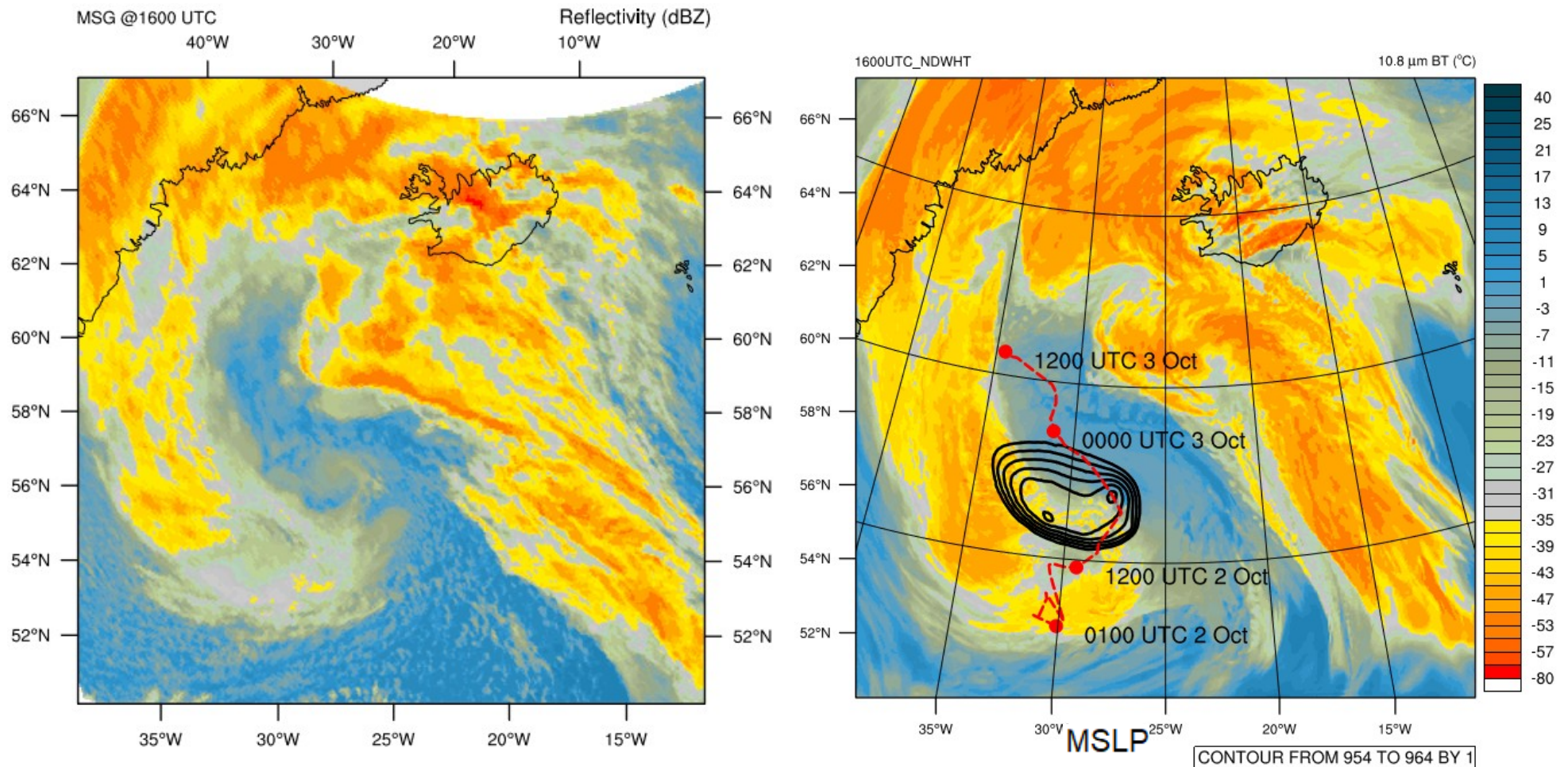
Delanoë et al., 2007

F6: 0900 UTC – 1130 UTC → WCB outflow

F7: 1300 UTC – 1600 UTC → WCB ascents

1. What are the characteristics of the ascents in the WCB of the Stalactite cyclone?

Stalactite cyclone track

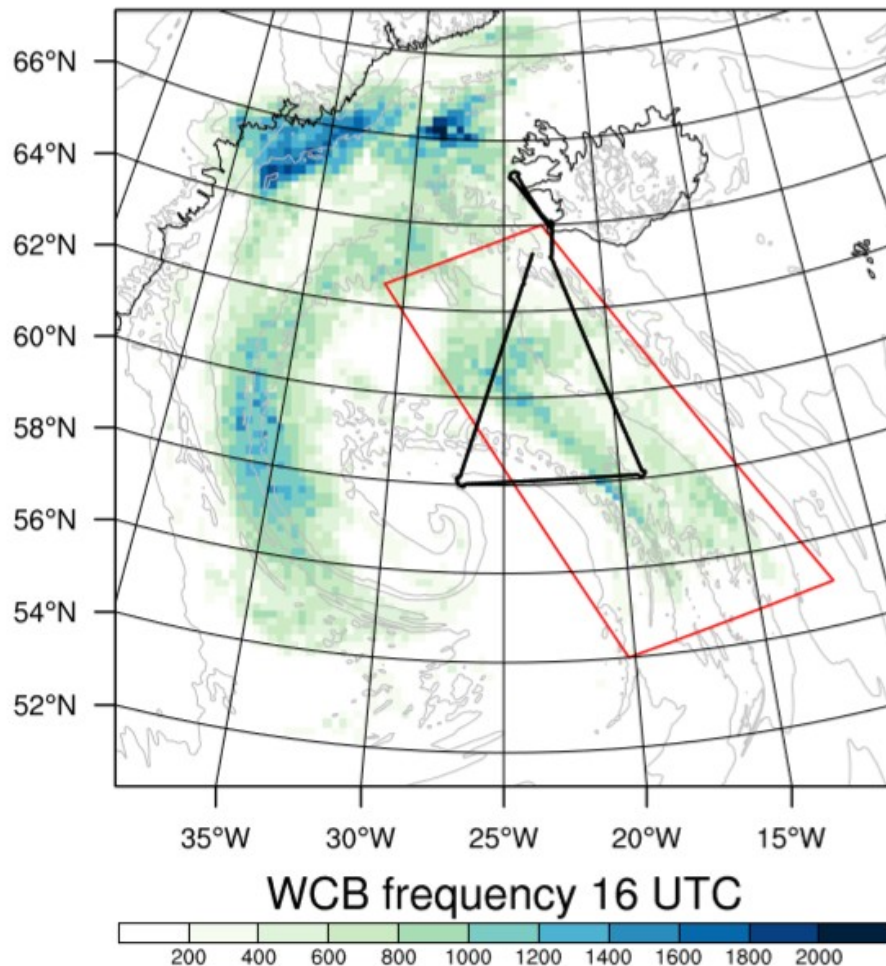


- Track and deepening well reproduced in REF
- Eastward shift of MSLP in REF (eg. cloud structure)

WCB identification at 1600 UTC

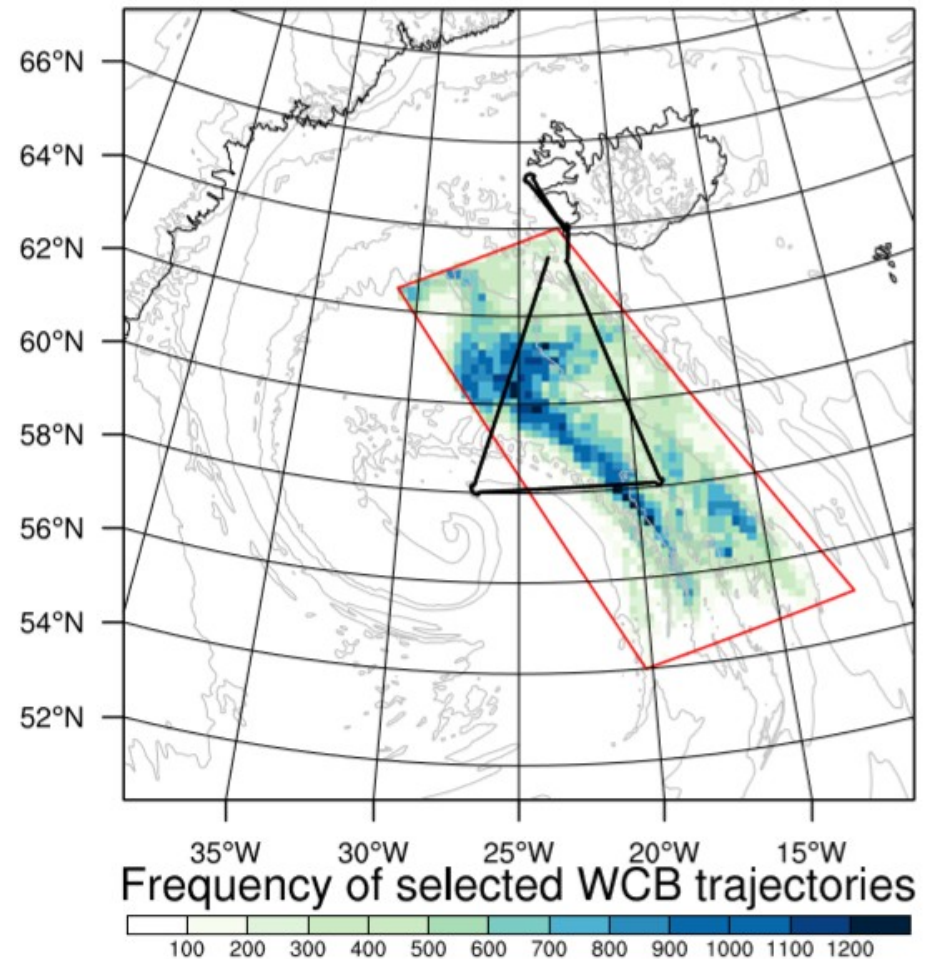
- **WCB:** ΔP_{12h} = pressure variation between 1000 and 2200 UTC (Oct. 2)

$\Delta P_{12h} \leq -150 \text{ hPa}$ (e.g 600 hPa in 48 h ; Madonna et al. 2014)



Red box : selected trajectories

THTE contours at $z \sim 1 \text{ km}$ (K)



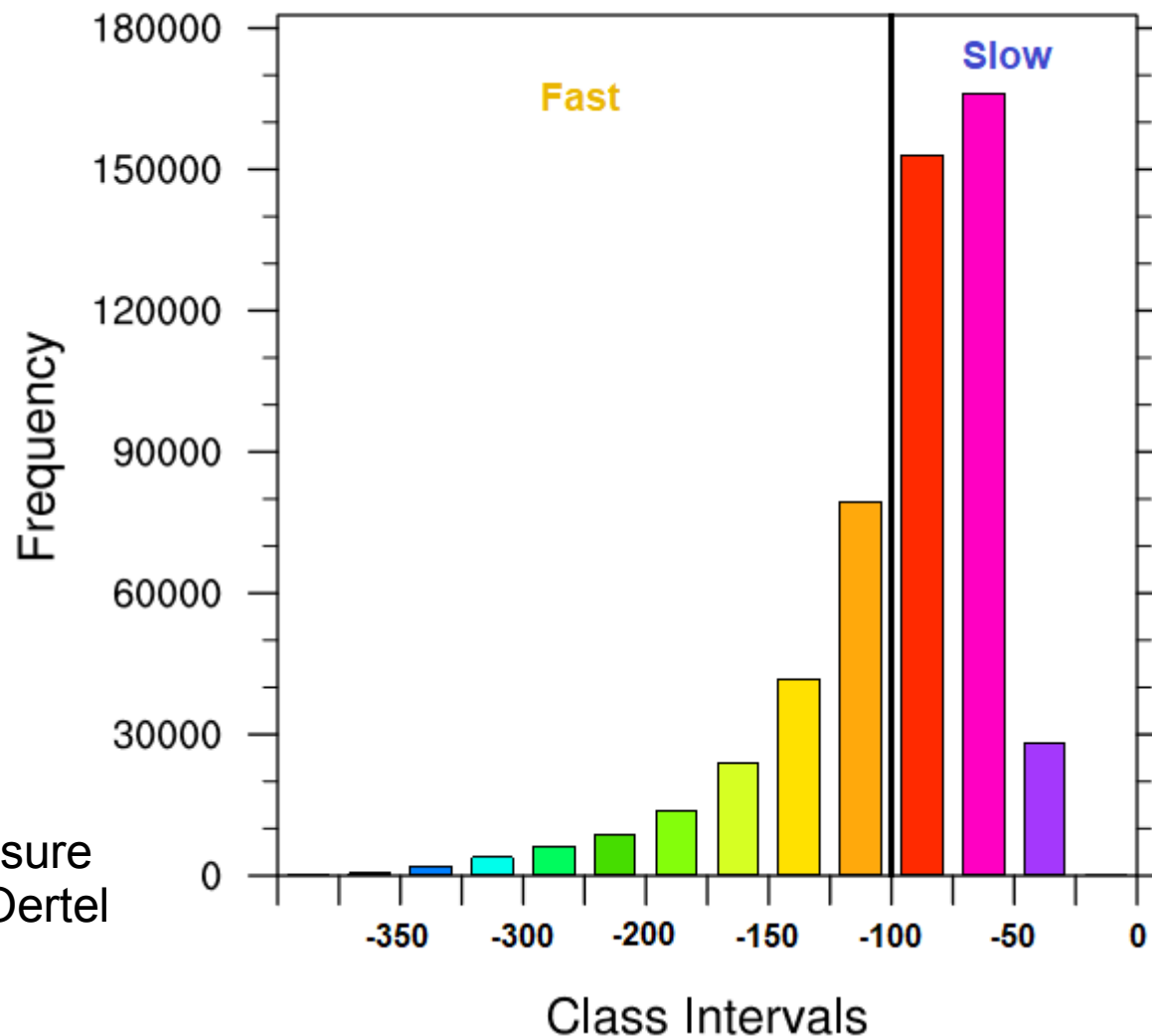
→ The plane flew over the WCB

Slow - Fast ascents identification in the red box

- **WCB:** ΔP_{12h} = pressure variation between 1000 and 2200 UTC (Oct. 2)

$\Delta P_{12h} \leq -150$ hPa (e.g 600 hPa in 48 h ; Madonna et al. 2014)

2h Pressure variation along WCB airmasses ($\text{hPa } 2\text{h}^{-1}$)



Fast ascents:

$$\Delta P_{2h} \leq -100 \text{ hPa}$$

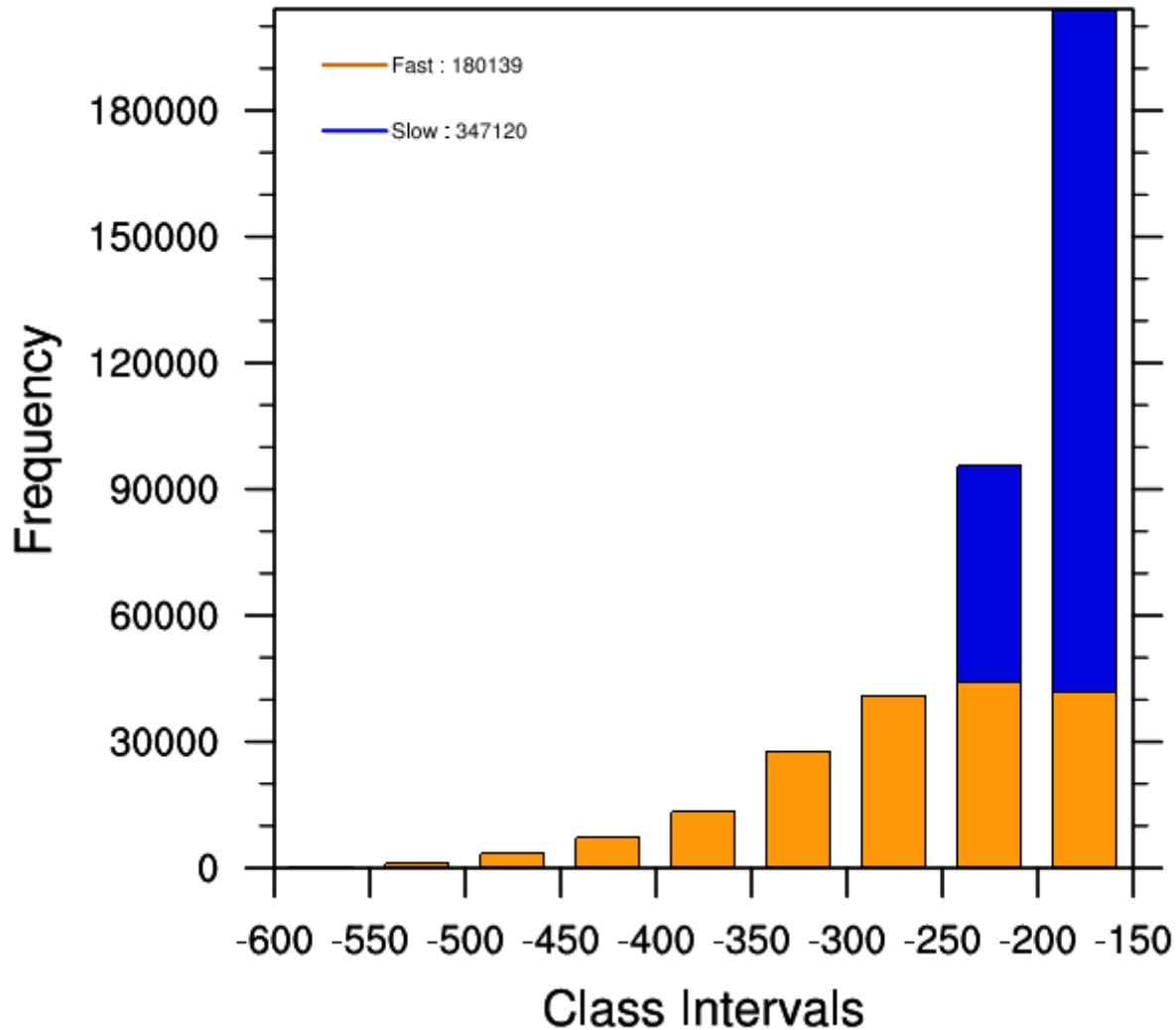
Slow ascents:

$$\Delta P_{2h} > -100 \text{ hPa}$$

Fast ascent pressure
criteria same as Oertel
et al. 2019a

Slow - Fast ascents elevation in 12 h evolution

Stacked histogram for 12h Pressure variation (hPa)

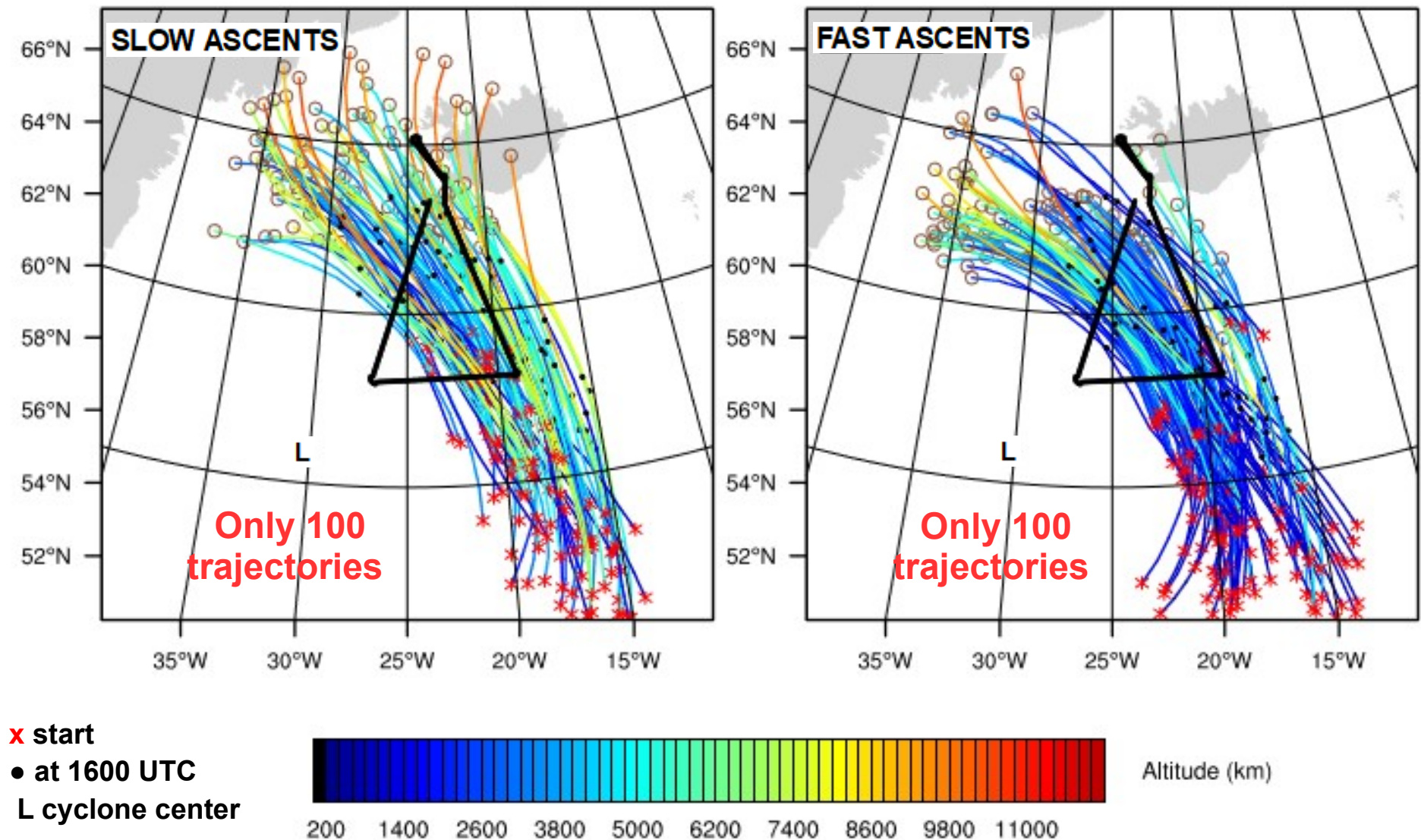


- $WCB_{tot} \sim 530\,000$ trajs
- **Slow** $\sim 2/3 WCB_{tot}$
- **Fast** $\sim 1/3 WCB_{tot}$

→ **Slow ascents**: $-150 \text{ hPa} \leq \Delta P_{12h} \leq -250 \text{ hPa}$

→ **Fast ascents**: $-150 \text{ hPa} \leq \Delta P_{12h} \leq -550 \text{ hPa}$

Air mass trajectories between 1000 – 2200 UTC



1a) What is the location of slow and fast ascents in the WCB?

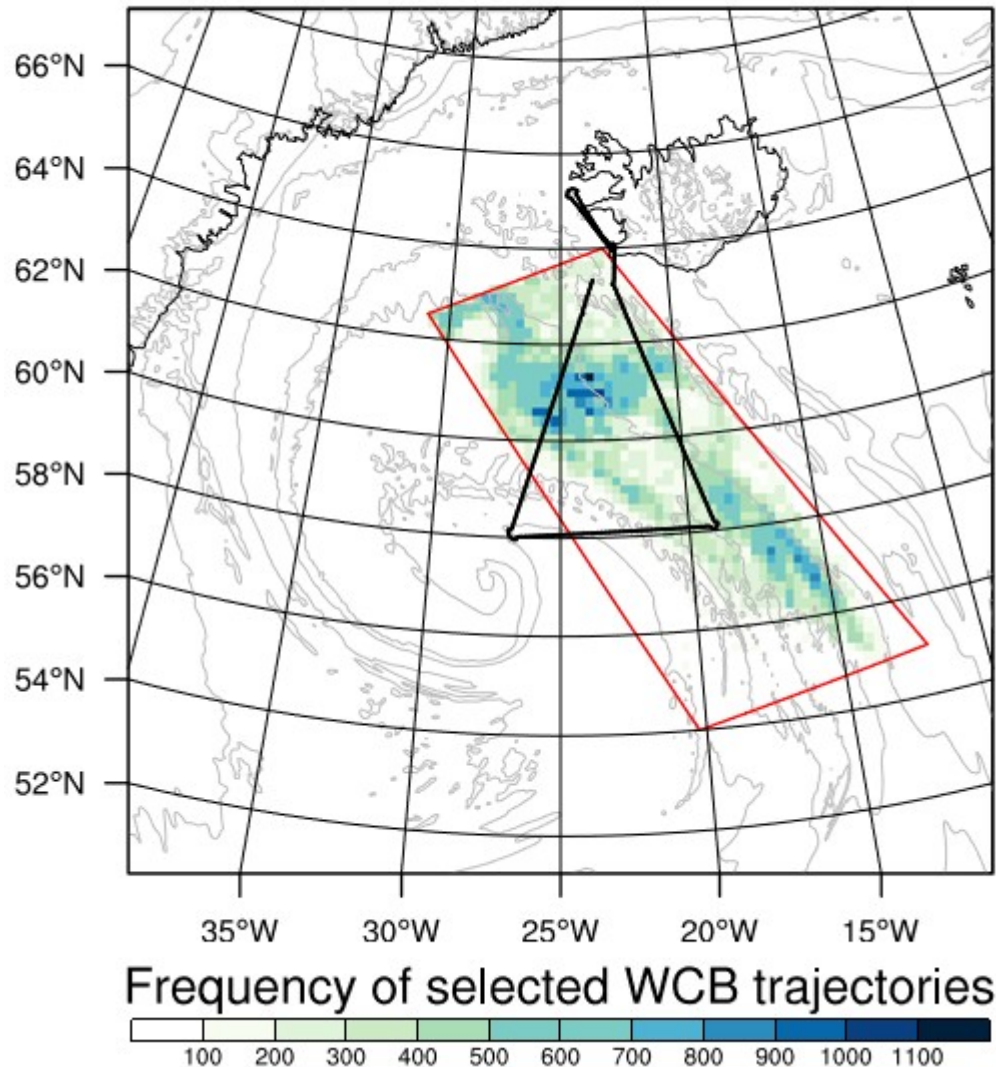
1b) How many fast and slow ascents have a cyclonic/anticyclonic trajectory?

1c) What is the altitude of fast and slow ascents?

1a) Slow – Fast ascents location at 1600 UTC

Slow: $-100 \text{ hPa} < \Delta P_{2h}$

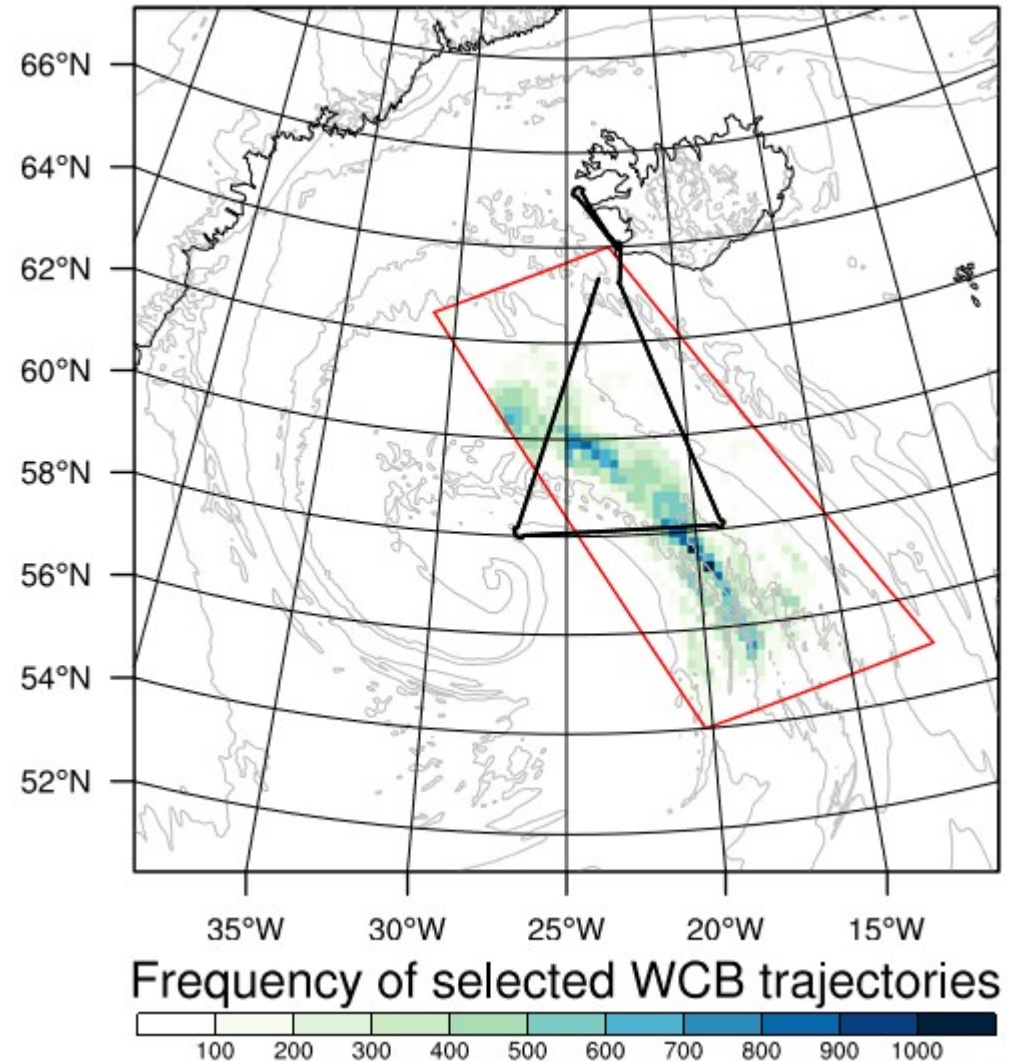
frequency_slow 16 UTC



→ Slow ascents in the core of WCB

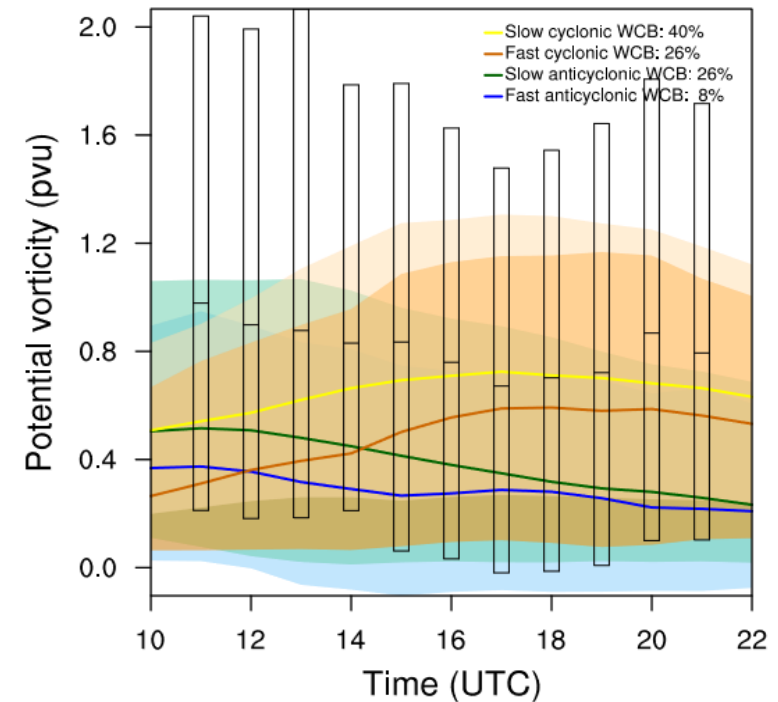
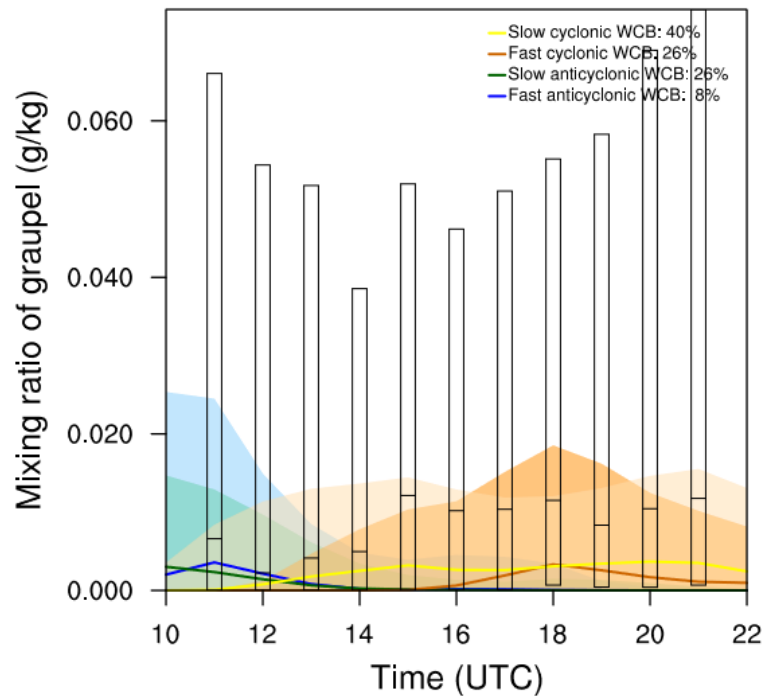
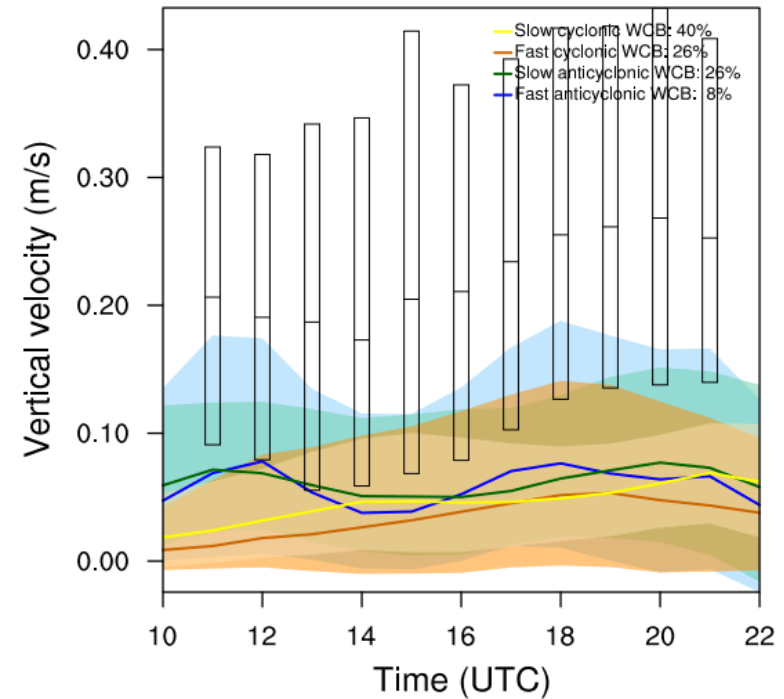
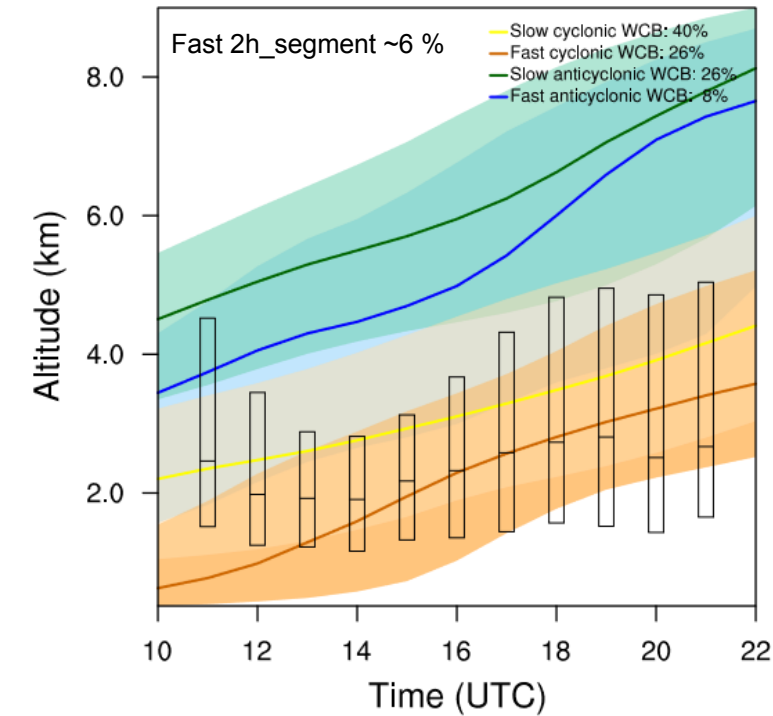
Fast: $\Delta P_{2h} \leq -100 \text{ hPa}$

frequency_fast 16 UTC



→ Fast ascents in the west side of WCB

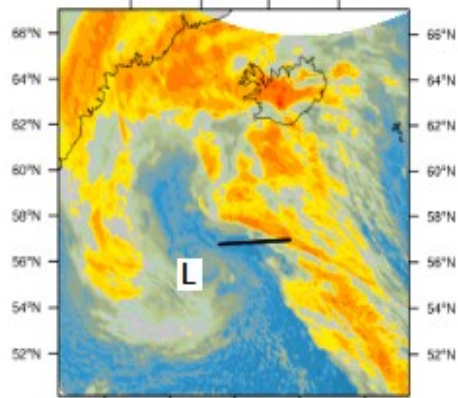
1b,c) Temporal evolution of air-mass characteristics



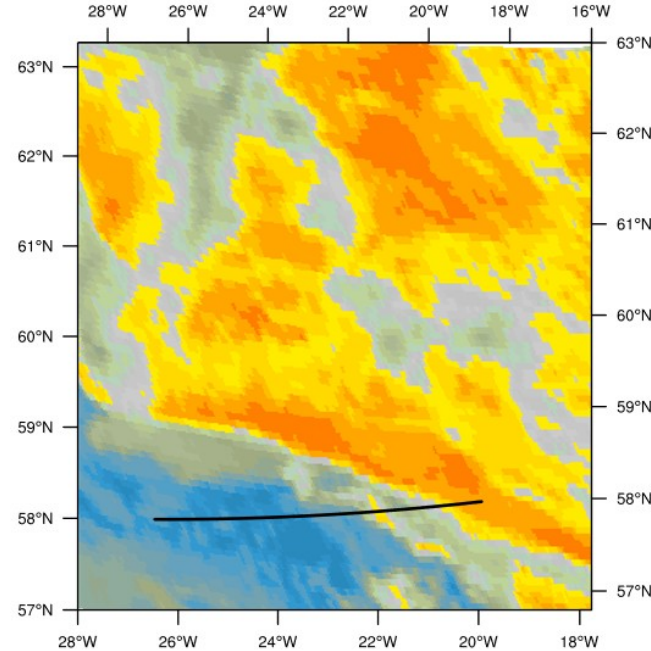
2. What are the processes leading to fast ascents?

Zoom at 1500 UTC

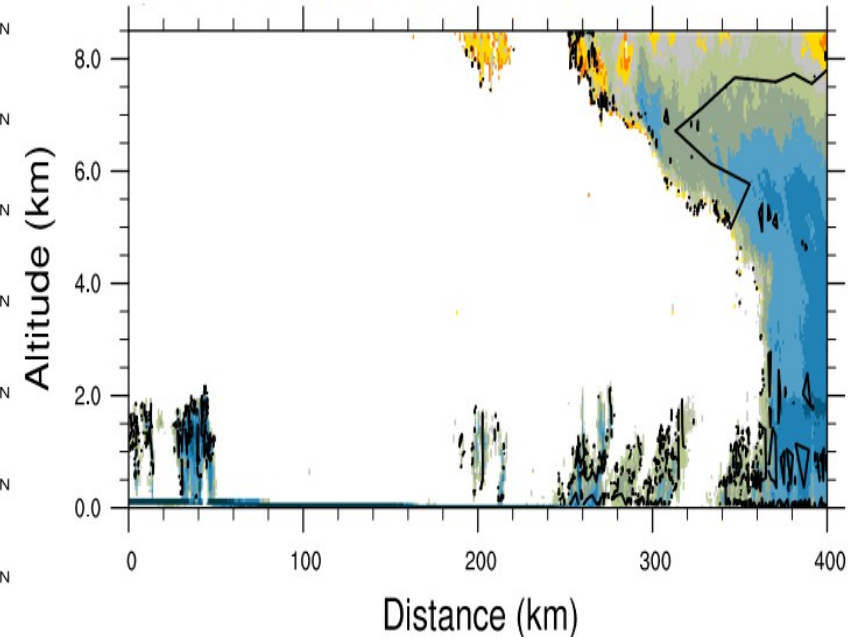
MSG @1500 UTC



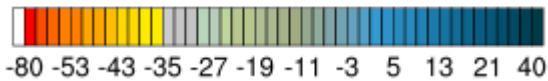
MSG @1500 UTC



RASTA Reflectivity (dBz) - Horizontal wind contours (m s⁻¹)

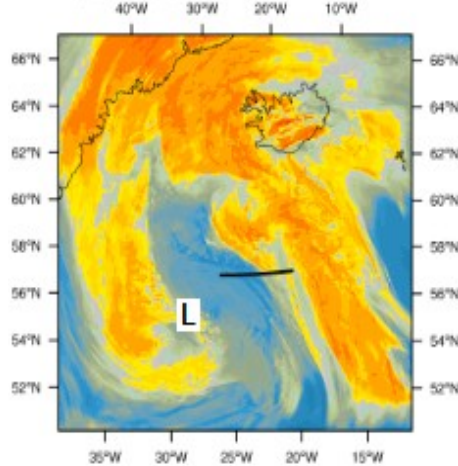


10.8 μ m Brightness Temperature (°C)



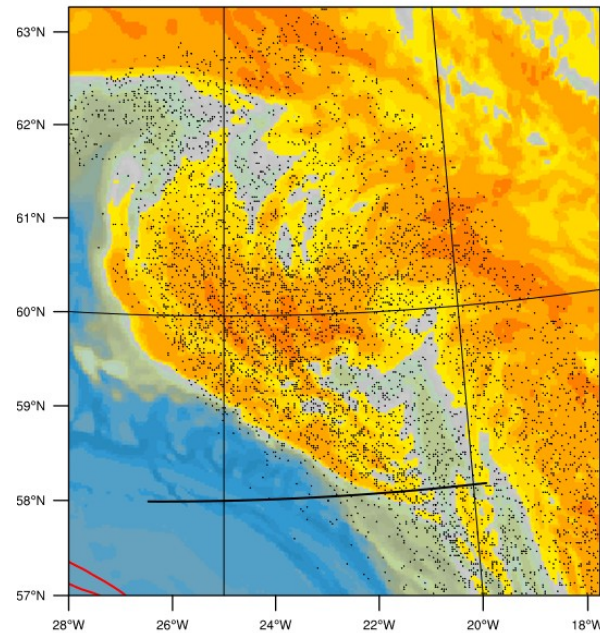
● WCB trajectories

Meso-NH @1500 UTC

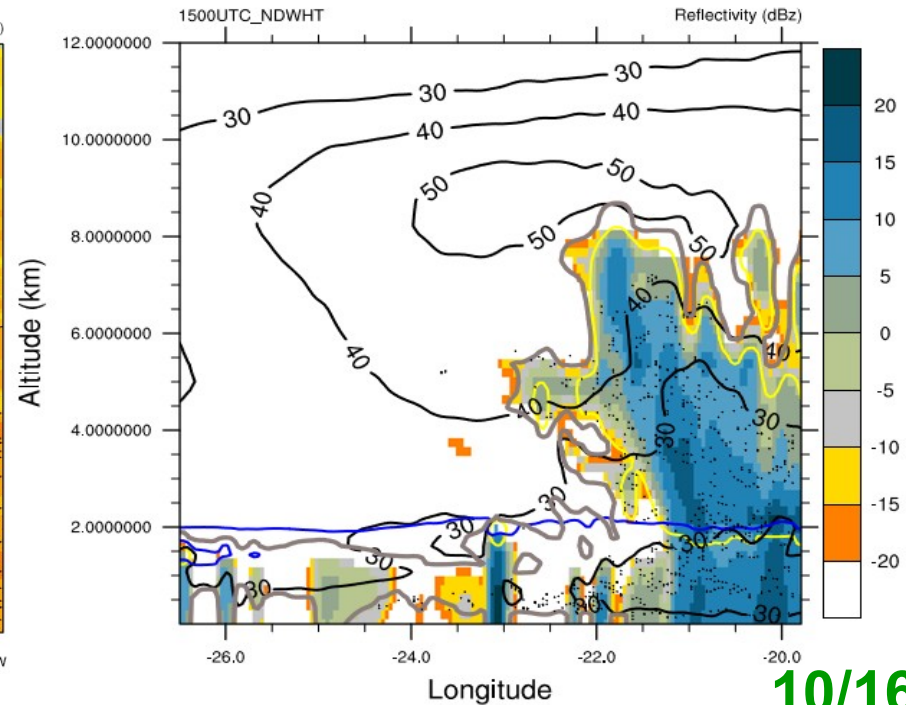


1500UTC_NDWHT

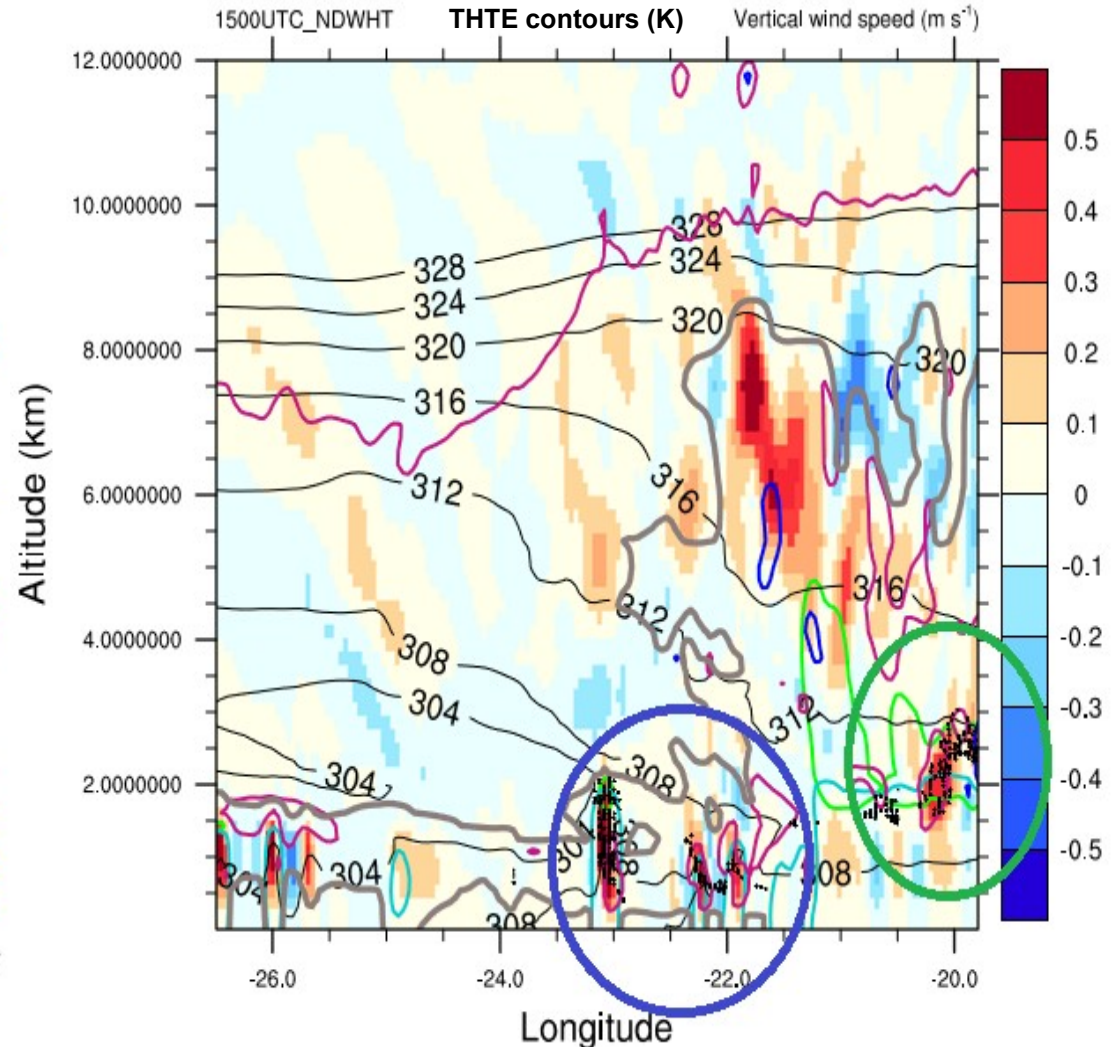
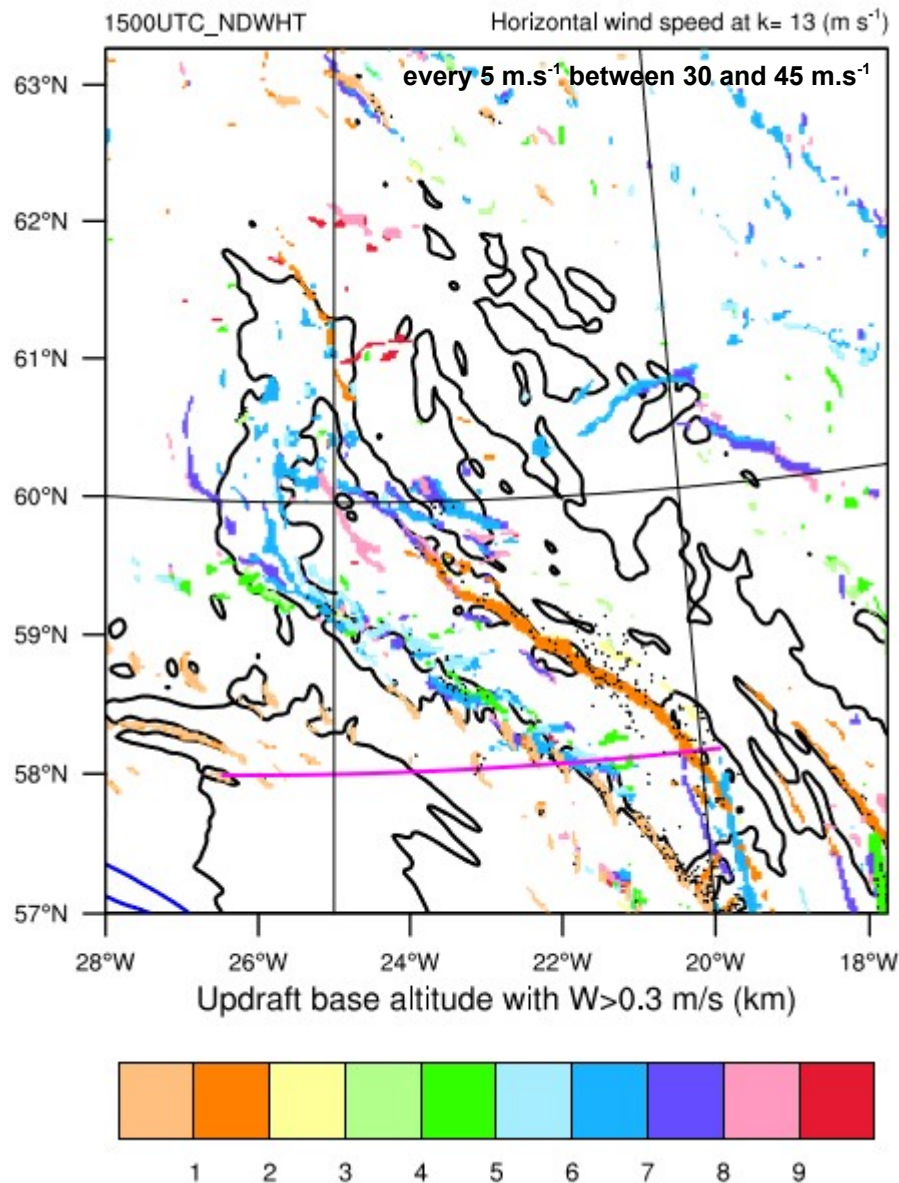
10.8 μ m BT (°C)



Horizontal wind contours (m.s⁻¹)



Fast ascents at 1500 UTC



• Fast Lag. trajs. at 1500 UTC
(every 20 trajectories)

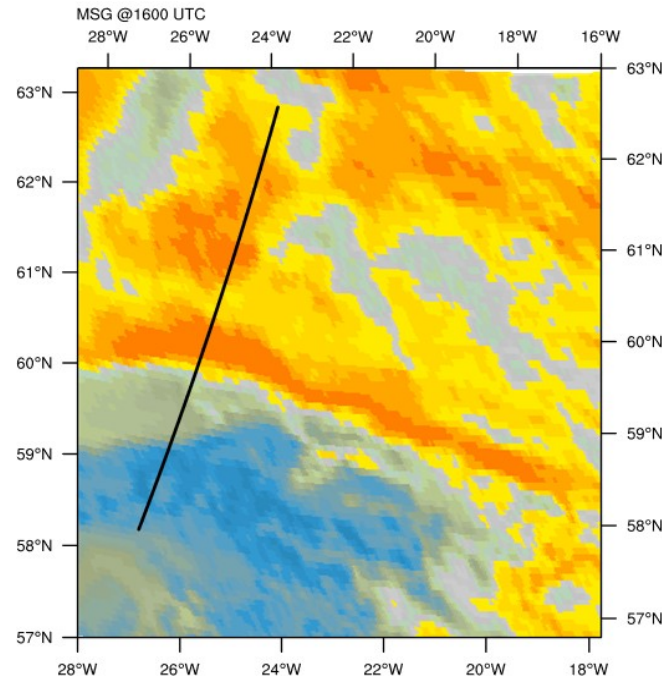
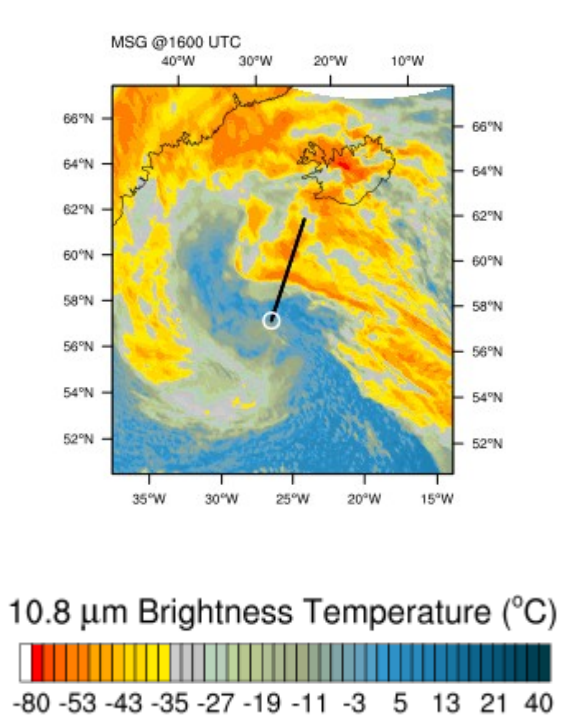
→ Good accordance between the location of Fast Lagrangian trajectories and the Updrafts base

→ **Extended low-level jet (LLJ)**

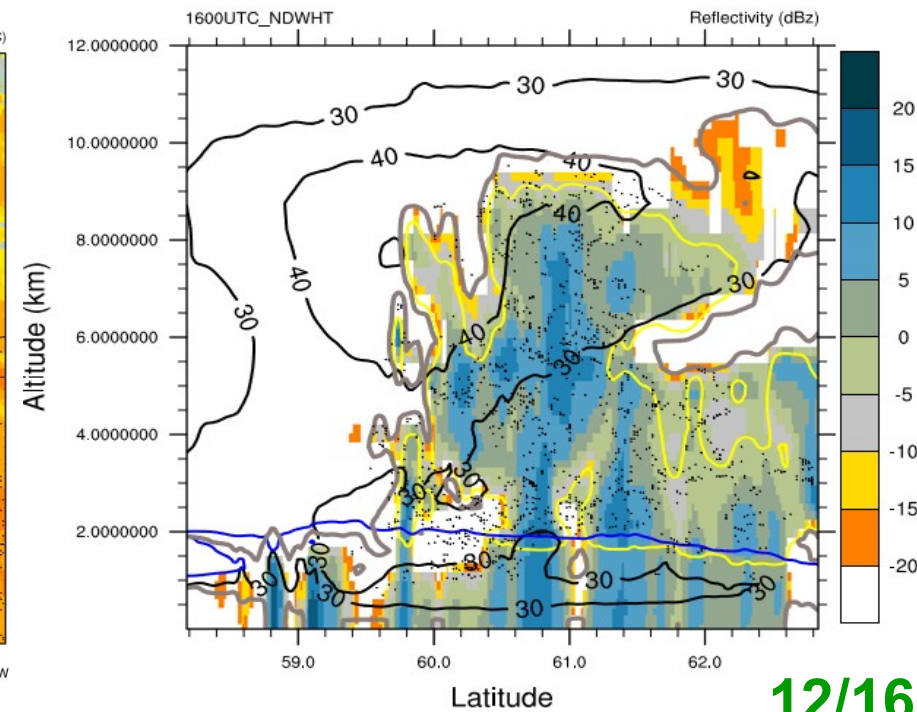
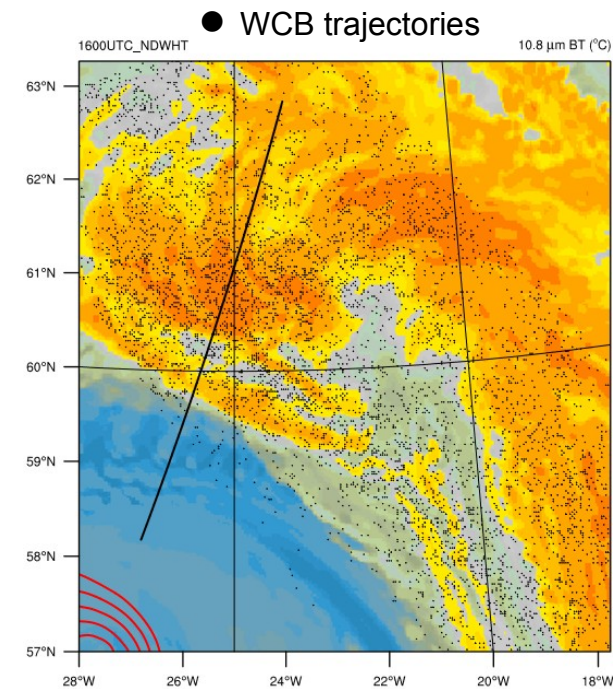
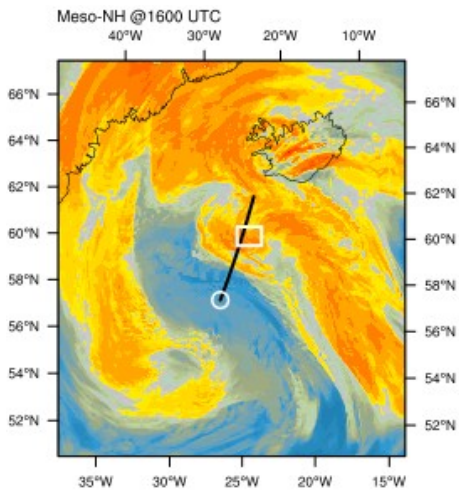
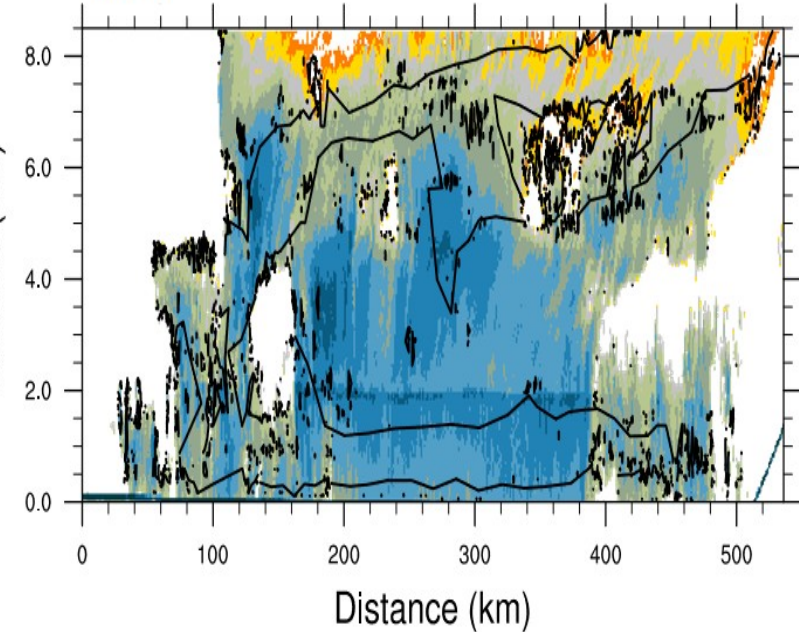
→ **Isolated shallow convection** located close to the cold front, at the west of the WCB and LLJ

→ **Organized low-level convection** located in the core of the WCB, at the east of LLJ

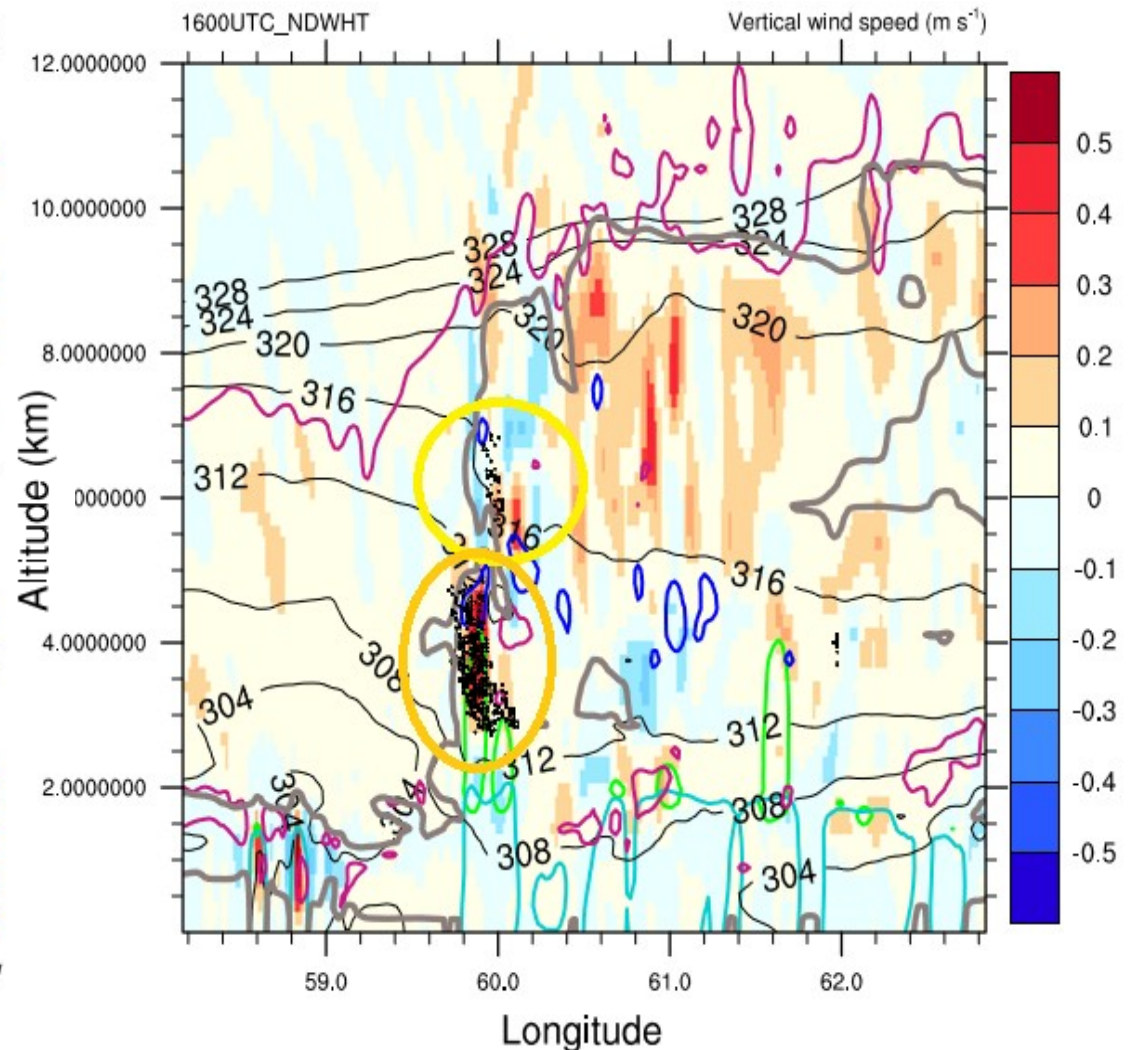
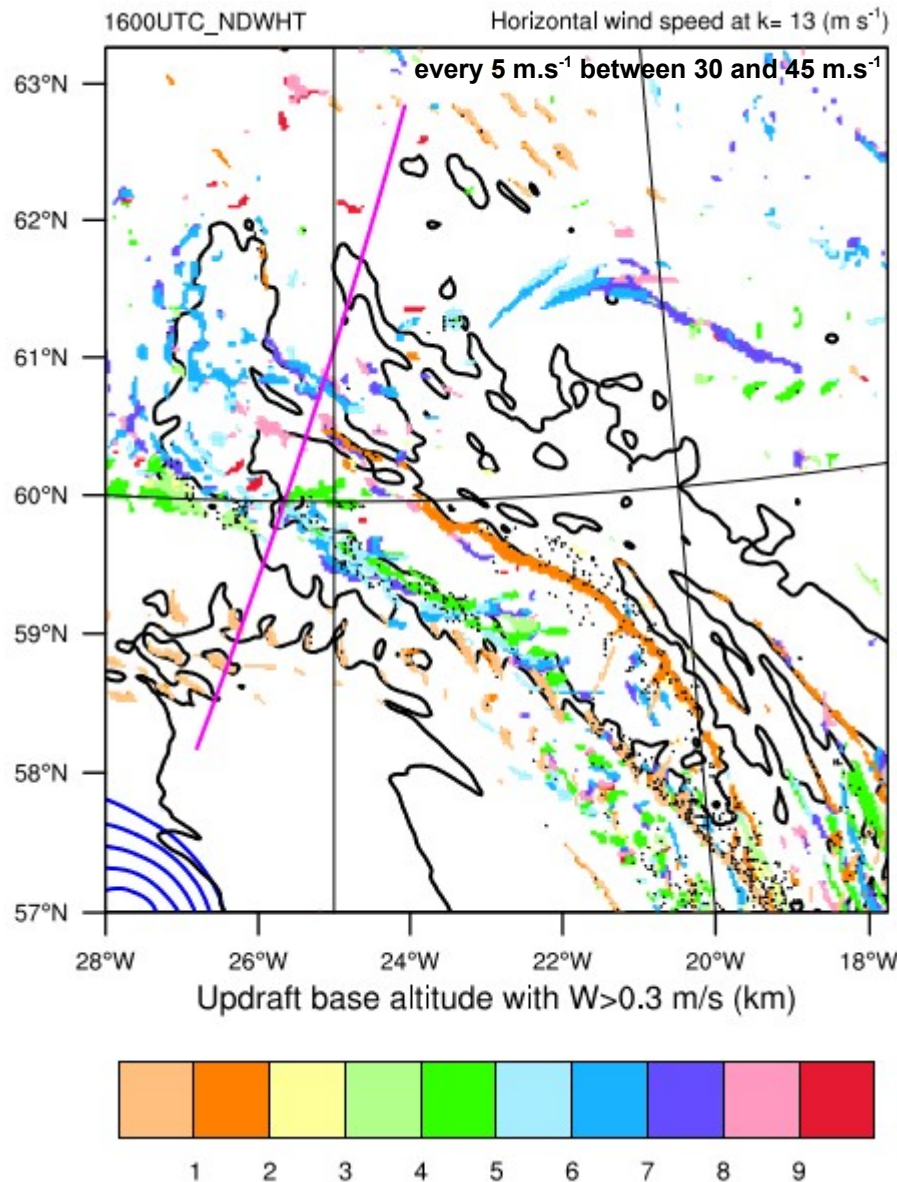
Zoom at 1600 UTC



RASTA Reflectivity (dBz) - Horizontal wind contours (m.s^{-1})



Fast ascents at 1600 UTC



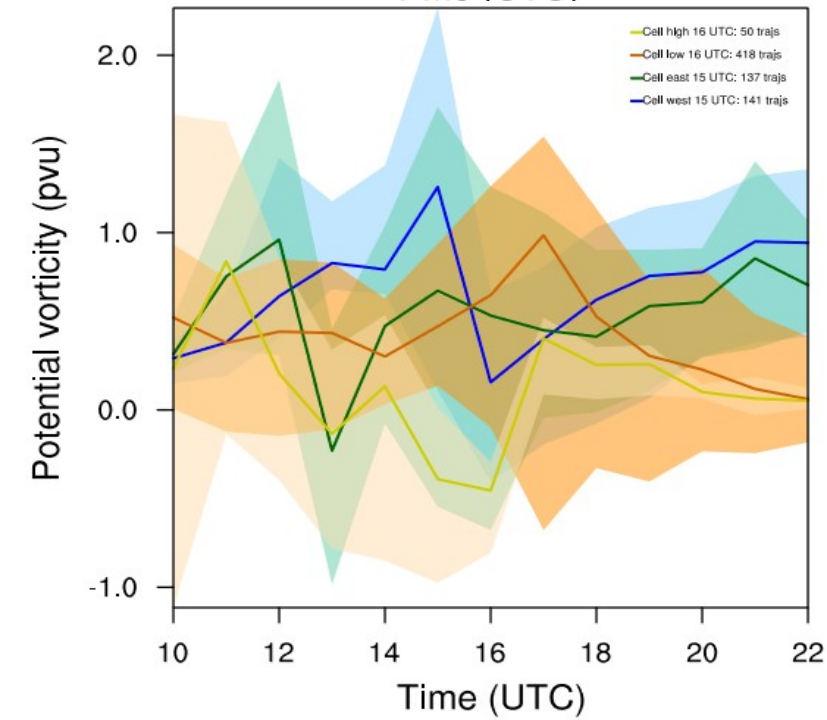
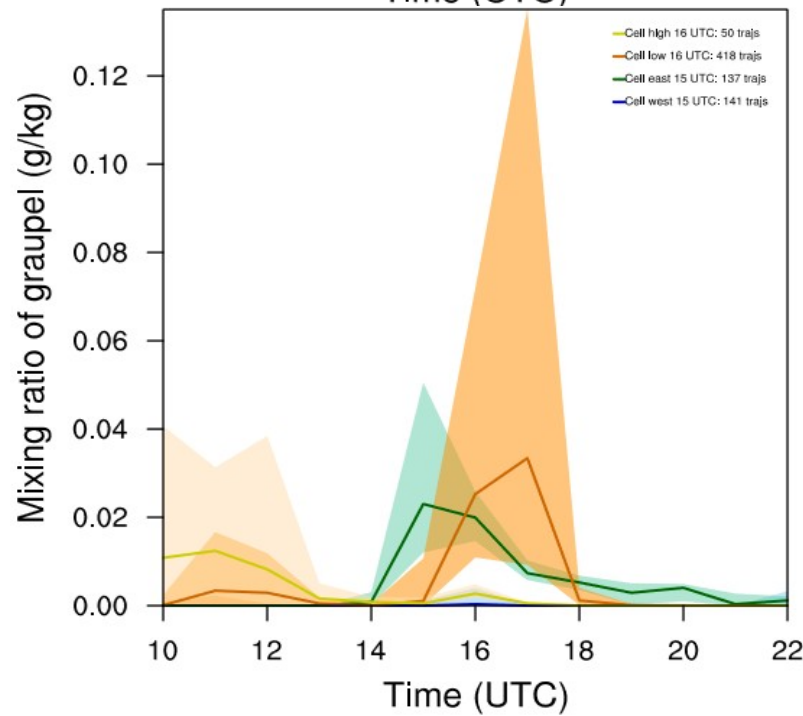
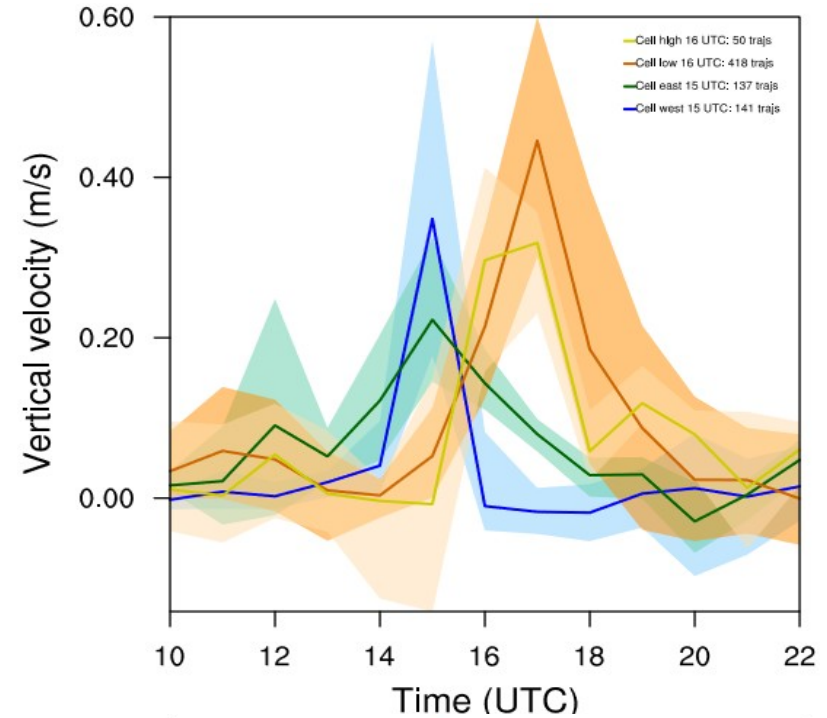
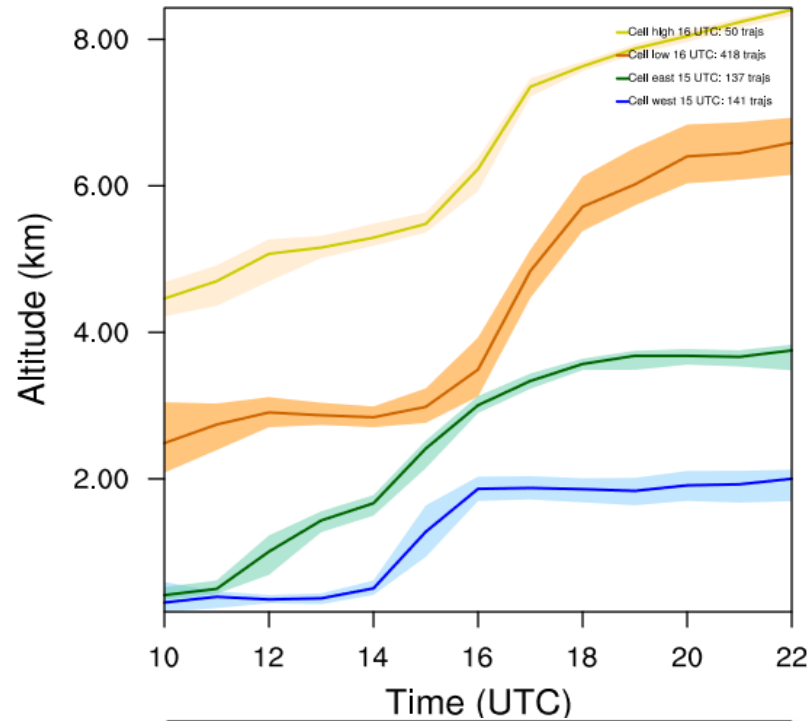
• Fast Lag. trajs. at 1500 UTC
(every 20 trajectories)

→ Good accordance between the location of Fast Lagrangian trajectories and the Updrafts base

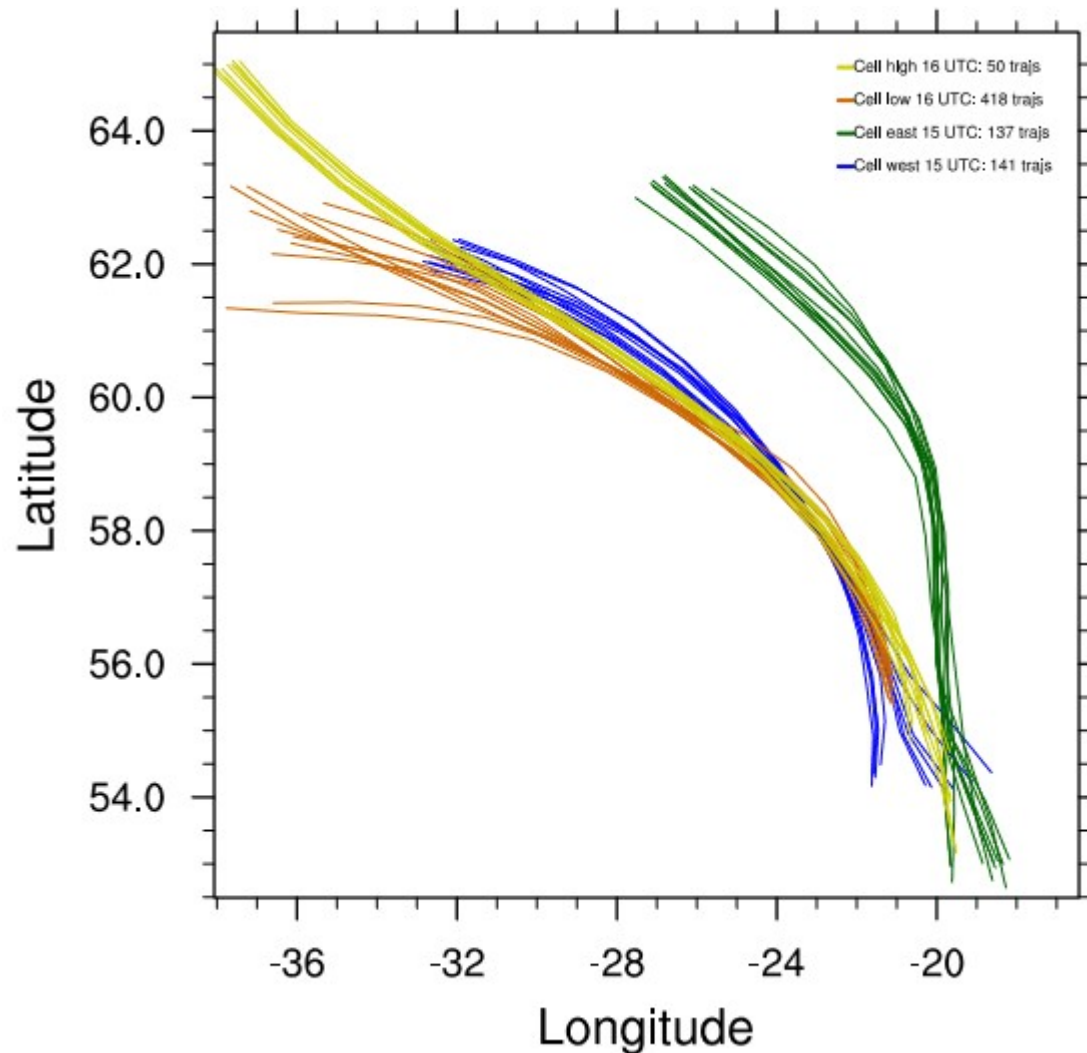
→ **Extended low-level jet (LLJ)**

→ **Isolated Mid-level (high/low) convection**
located to the west of the WCB and of the
LLJ

Temporal evolution of 4 identified convection cells



Trajectories of 4 identified convection cells (10-22 UTC)



→ Différence d'altitude de $z=2$ km pour les trajectoires rapides identifiées à 1600 UTC font qu'elles empruntent une tournure cyclonique ou anticyclonique en fin de trajectoire

Conclusions

1. What are the characteristics of the ascents in the WCB of the Stalactite cyclone?

- **1/3 of fast ascents** ($-100 \text{ hPa} \geq \Delta P_{2h}$) embedded in the slantwise WCB ascents
- **2/3 of cyclonic trajectories, 3/5 for slow ascents and 4/5 for fast ascents**
- **anticyclonic ascents located at higher altitude than cyclonic ascents**
- **slow ascents in the core of WCB and fast ascents in its west side, close to the cold front**

2. What are the processes leading to fast ascents?

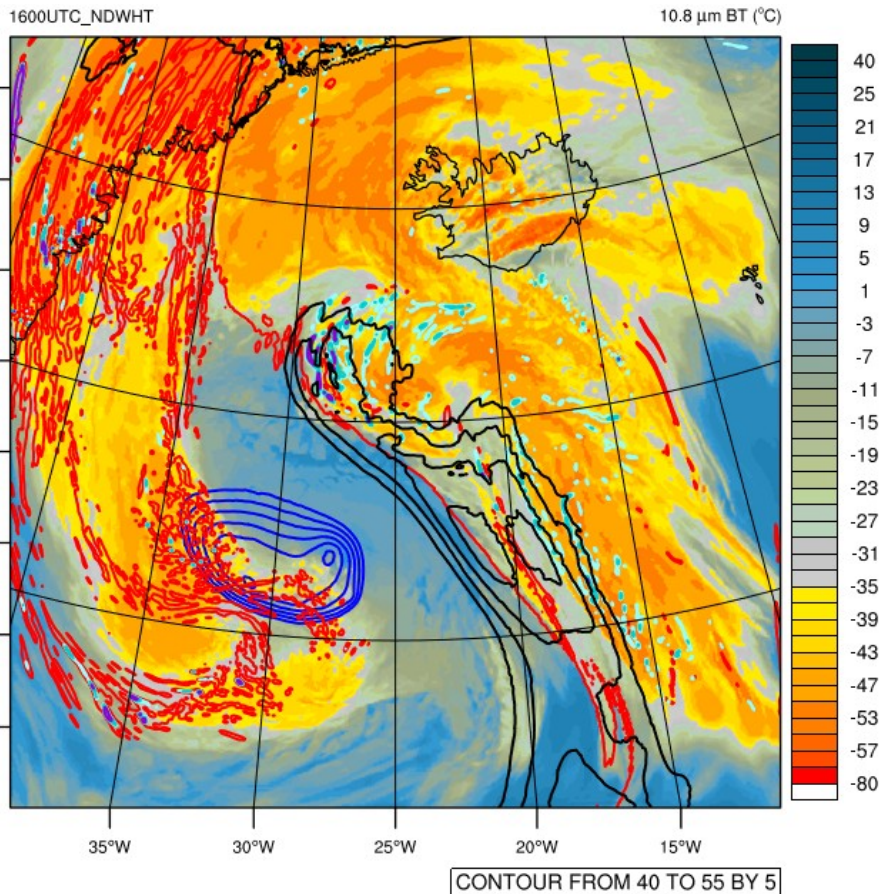
- **Isolated shallow convection** : west of the WCB and LLJ, low-layer ($500 \text{ m} < z < 2 \text{ km}$), cyclonic trajectory, $PV_{\text{mediane}} > 0 \text{ PVU}$
- **Organized low-level convection** : WCB core at $500 \text{ m} < z < 4 \text{ km}$, east of the LLJ, graupel, cyclonic trajectory, $PV_{\text{mediane}} > 0 \text{ PVU}$ mainly
- **Isolated Mid-level (*high / low*) convection** : WCB at $2 \text{ km} < z < 8 \text{ km}$, graupel, *anticyclonic/cyclonic trajectories*, $PV_{\text{mediane}} \sim 0 \text{ PVU}$ with **negative values**

Perspectives : negative PV bands in the cloud head !

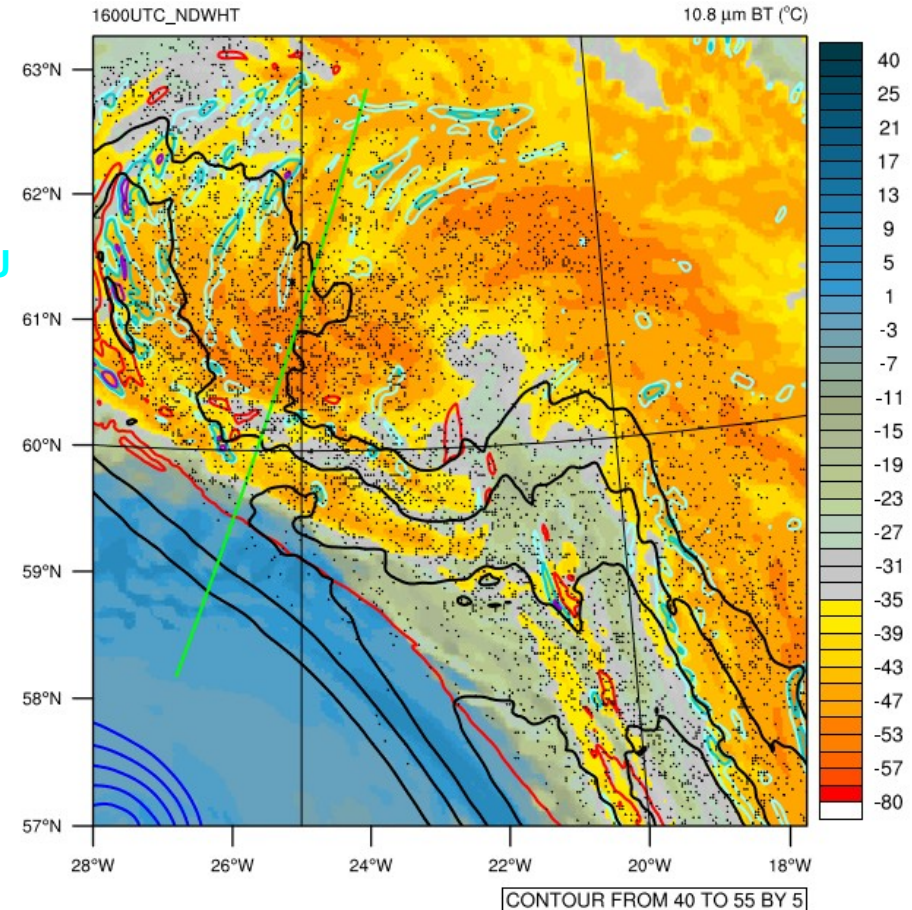
1600 UTC

At $\theta=320$ K

1600 UTC



+2 PVU
-0.5 / -1 PVU
-2 PVU



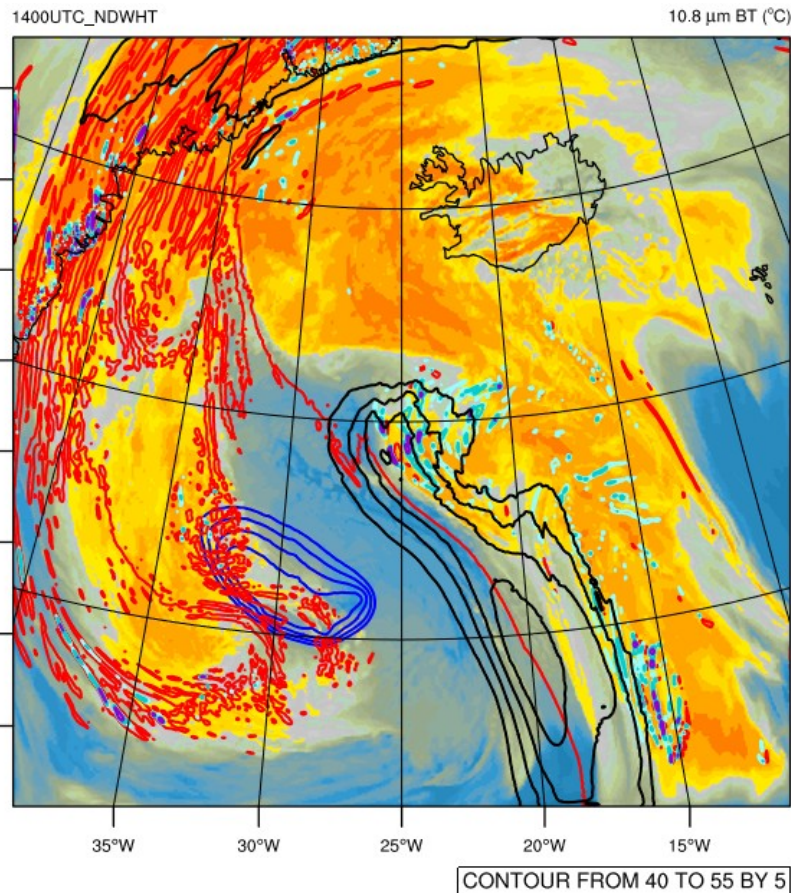
Horizontal wind speed contours
every 5 m.s⁻¹ between 40 and 55 m.s⁻¹

→ Negative PV bands follow
the *cloud head* structure as well
as the Jet Stream

→ Jet Stream is locally
intensified by the PV
gradient

Origin of the negative PV bands?

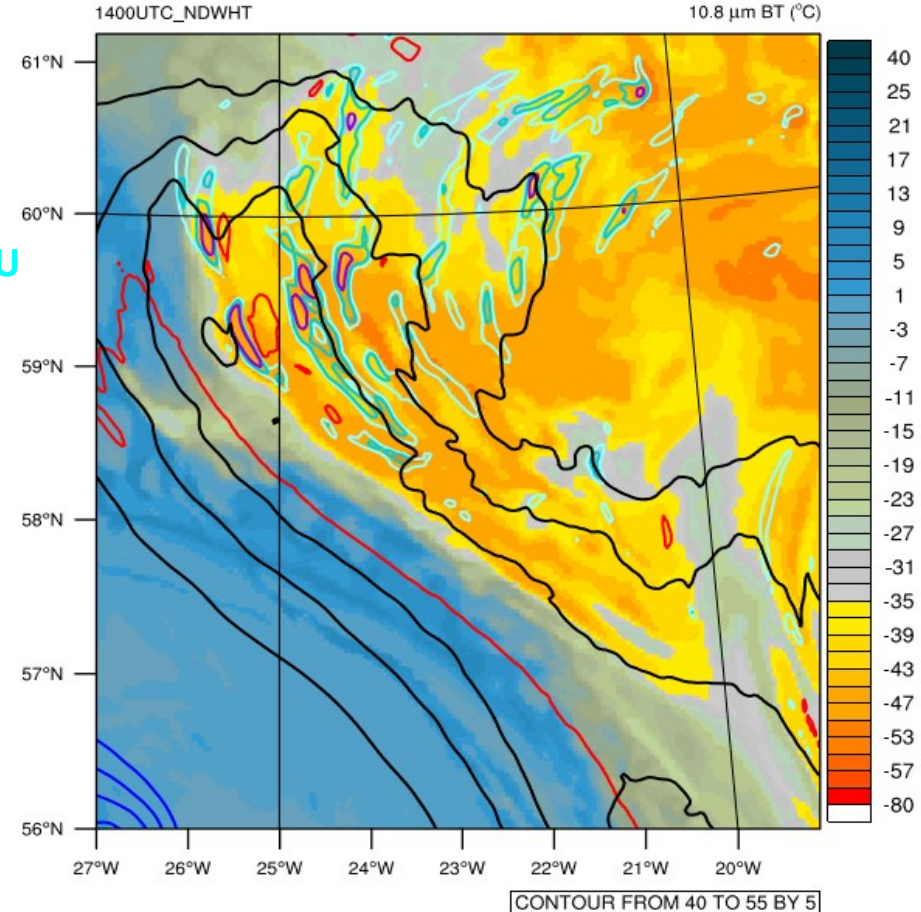
1400 UTC



At $\theta=320$ K

+2 PVU
-0.5 / -1 PVU
-2 PVU

1400 UTC



Horizontal wind speed contours
every 5 m.s^{-1} between 40 and 55 m.s^{-1}

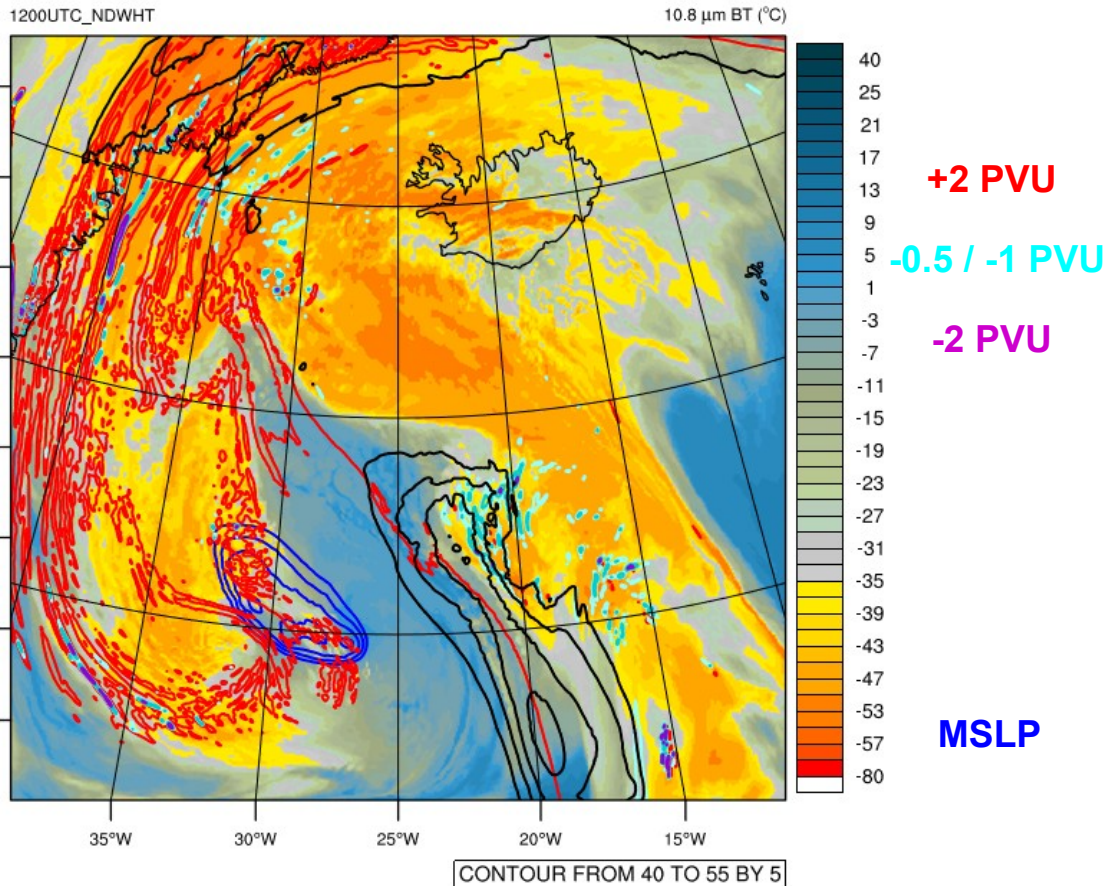
→ Negative PV bands follow
the *cloud head* structure as well
as the Jet Stream

→ Jet Stream is locally
intensified by the PV
gradient

Origin of the negative PV bands?

1200 UTC

At $\theta=320$ K



→ Negative PV bands persist with time...

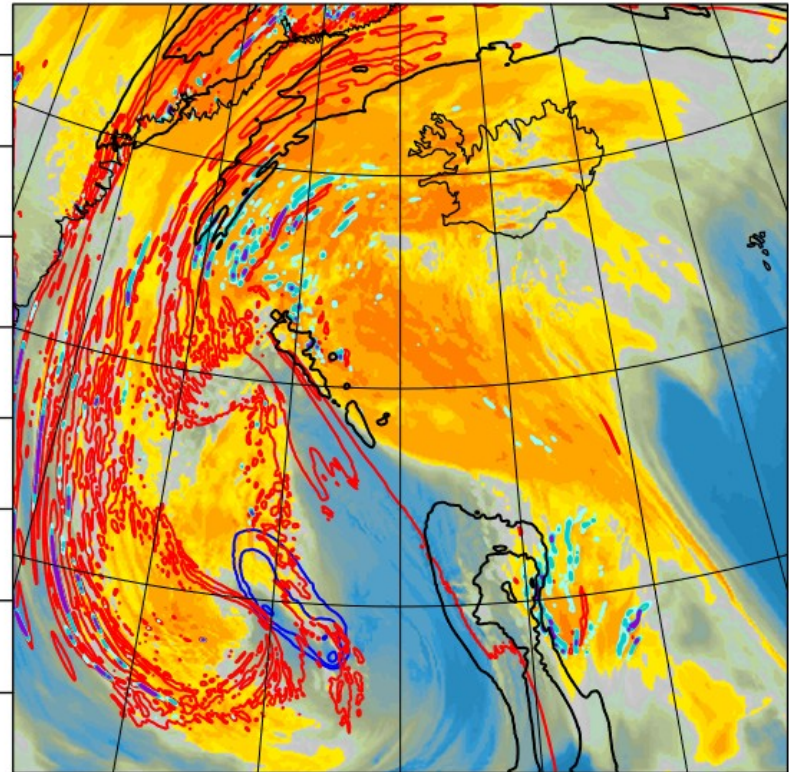
Horizontal wind speed contours
every 5 m.s⁻¹ between 40 and 55 m.s⁻¹

→ Negative PV bands follow
the *cloud head* structure as well
as the Jet Stream

Origin of the negative PV bands?

1000 UTC

1000UTC_NDWHT 10.8 μm BT ($^{\circ}\text{C}$)



CONTOUR FROM 40 TO 55 BY 5

At $\theta=320$ K

+2 PVU

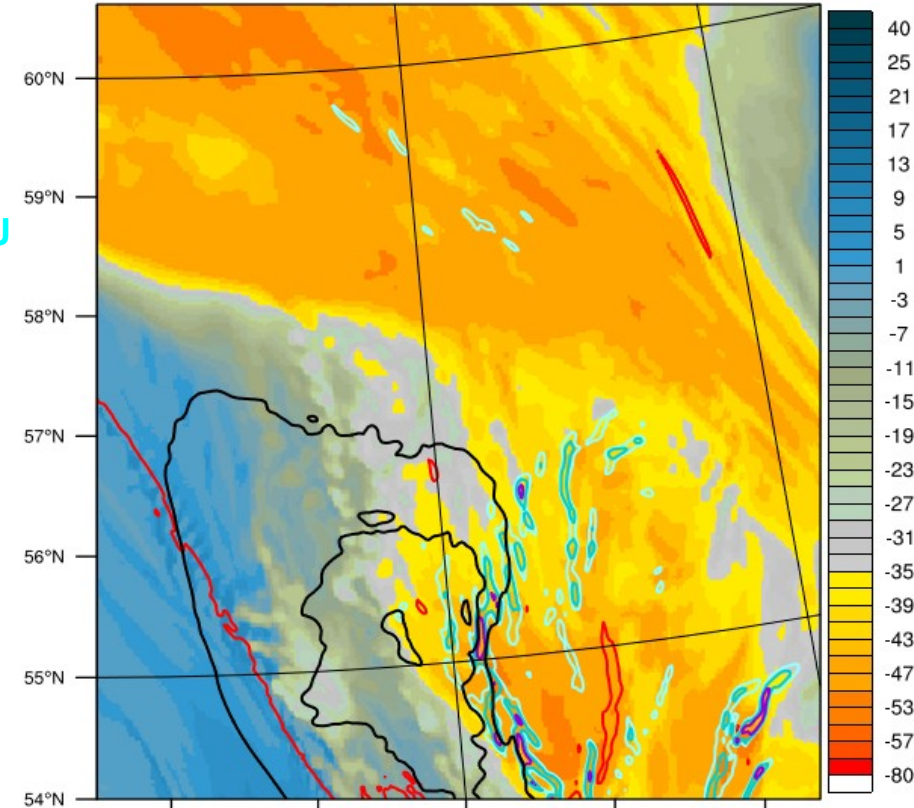
-0.5 / -1 PVU

-2 PVU

MSLP

1000 UTC

1000UTC_NDWHT 10.8 μm BT ($^{\circ}\text{C}$)



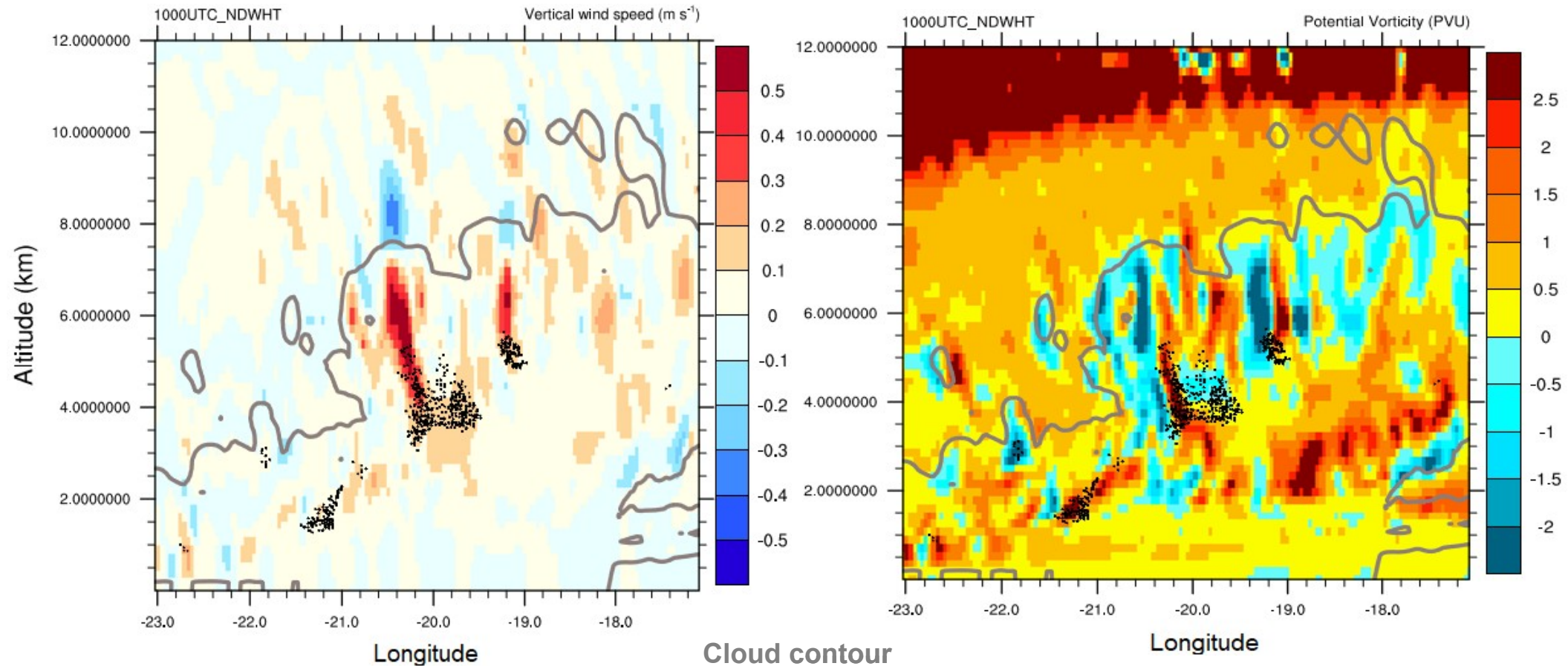
CONTOUR FROM 40 TO 55 BY 5

Horizontal wind speed contours
every 5 m.s^{-1} between 40 and 55 m.s^{-1}

→ Negative PV bands follow
the *cloud head* structure as well
as the Jet Stream

→ Negative PV bands
formation

Vertical cross section at 1000 UTC

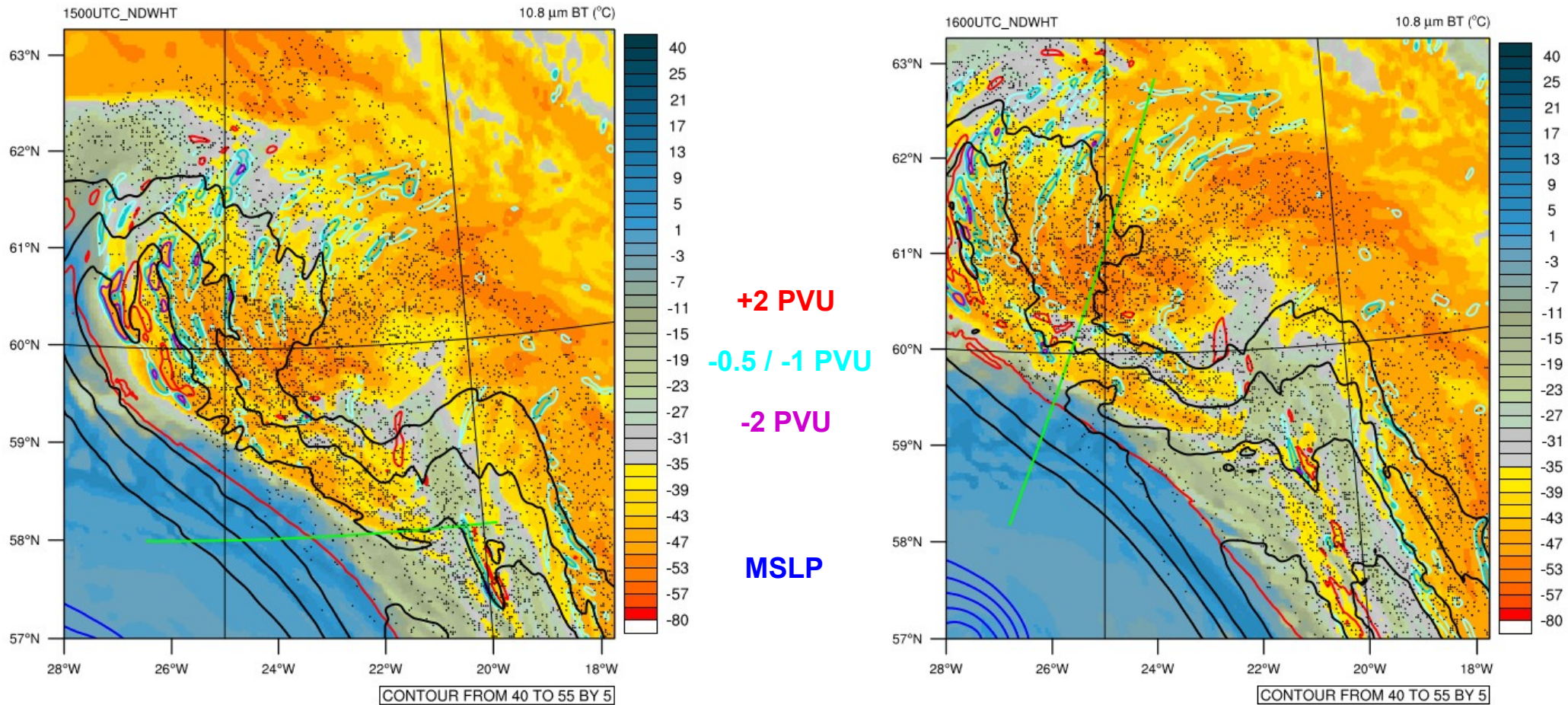


→ Fast Lagrangian trajectories in mid-level convective regions

→ **Negative PV** in altitude located in these regions

ANNEXES

Jet Stream and PV at θ_{320K}



Horizontal wind speed contours
every 5 m.s^{-1} between 40 and 55 m.s^{-1}

→ Negative PV bands
follow the *cloud head*
structure as well as the
Jet Stream

→ **What is the origin of
the negative PV bands?**

→ Jet Stream is locally
intensified by the PV
gradient

Origin of the negative PV bands?

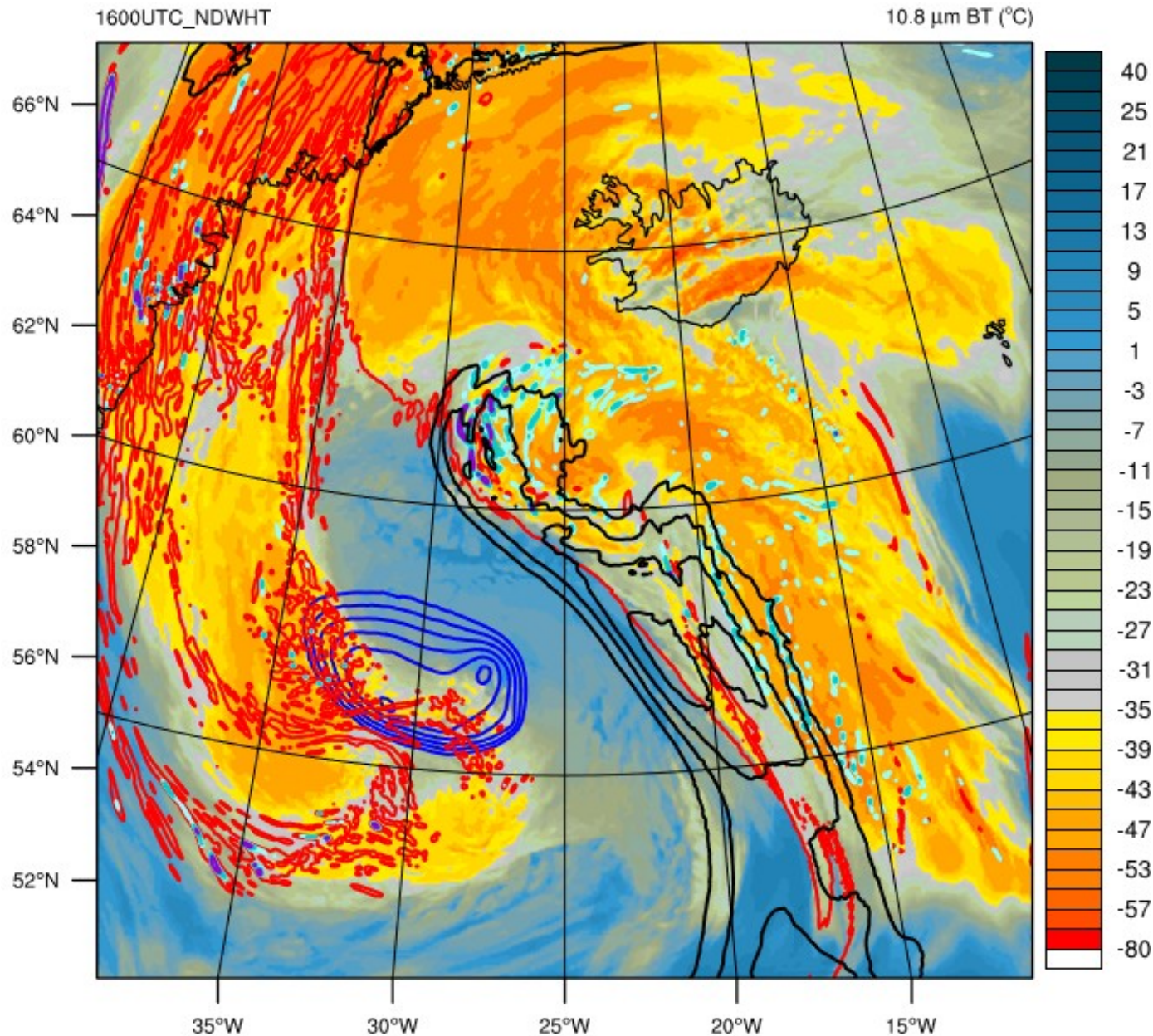
1600 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

Origin of the negative PV bands?

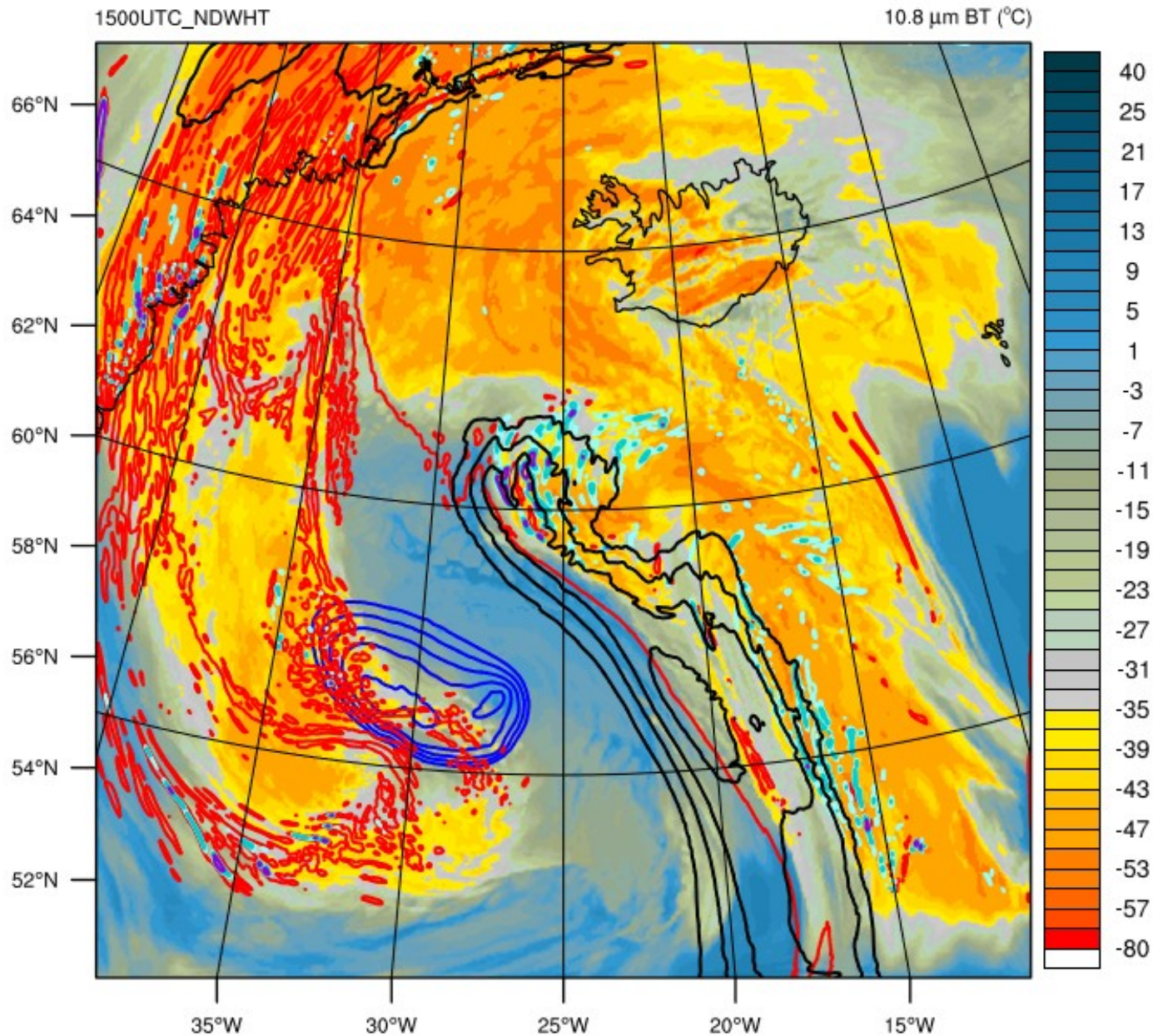
1500 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



Negative PV bands
follow the *cloud
head* structure
observed at 1600

CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

Origin of the negative PV bands?

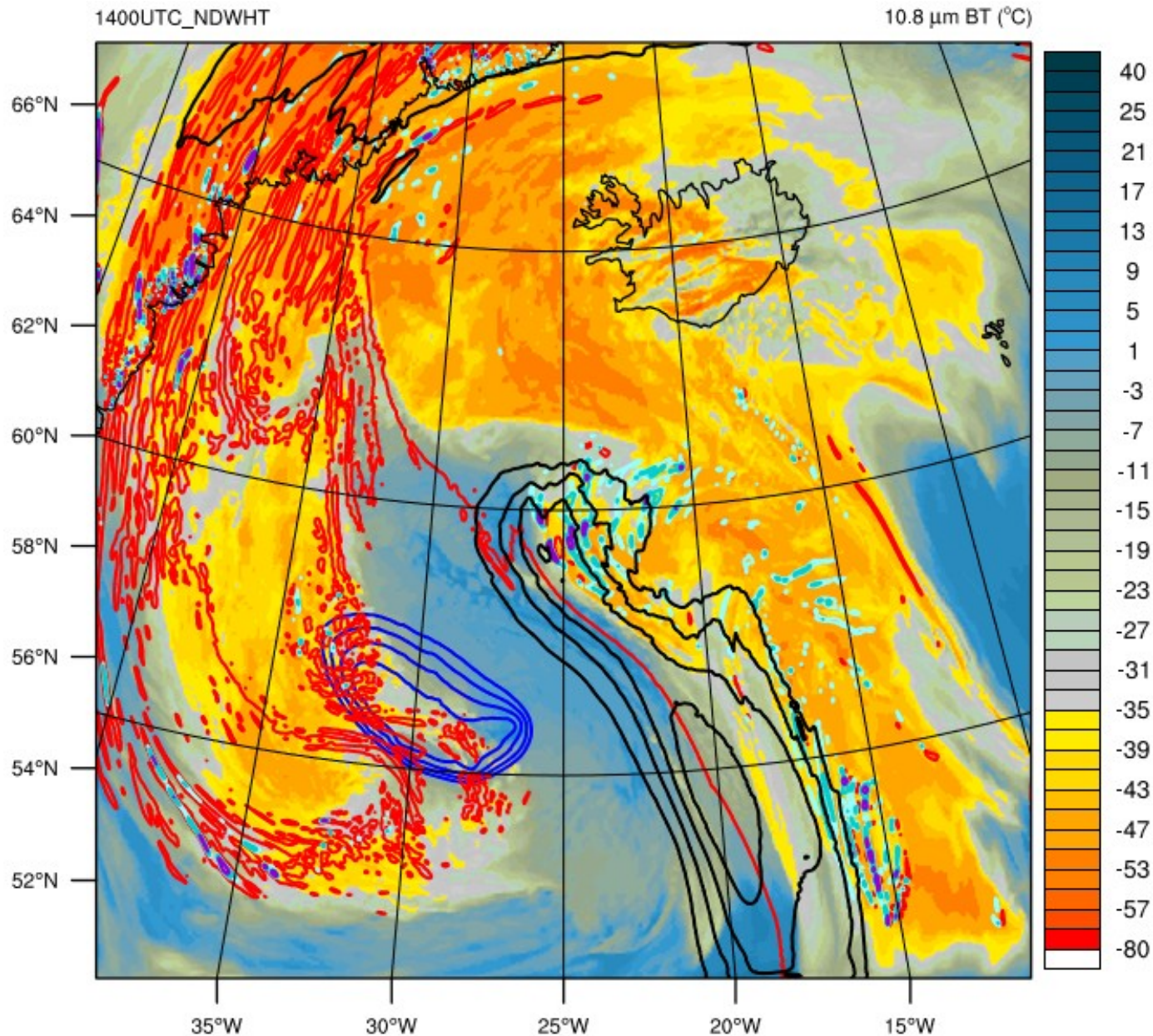
1400 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



at $\theta=320$ K

Origin of the negative PV bands?

1400 UTC

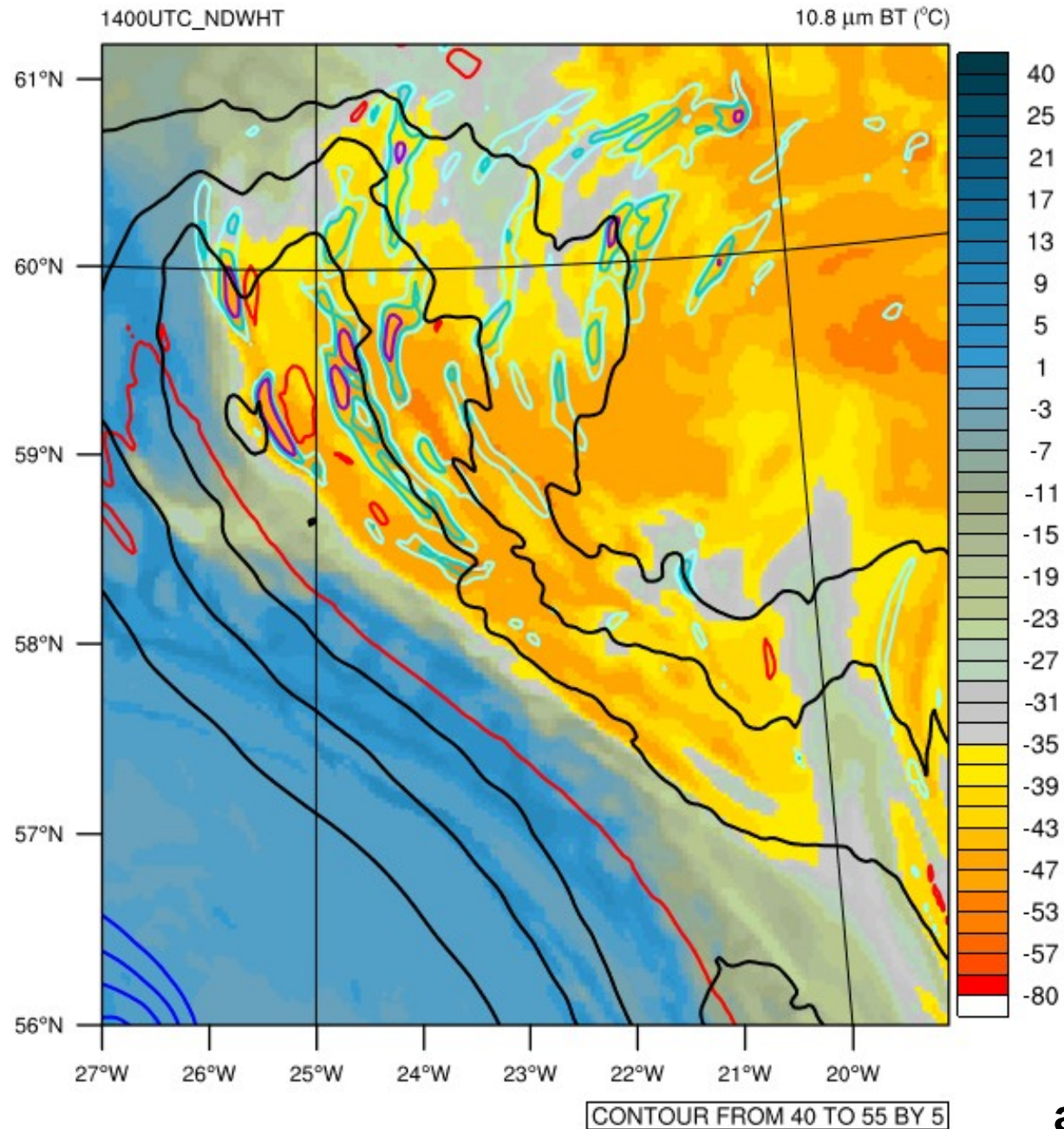
Zoom

+2 PVU

-0.5 / -1 PVU

-2 PVU

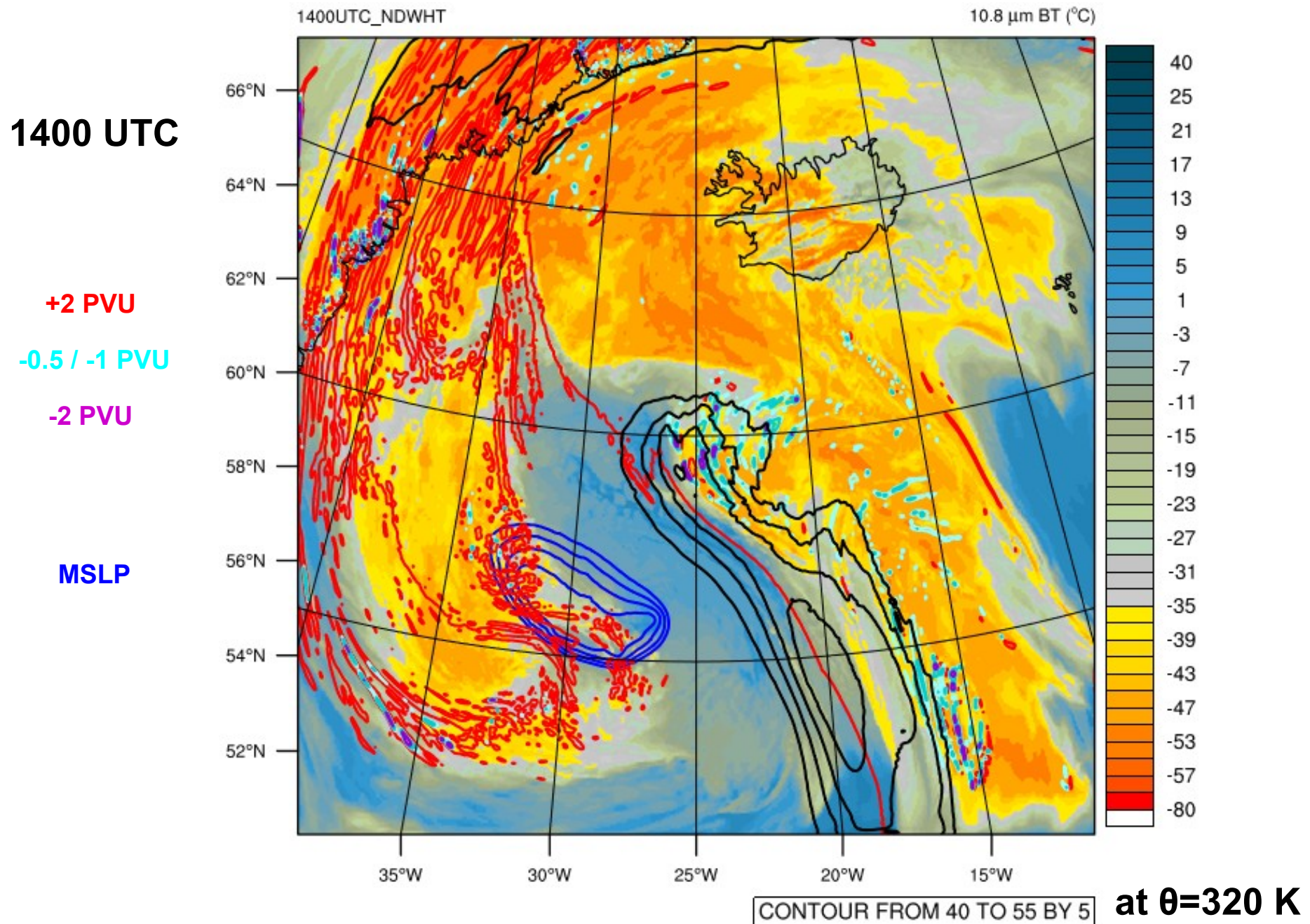
MSLP



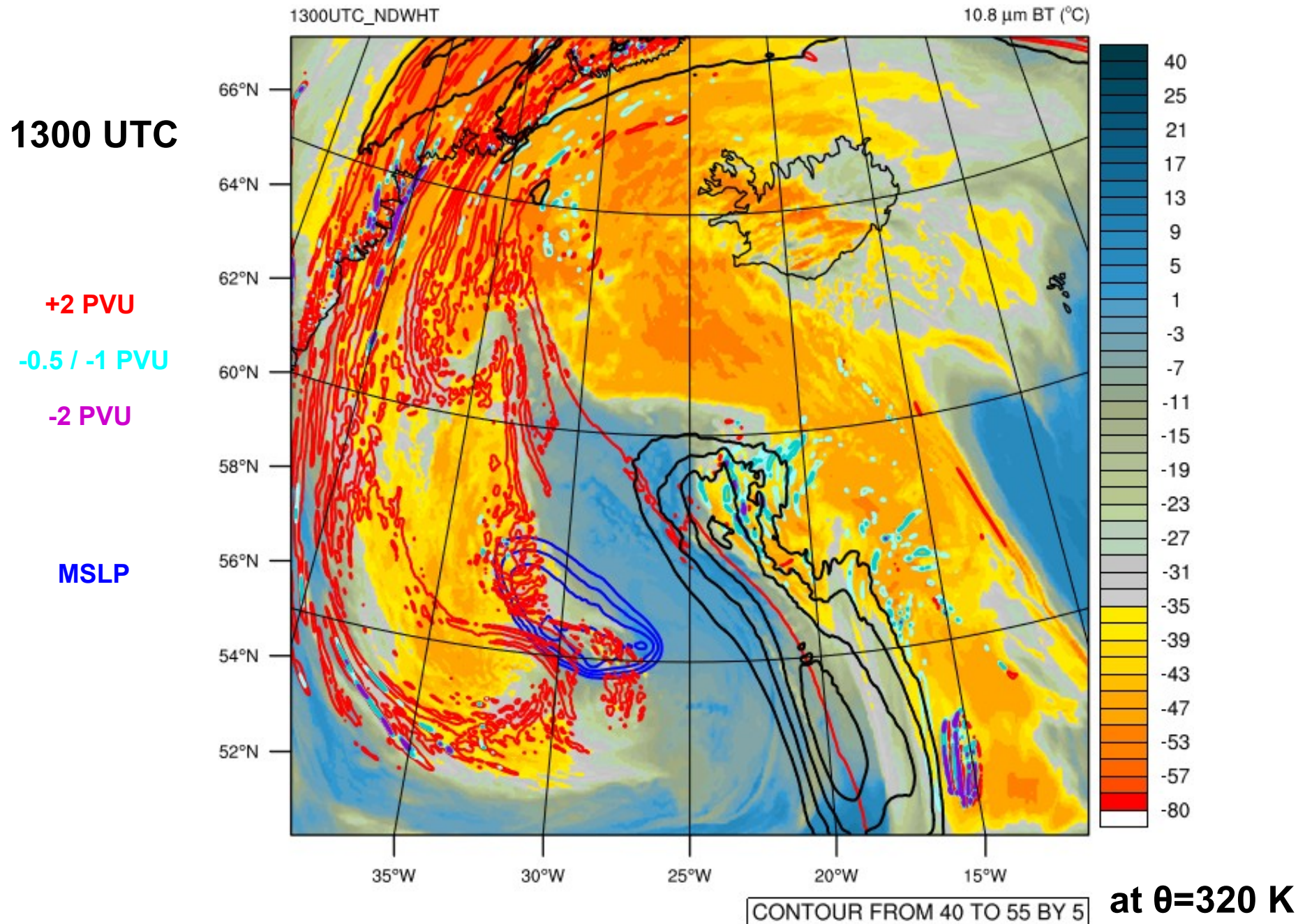
Jet Stream is locally intensified by the PV gradient

at $\theta=320$ K

Origin of the negative PV bands?



Origin of the negative PV bands?



Origin of the negative PV bands?

1200 UTC

+2 PVU

-0.5 / -1 PVU

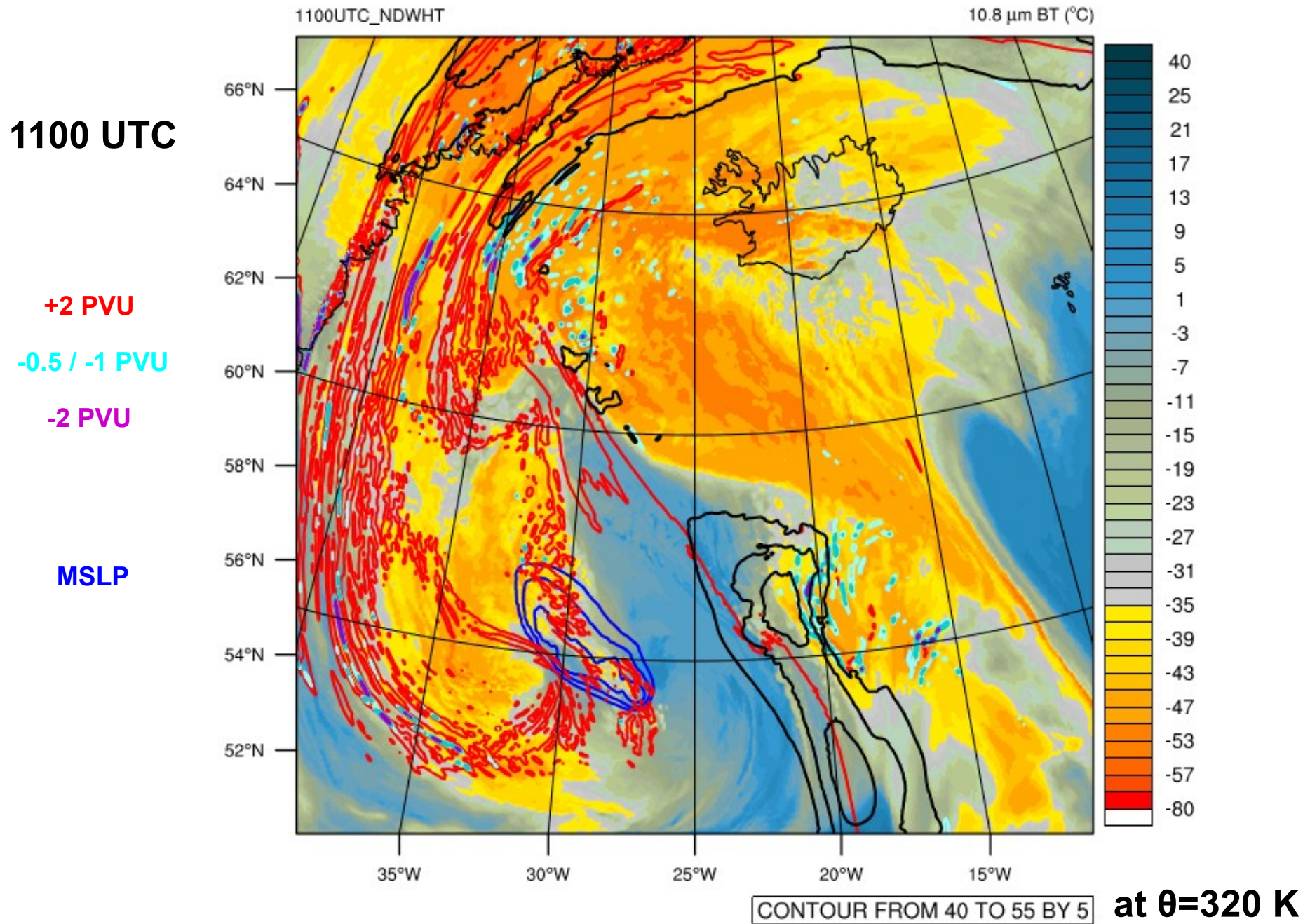
-2 PVU

MSLP

Negative PV
bands persist
with time

at $\theta=320$ K

Origin of the negative PV bands?



Origin of the negative PV bands?

1000 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP

Negative PV
bands formation

at $\theta=320$ K

Origin of the negative PV bands?

1000 UTC

Zoom

+2 PVU

-0.5 / -1 PVU

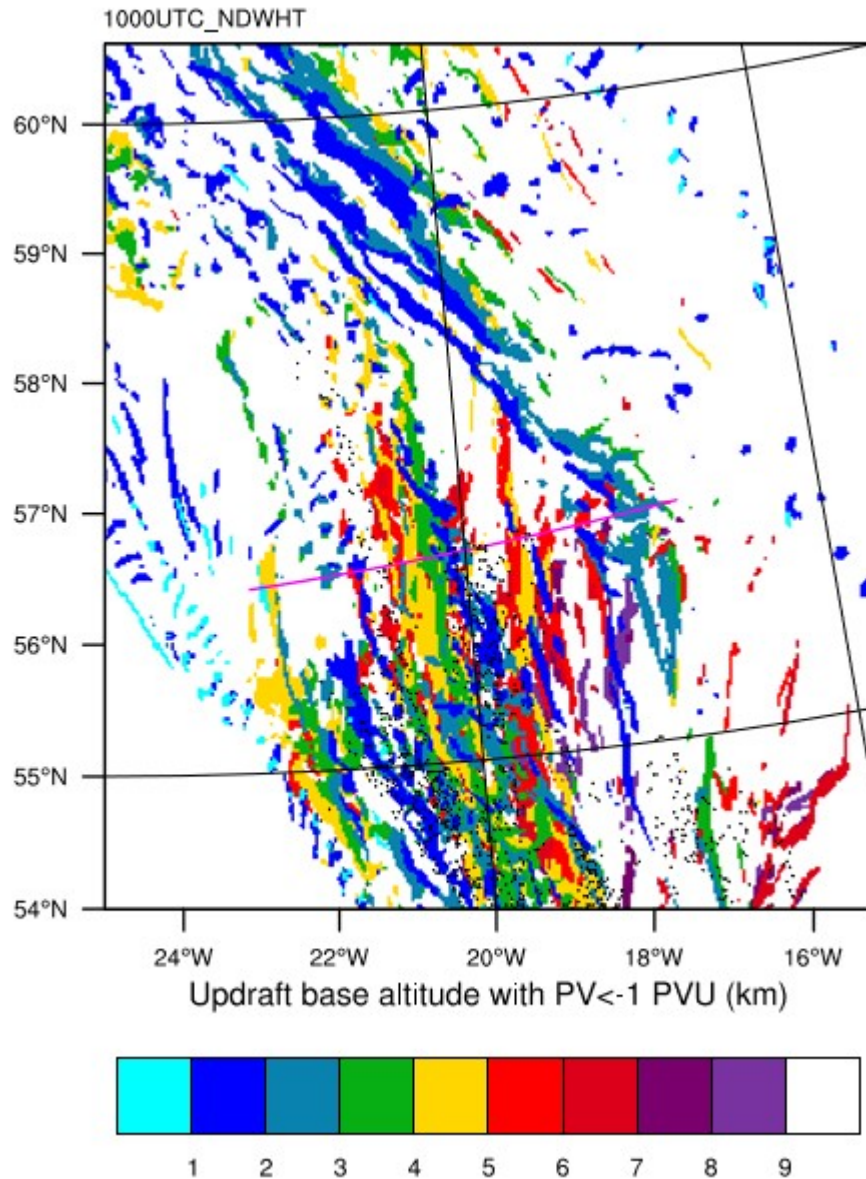
-2 PVU

MSLP

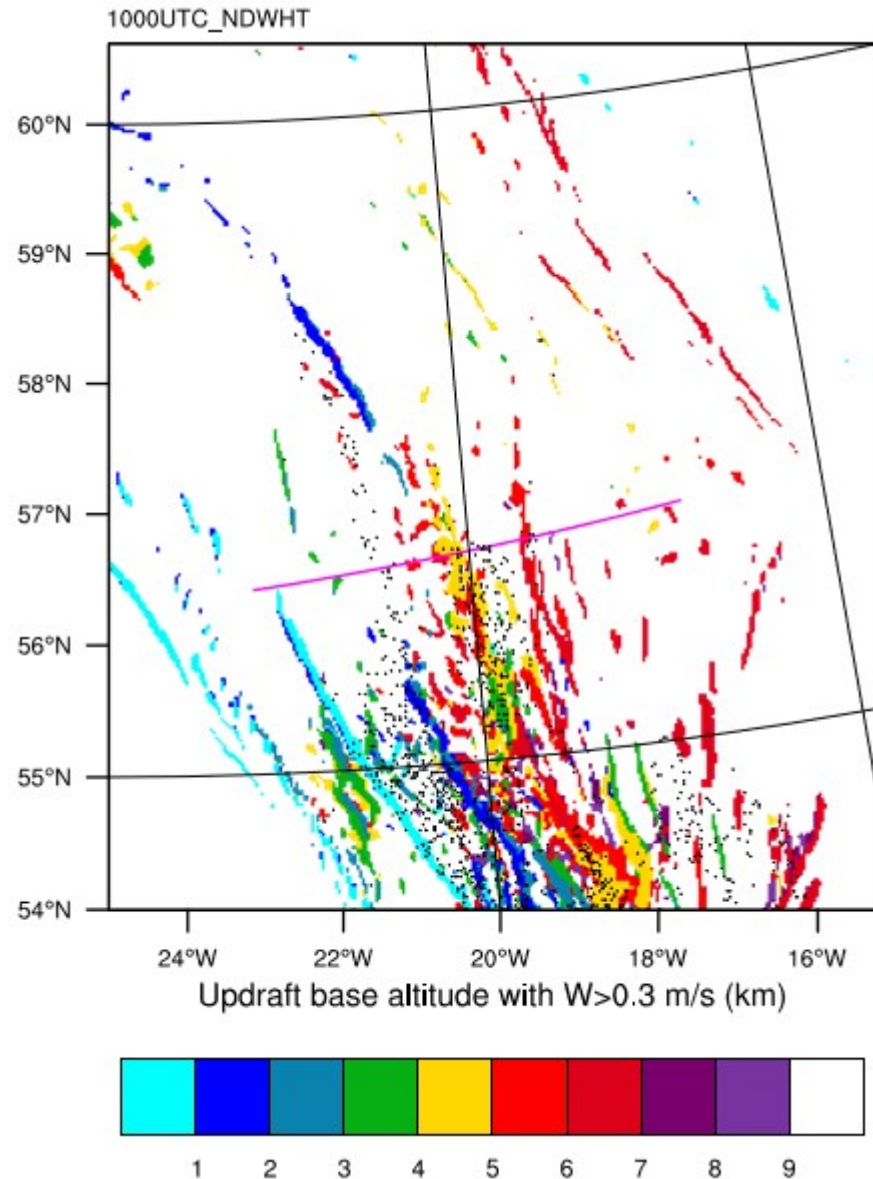
Negative PV
bands formation

at $\theta=320$ K

Negative PV bands at 1000 UTC

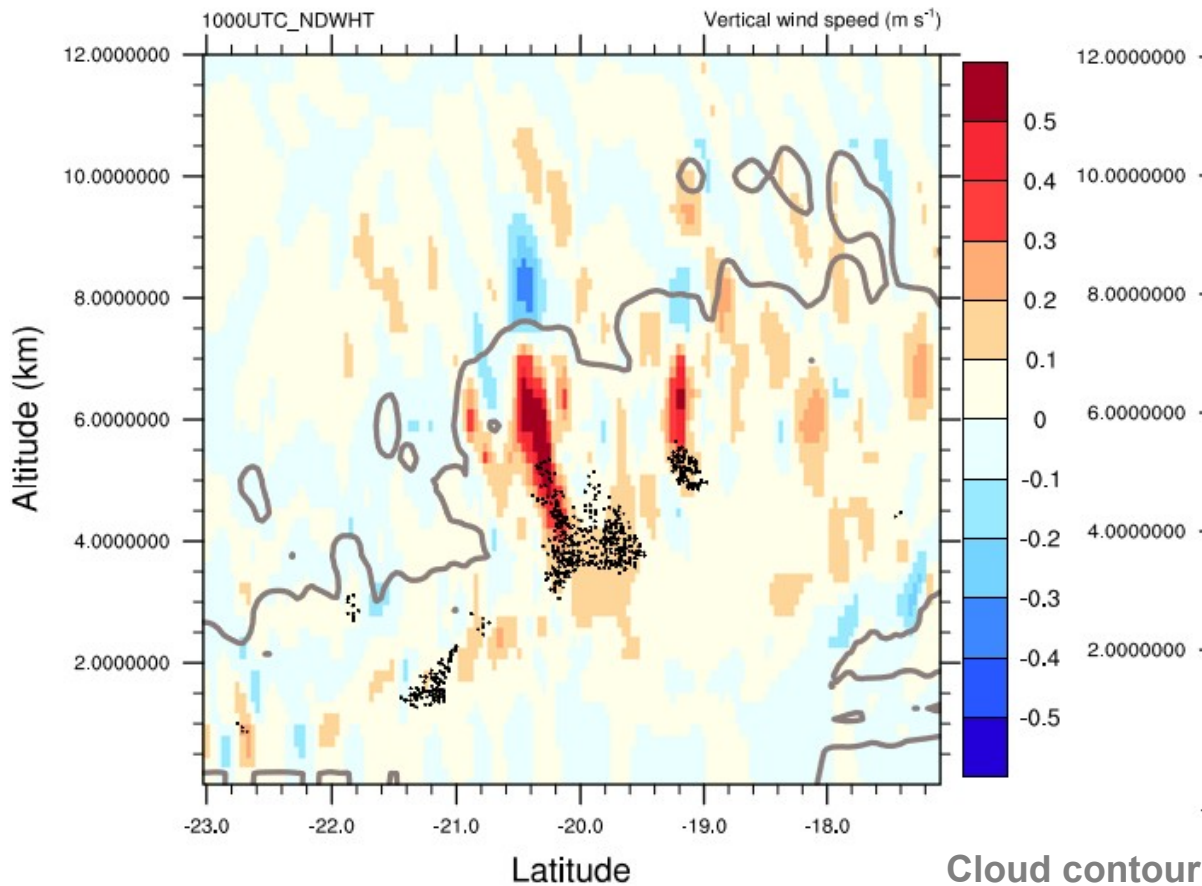


Altitude and location of negative PV bands correspond to those of Updrafts $z \sim 5$ km



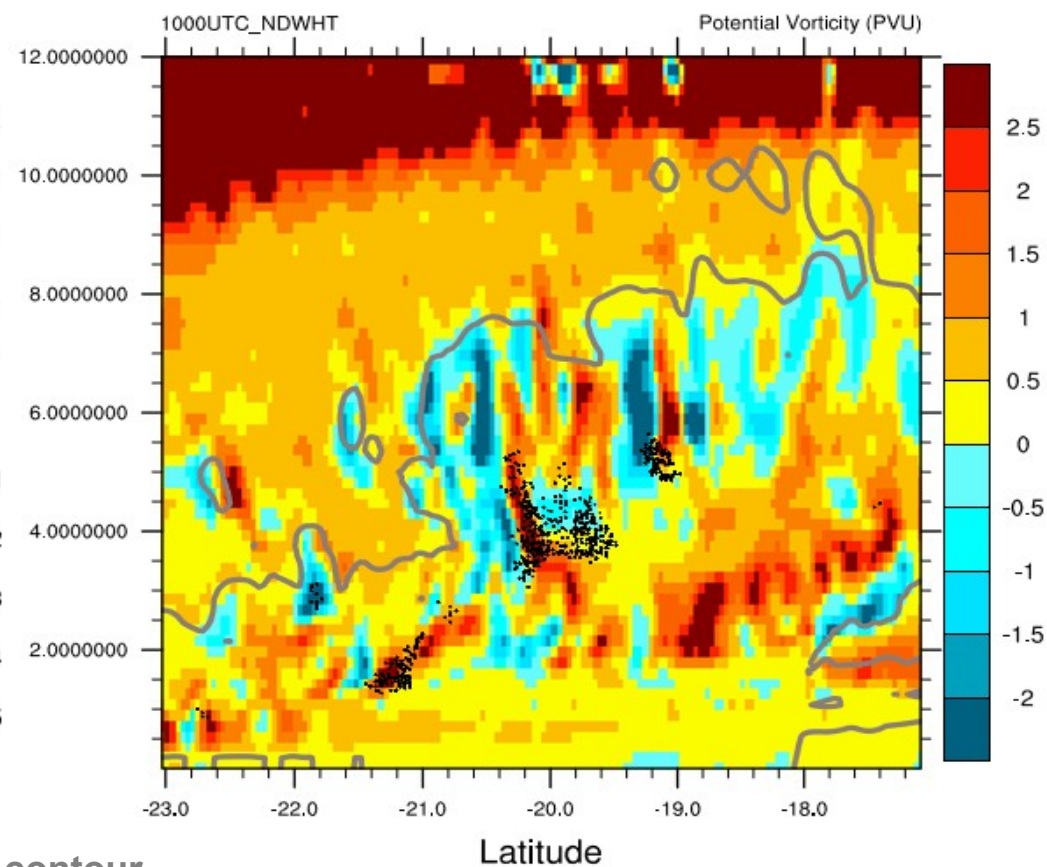
Fast Lagrangian trajectories (every 20 trajs) located on Updrafts bases

Vertical cross section at 1000 UTC



Fast Lagrangian trajectories located close to
Updrafts regions ($W > 0.3 \text{ m.s}^{-1}$)

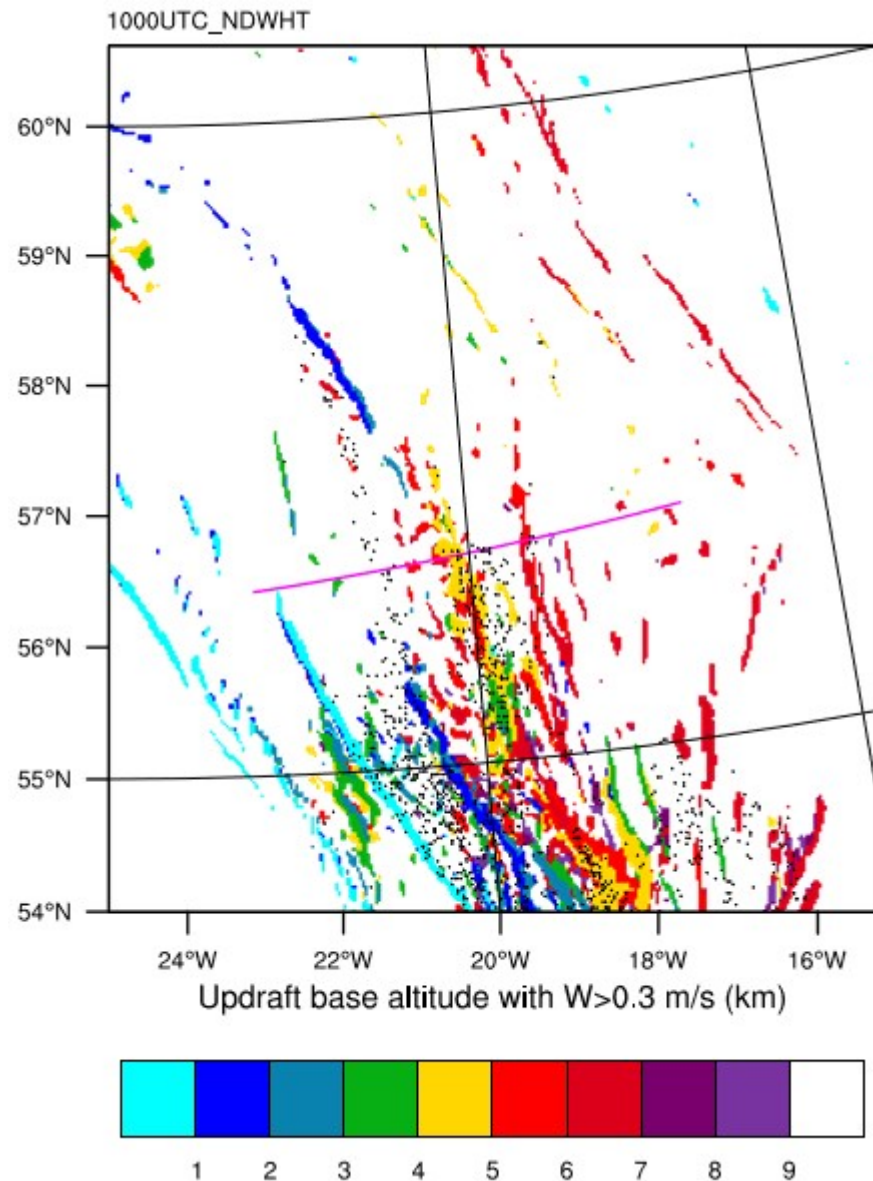
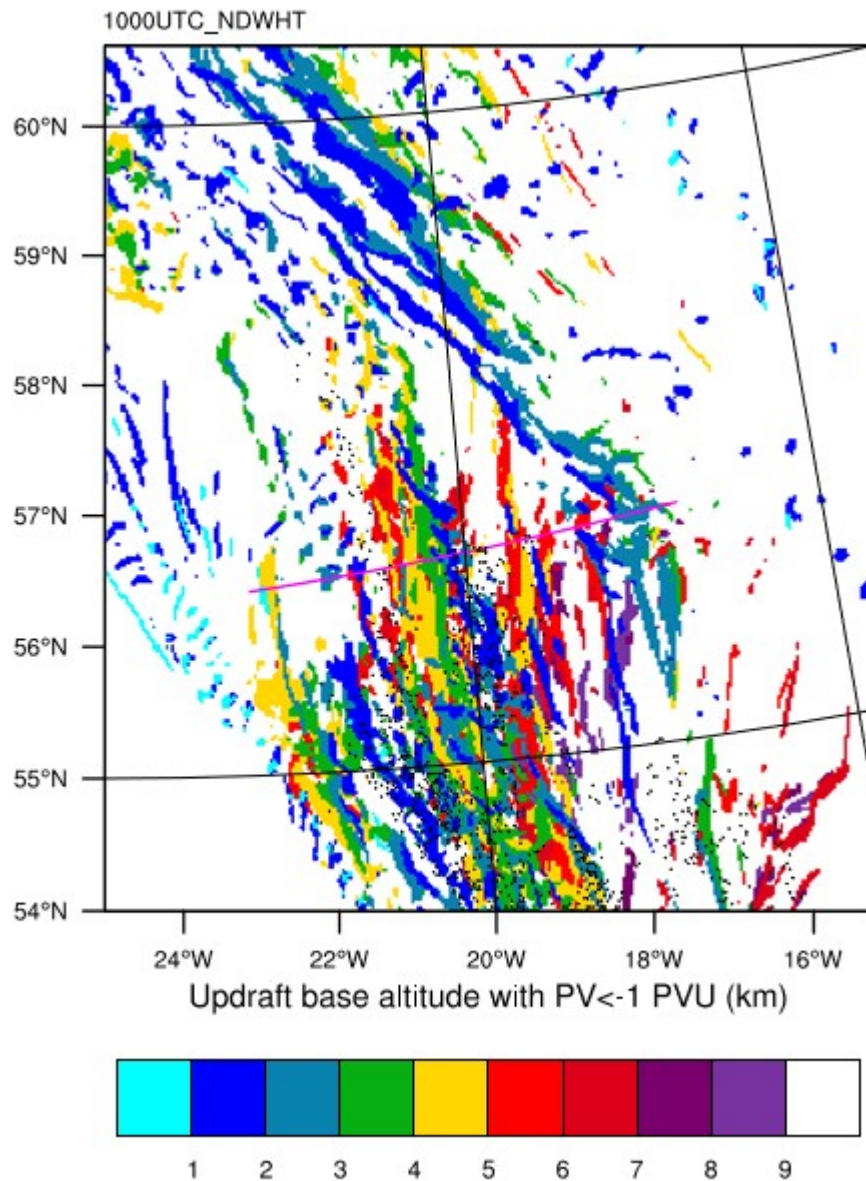
→ Mid-level / High-level convection



PV dipoles located in regions of strong vertical
movements

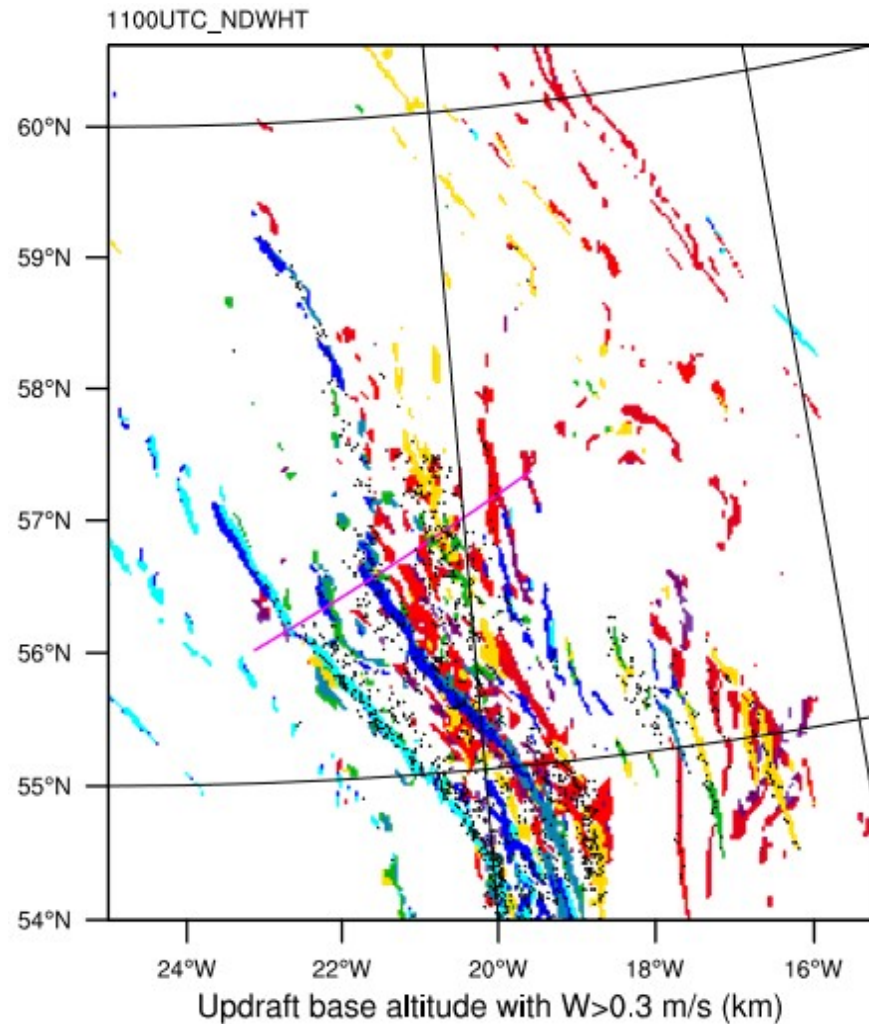
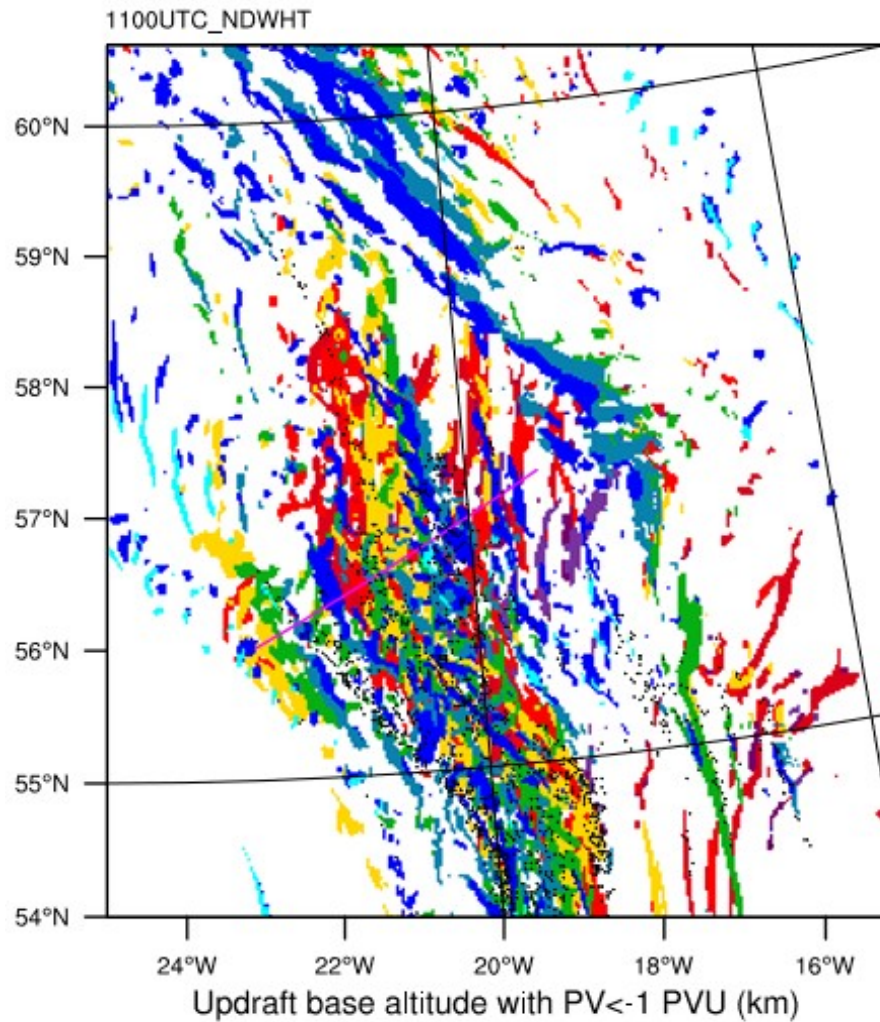
→ Diabatic production of PV dipoles?

Negative PV bands at 1000 UTC

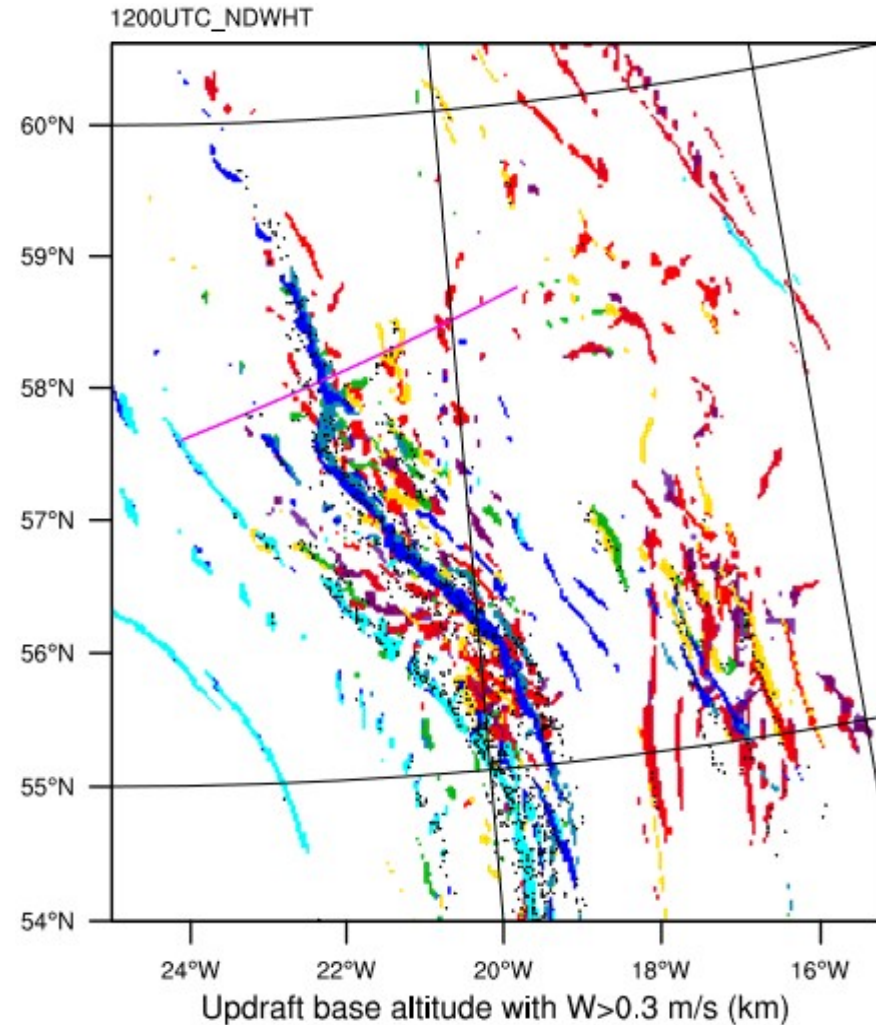
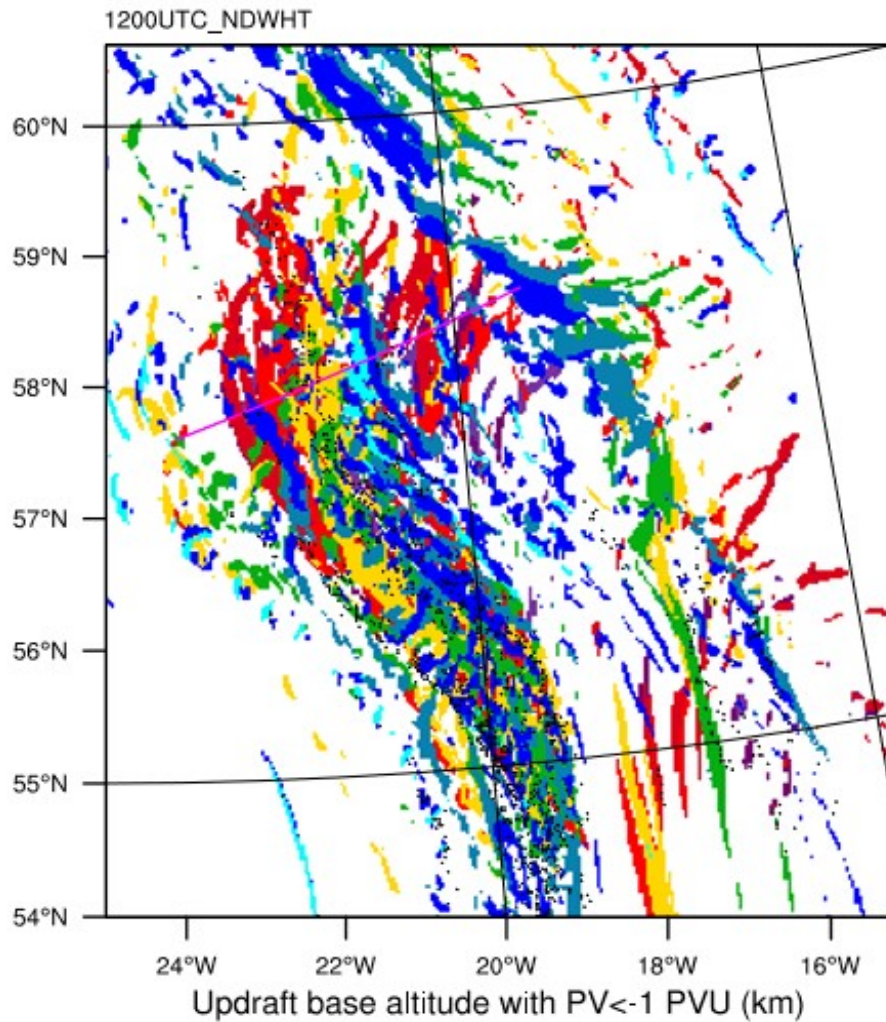


Narrow negative PV bands base located at ~5 km of altitude
as well as Updrafts base

Negative PV bands at 1100 UTC

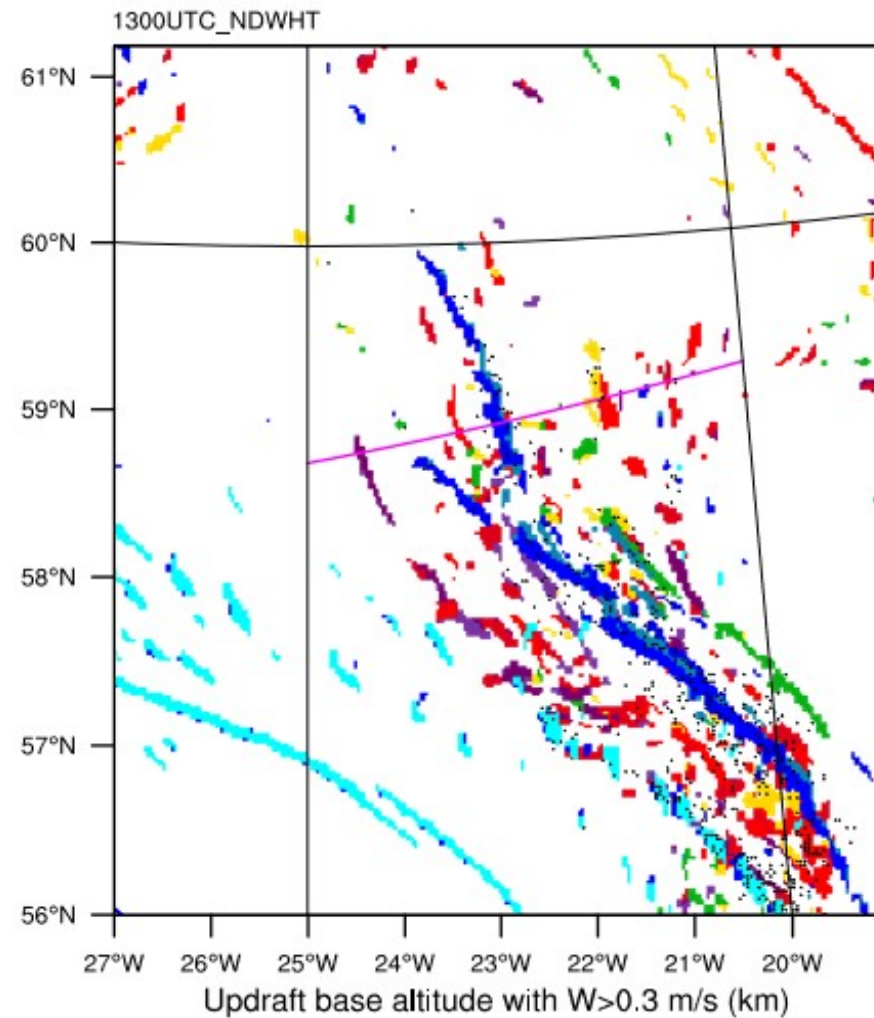
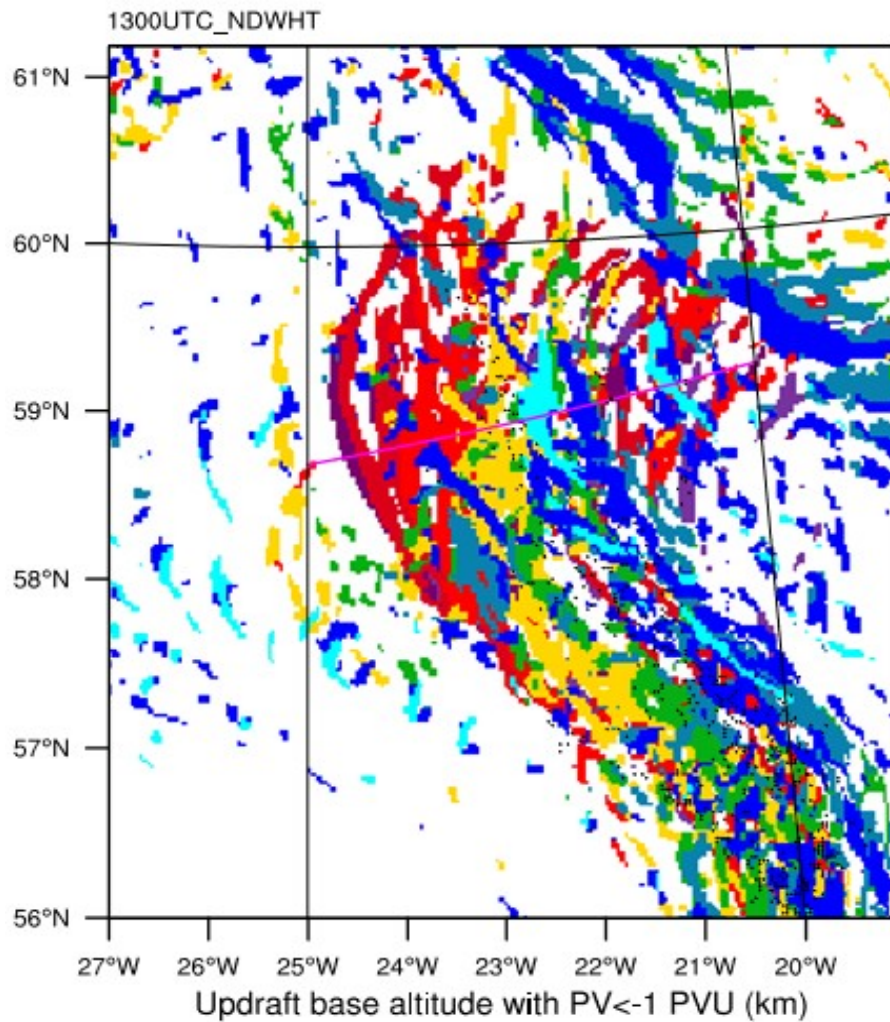


Negative PV bands at 1200 UTC



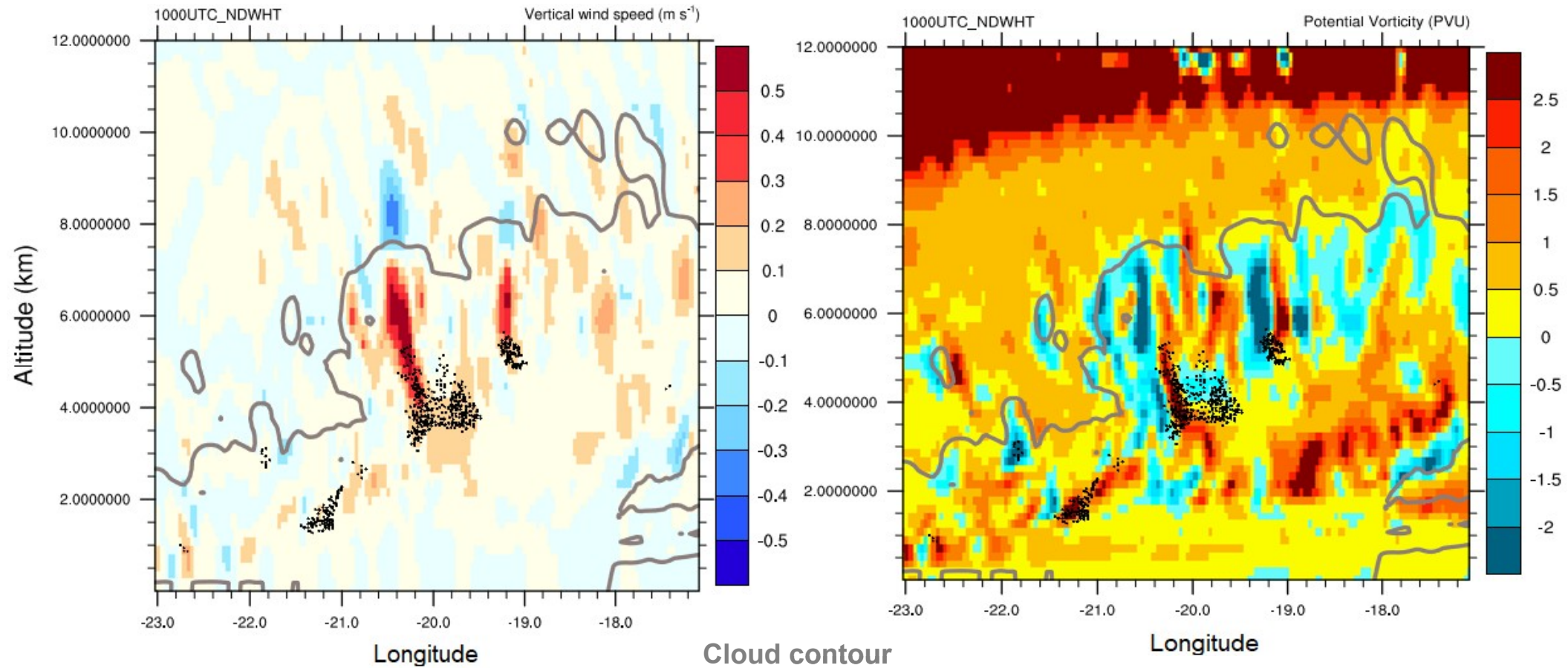
Negative PV bands base located between 5 and 6 km of altitude

Negative PV bands at 1300 UTC



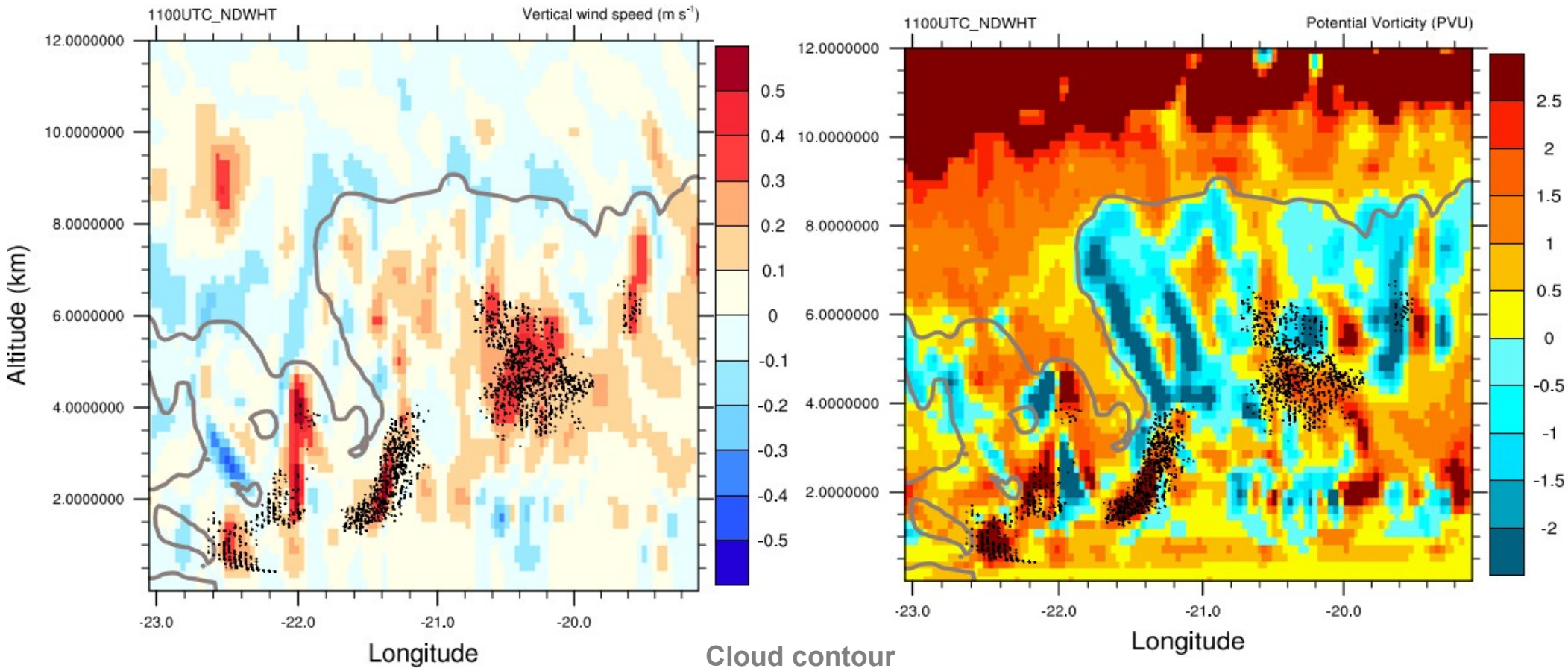
Negative PV bands base located between 6 and 7 km of altitude
as well as Updrafts base

Vertical cross section at 1000 UTC



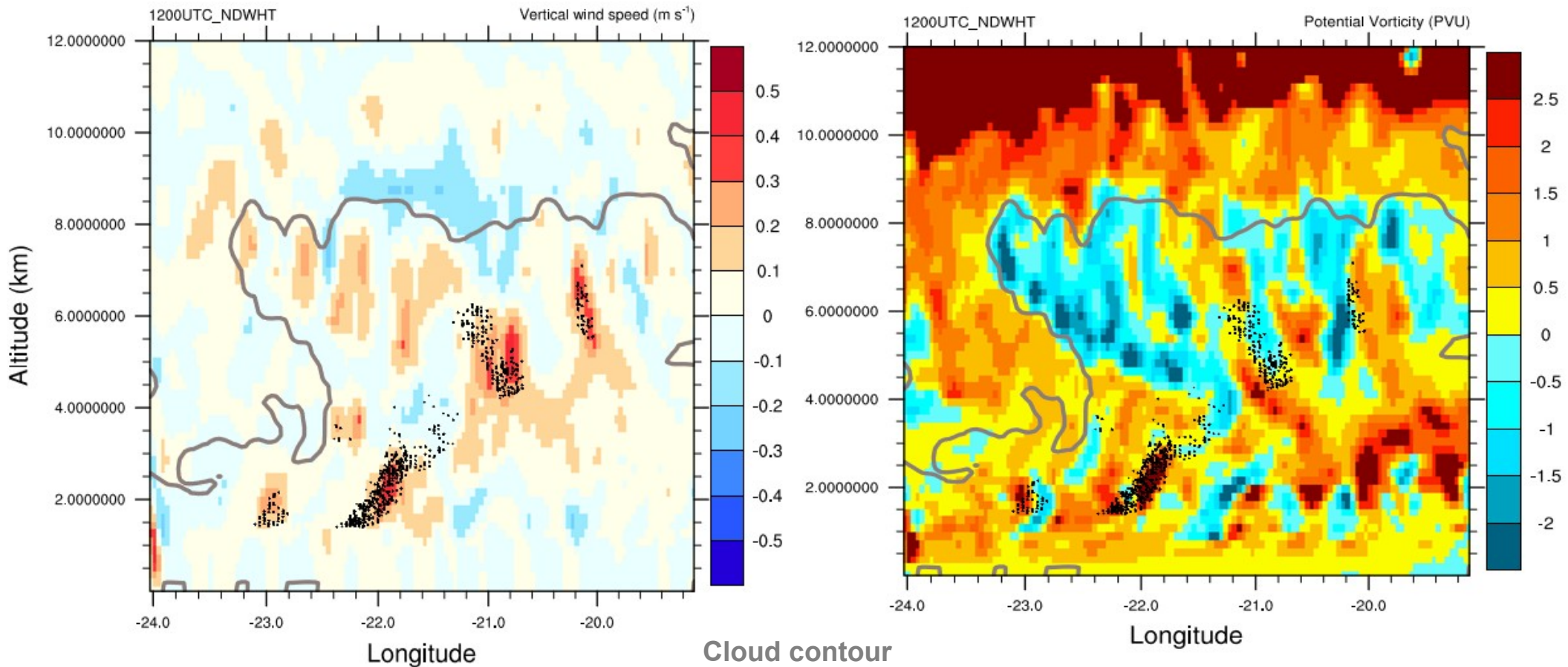
- Cloud top at $z \sim 8$ km
- **Negative PV** in altitude located in mid-level convection regions
 - Other PV dipoles in convection regions within the cloud

Vertical cross section at 1100 UTC



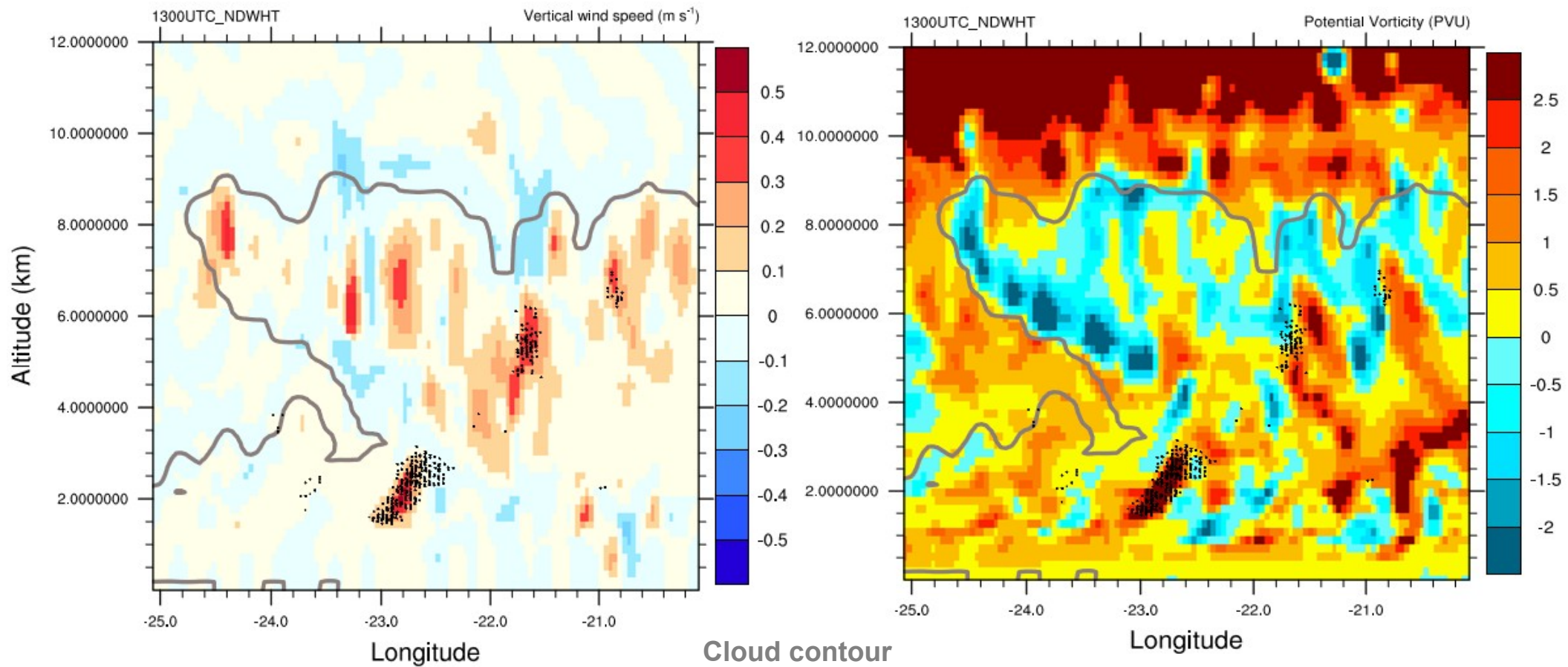
- Cloud top at $z \sim 8.5$ km
- **Vertically elongated negative PV band** (advected) close to the edge of the cloud
 - Other PV dipoles in Updrafts regions within the cloud
 - **Strong positive** $\text{PV} > 2$ PVU in low-layers, in Updrafts regions

Vertical cross section at 1200 UTC



- Cloud top at $z \sim 8.5$ km
- **Horizontally elongated** negative PV band (advected) close to the edge of the cloud
- **A priori disconnected from Updrafts regions**
- **Strong positive** $\text{PV} > 2$ PVU in low-layers, in Updrafts regions

Vertical cross section at 1300 UTC



- Cloud top at $z \sim 8.5$ km
- **Horizontally elongated** negative PV band close to the edge of the cloud
 - A priori disconnected from Updrafts regions?
- **Strong positive** $PV > 2$ PVU in low-layers, in Updrafts regions

End of the negative PV bands?

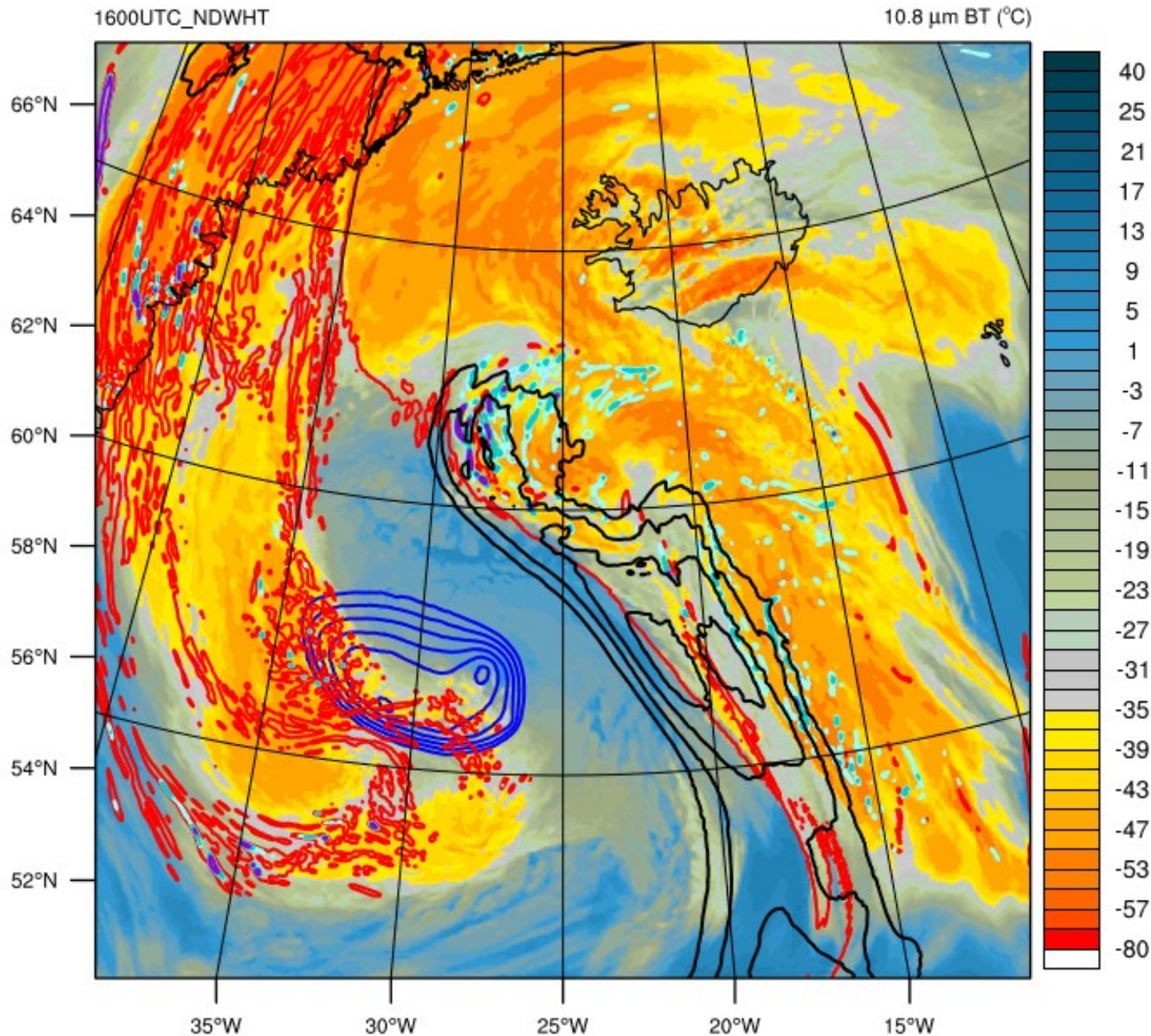
1600 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

End of the negative PV bands?

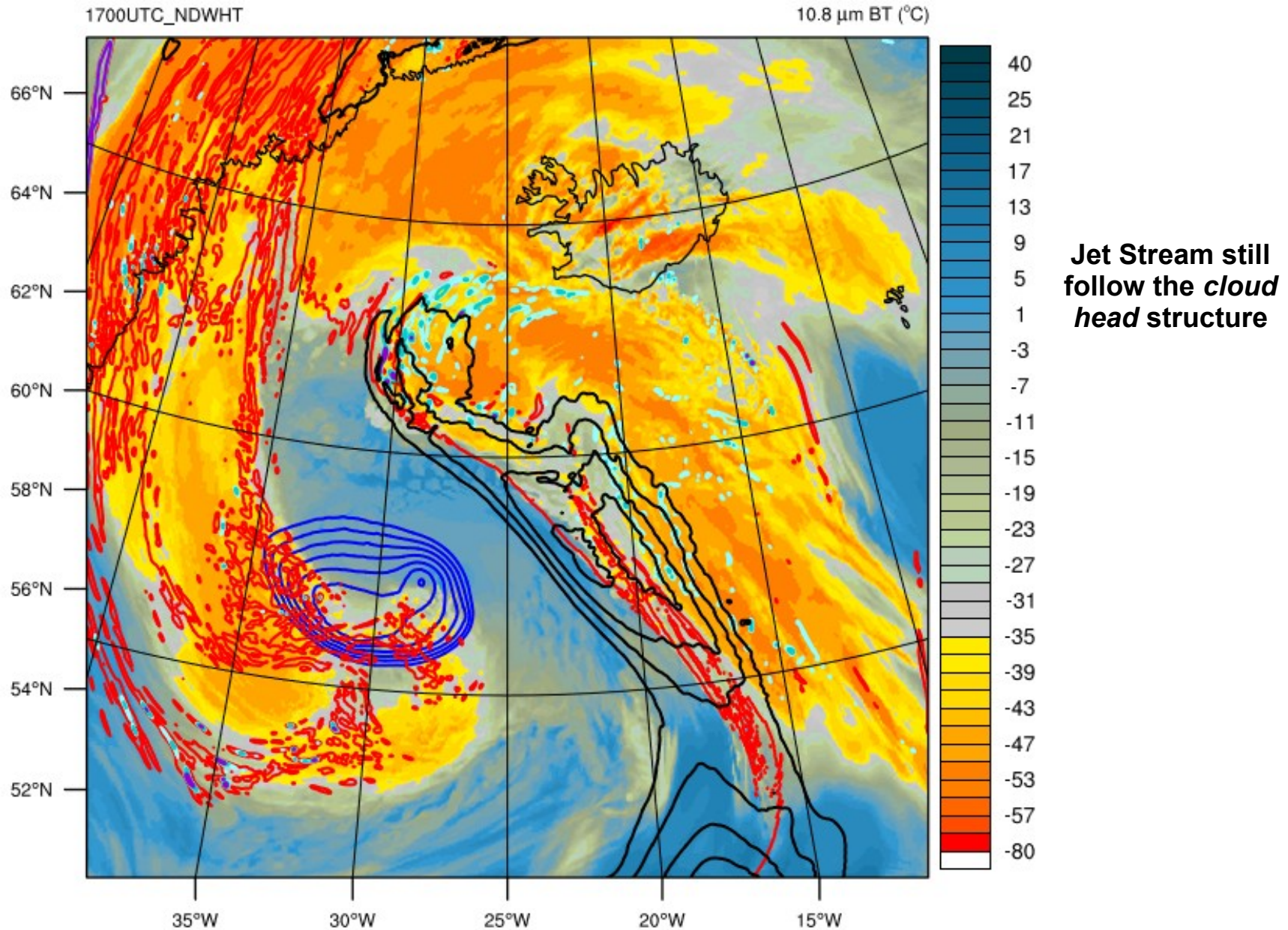
1700 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

End of the negative PV bands?

1700 UTC

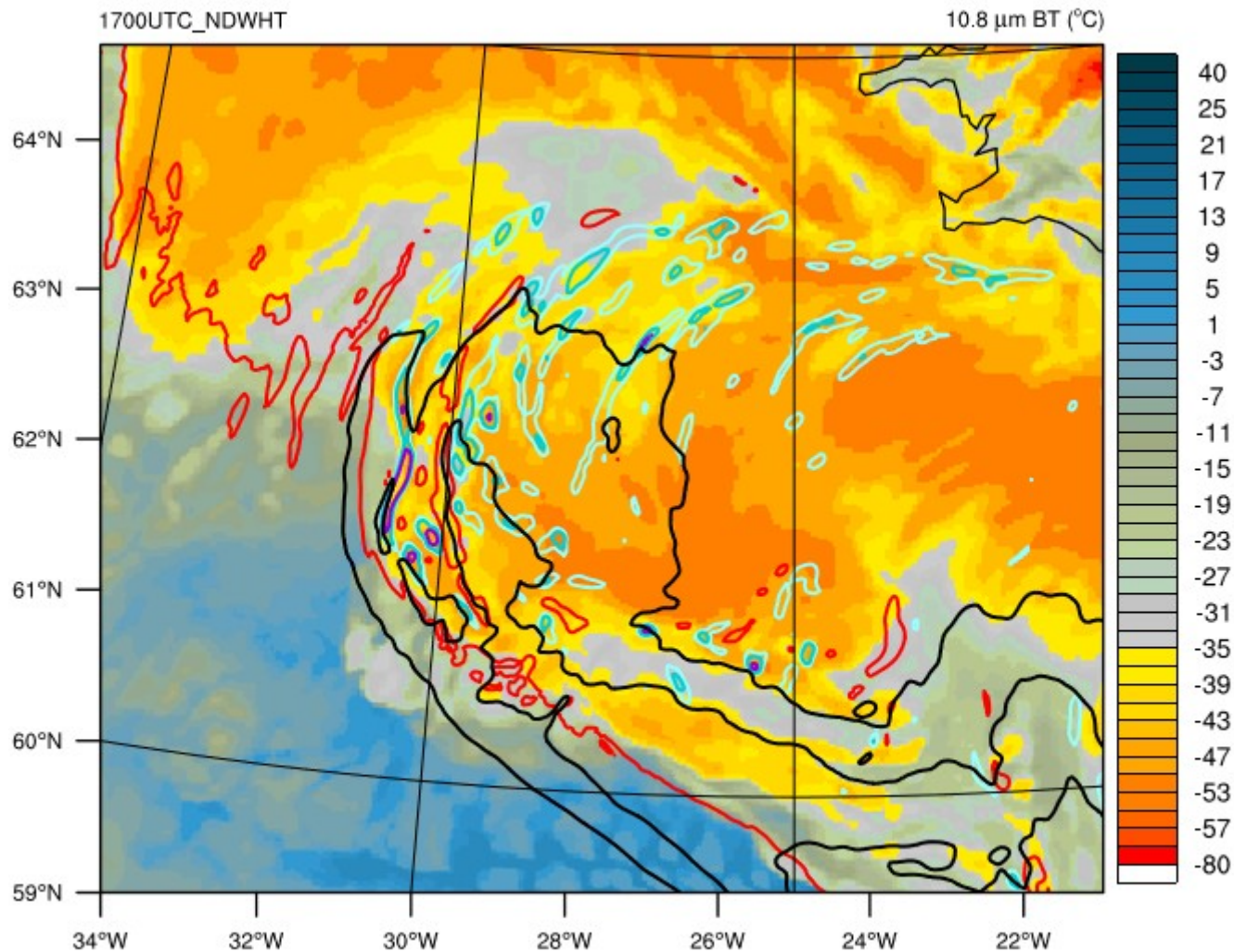
Zoom

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



Jet Stream is locally intensified by the PV gradient

at $\theta=320$ K

End of the negative PV bands?

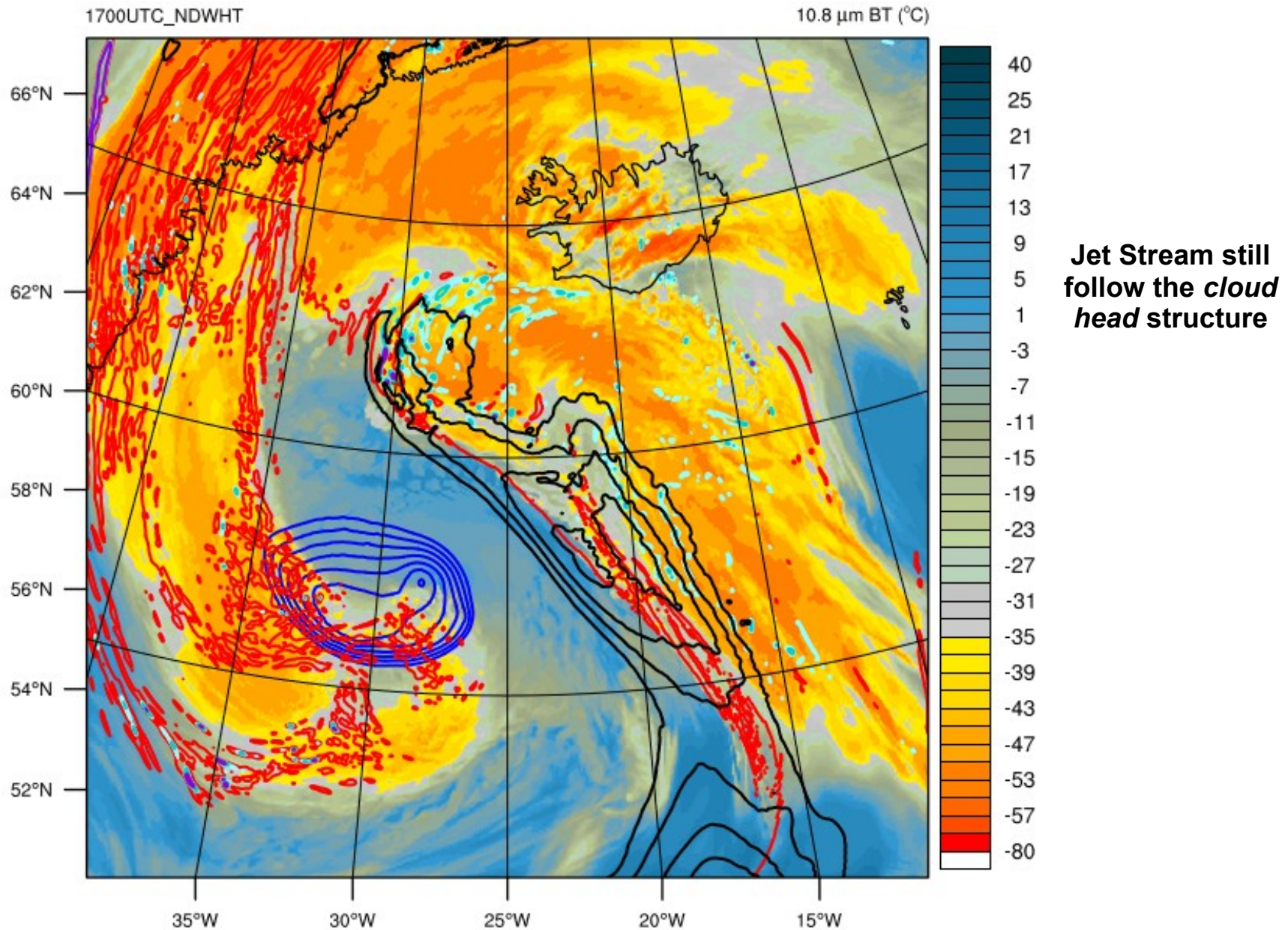
1700 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



at $\theta=320$ K

End of the negative PV bands?

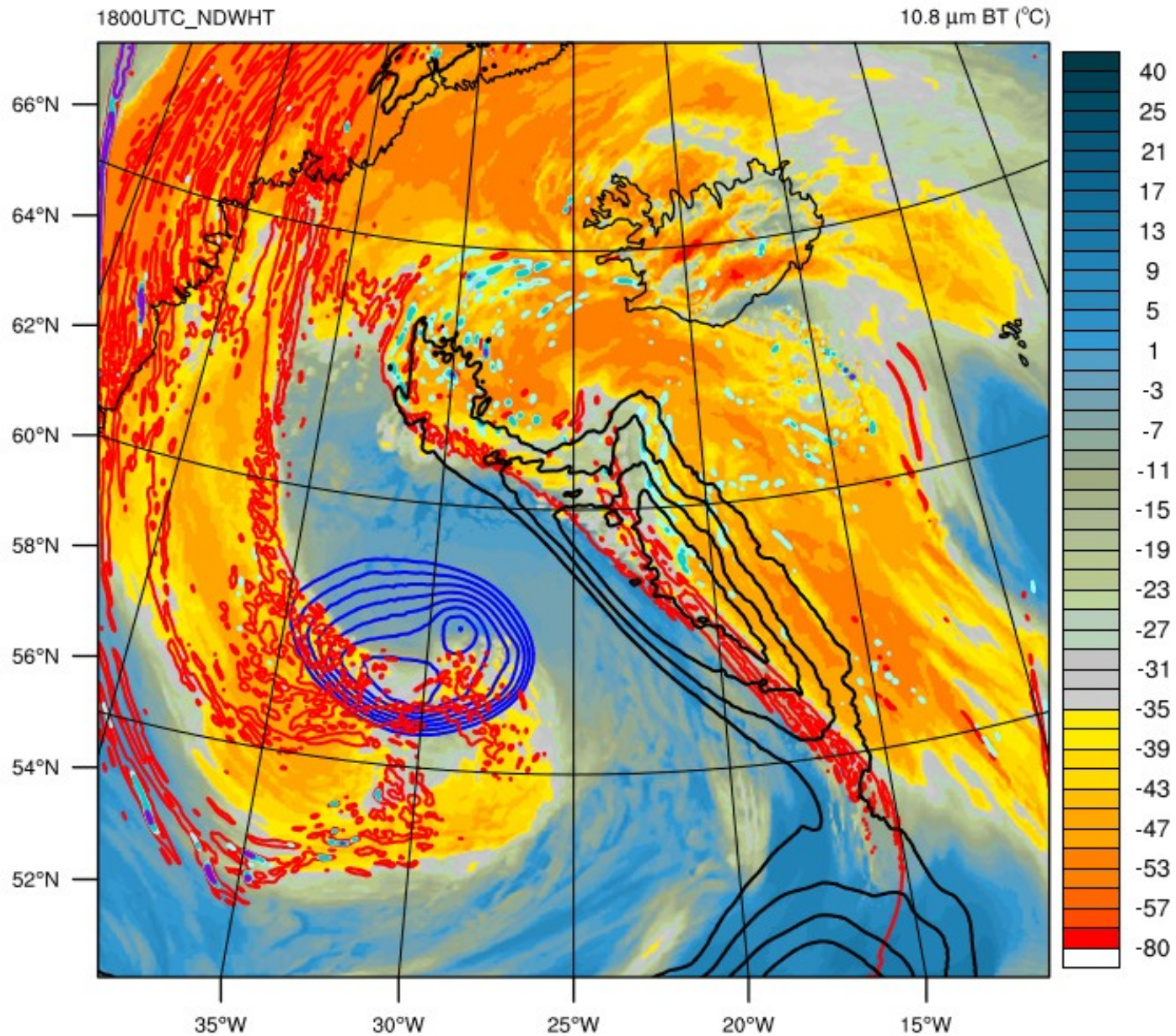
1800 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

End of the negative PV bands?

1800 UTC

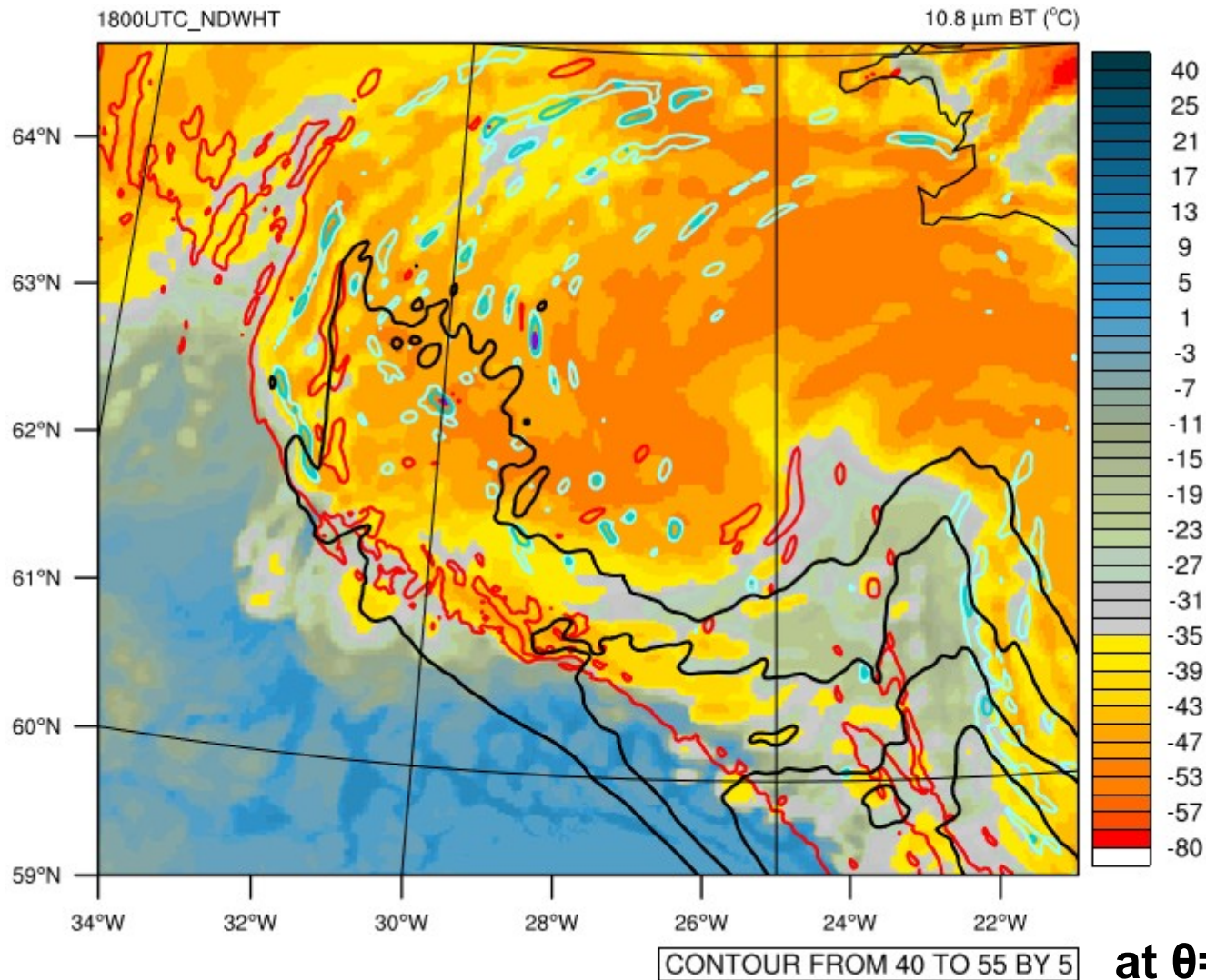
Zoom

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



End of the negative PV bands?

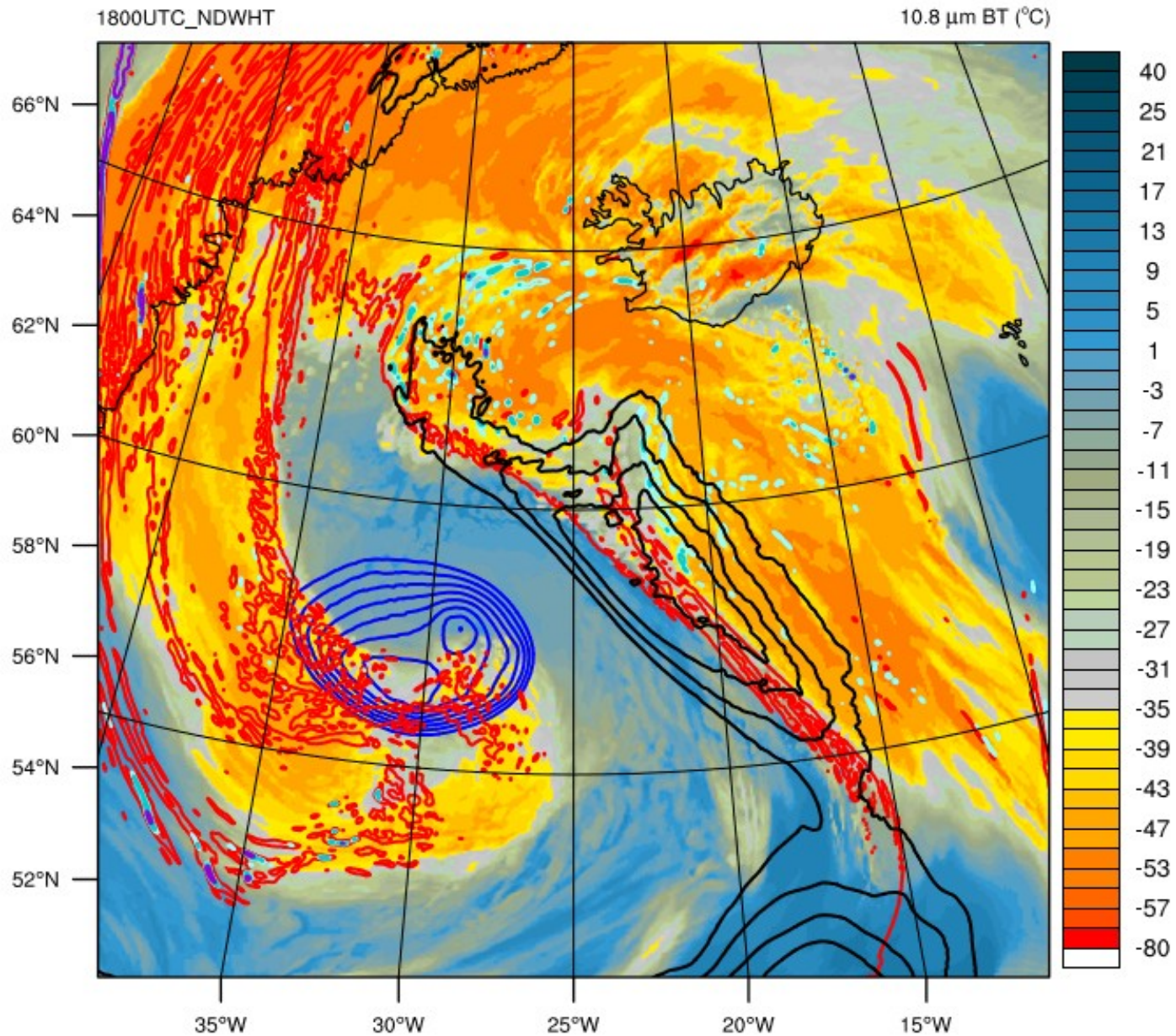
1800 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

End of the negative PV bands?

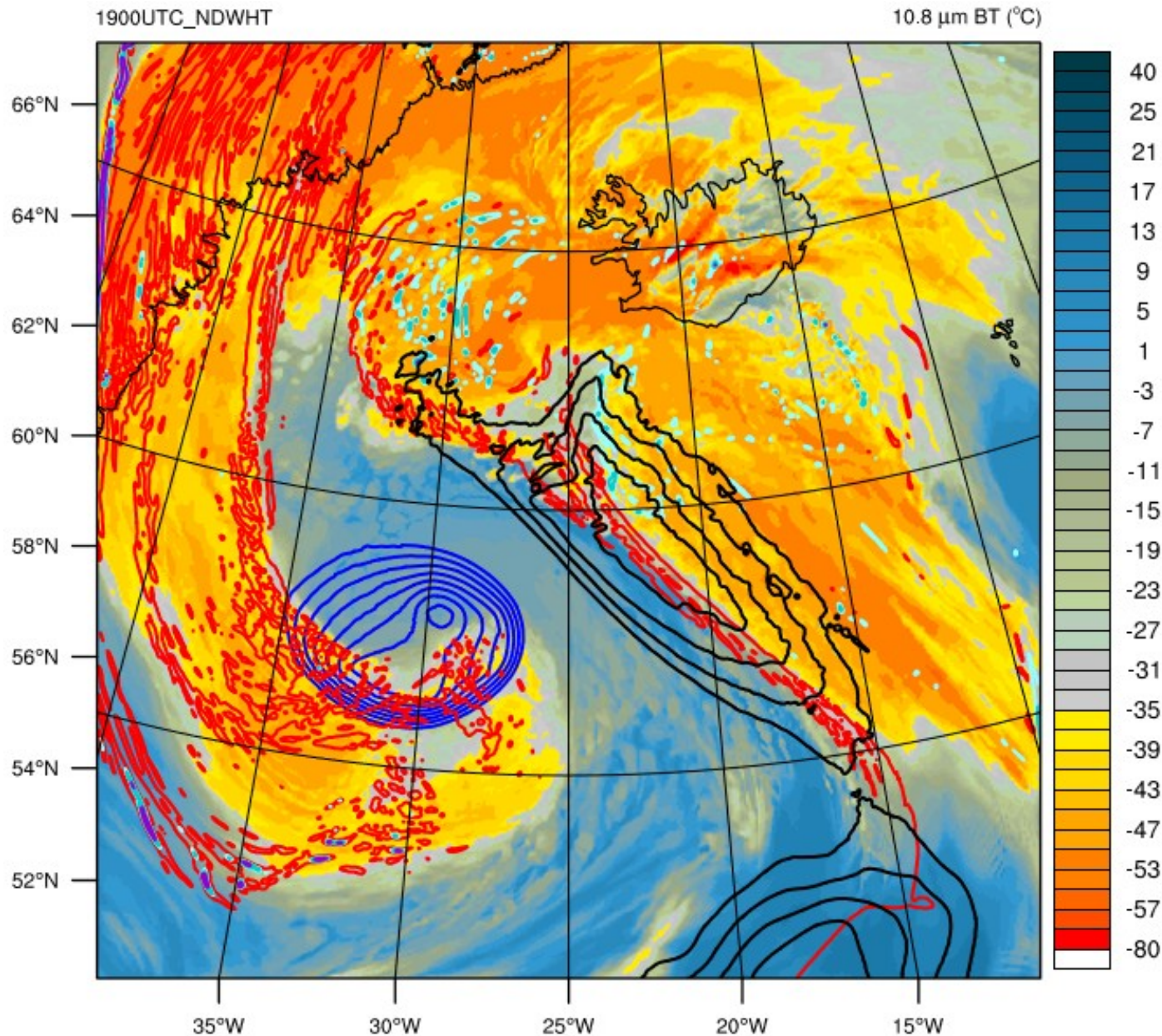
1900 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



North part of the
Jet Stream
disappears
because of the
cloud head and
negative PV
bands advection

CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

End of the negative PV bands?

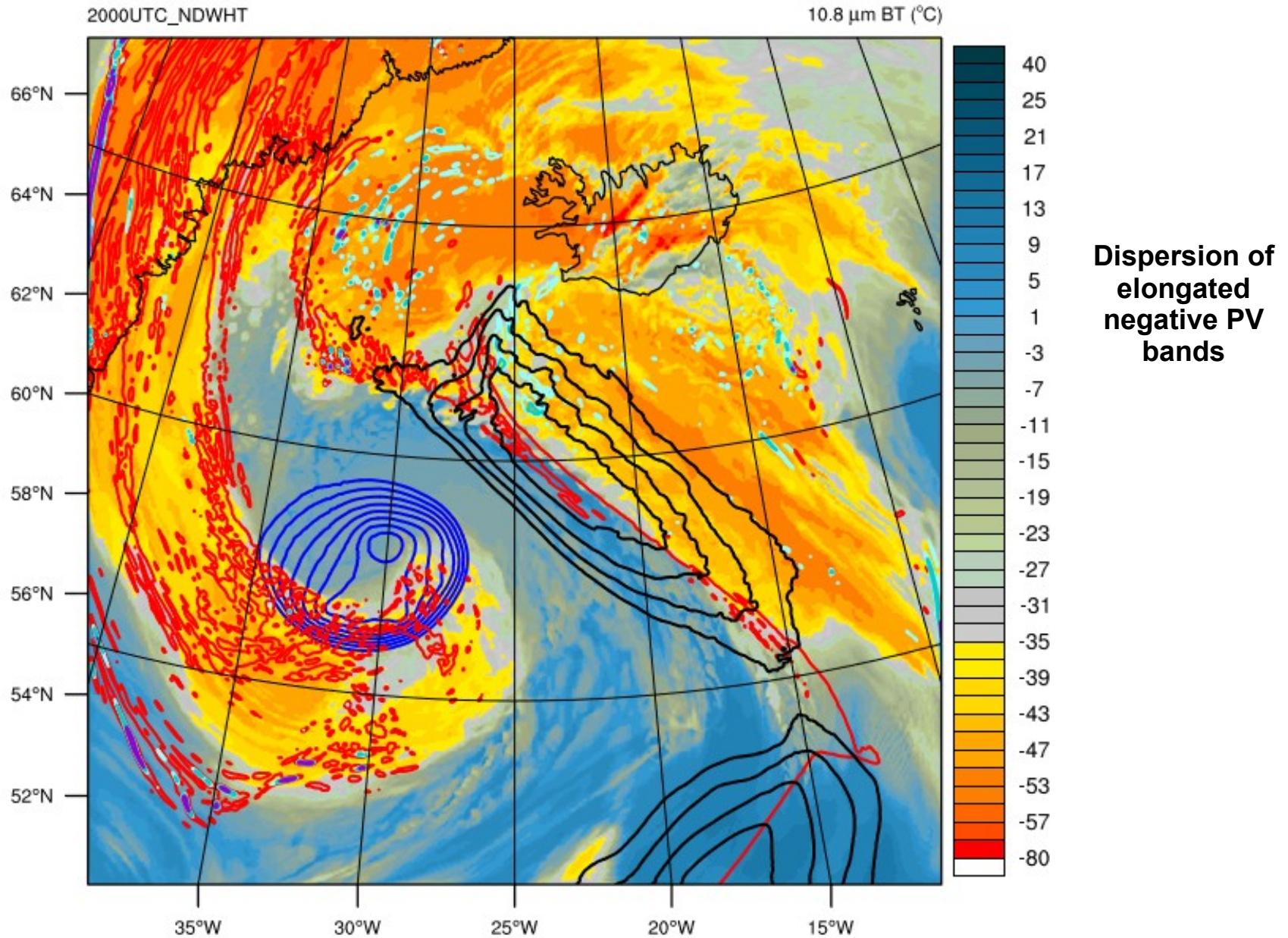
2000 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

End of the negative PV bands?

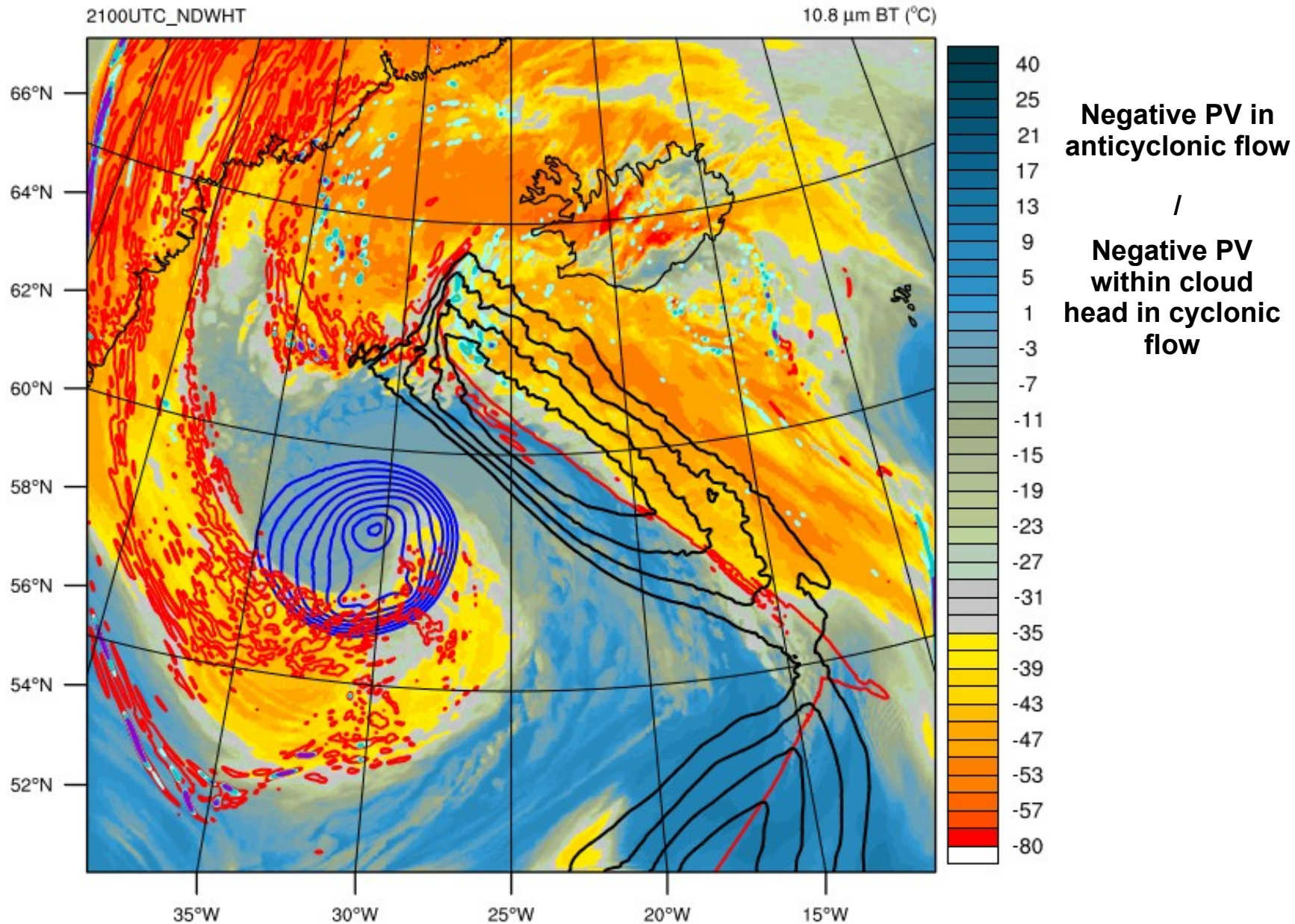
2100 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



CONTOUR FROM 40 TO 55 BY 5

at $\theta=320$ K

End of the negative PV bands?

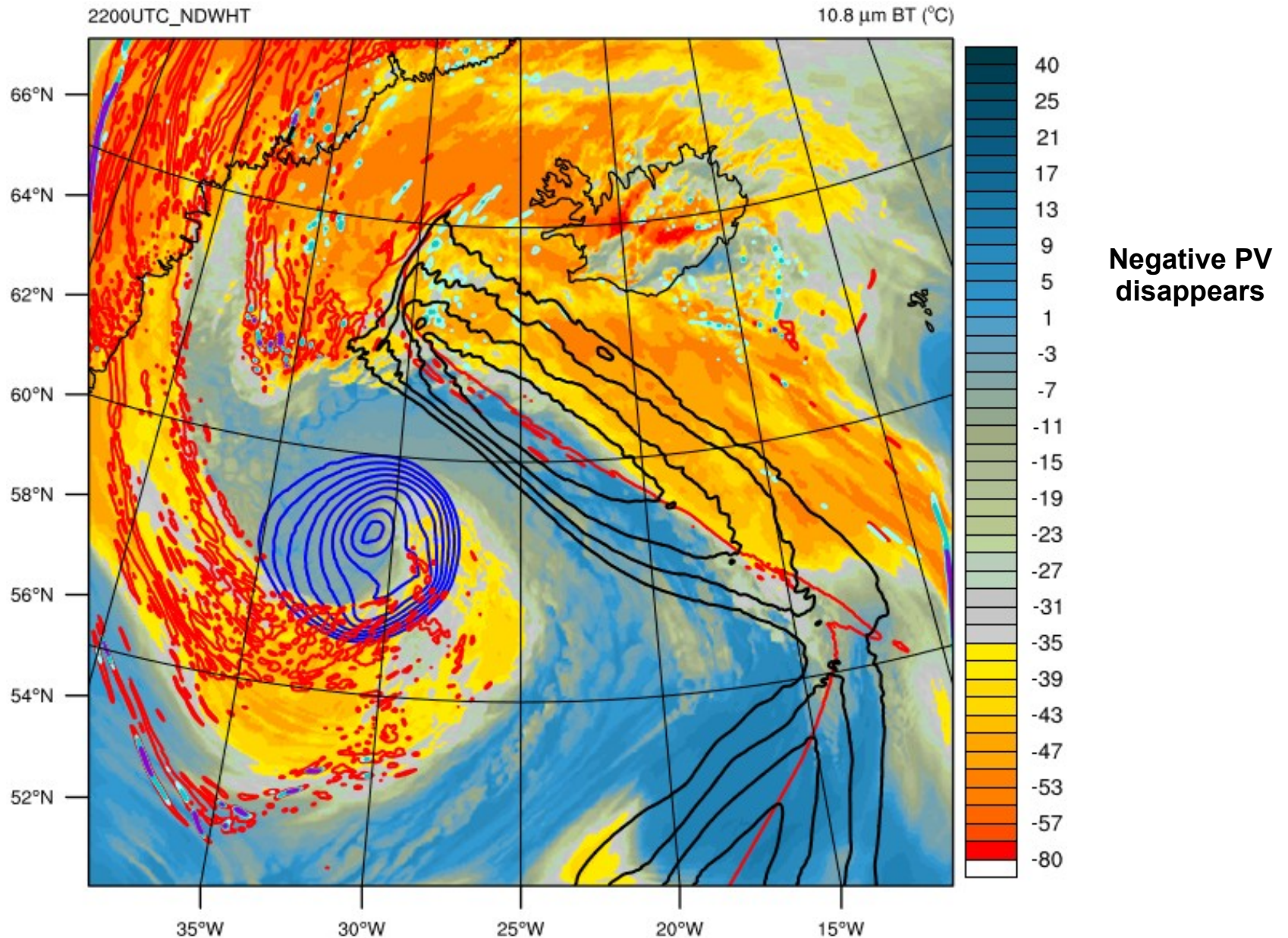
2200 UTC

+2 PVU

-0.5 / -1 PVU

-2 PVU

MSLP



at $\theta=320$ K