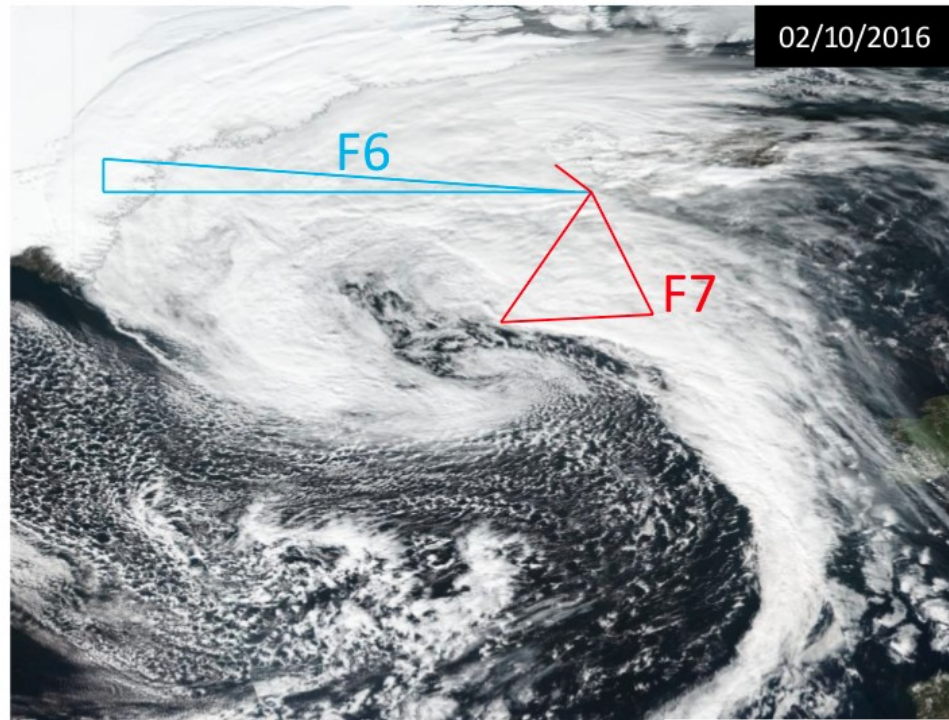


Influence of diabatic processes in warm conveyor belt of the Stalactite Cyclone in Meso-NH simulations and comparison between two microphysical schemes



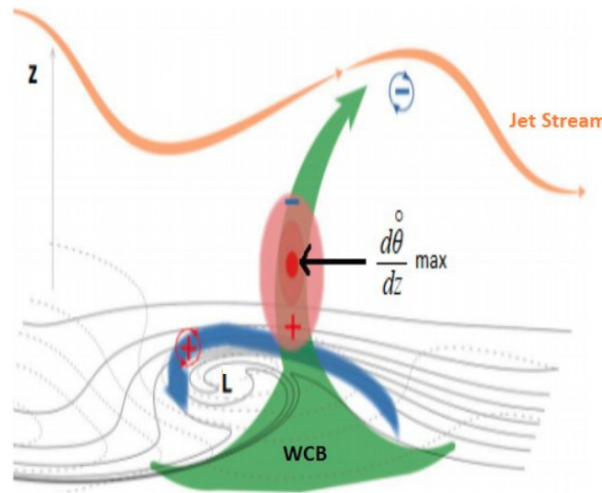
MODIS

Figure M. Wimmer

Marie Mazoyer (CNRM), Philippe Arbogast (CNRM), Didier Ricard (CNRM), Gwendal Rivière (LMD), Julien Delanoë (LATMOS), Quitterie Cazenave (LATMOS), Jacques Pelon (LATMOS), Benoit Vié (CNRM), Christine Lac (CNRM)

Diabatic processes in warm conveyor belt

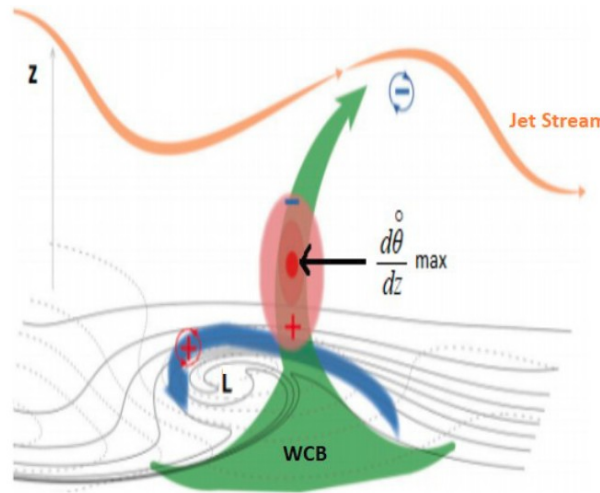
- Extratropical cyclones : Prediction error in numerical weather prediction models -> **PV shift at the jet stream level**
- May be due to **representation of diabatic processes** through latent heat release (Roberston and Smith, 1983)
- **Microphysical** processes of importance (Joos and Wernli, 2012, Joos and Forbes, 2016)
- Many **uncertainties** in representation of warm-phase, mixed-phase and ice-phase microphysics and the role of aerosols (Thompson and Eidhammer, 2014)



Joos and Wernli (QJRMS, 2012)

Diabatic processes in warm conveyor belt

- Extratropical cyclones : Prediction error in numerical weather prediction models -> **PV shift at the jet stream level**
- May be due to **representation of diabatic processes** through latent heat release (Roberston and Smith, 1983)
- **Microphysical** processes of importance (Joos and Wernli, 2012, Joos and Forbes, 2016)
- Many **uncertainties** in representation of warm-phase, mixed-phase and ice-phase microphysics and the role of aerosols (Thompson and Eidhammer, 2014)

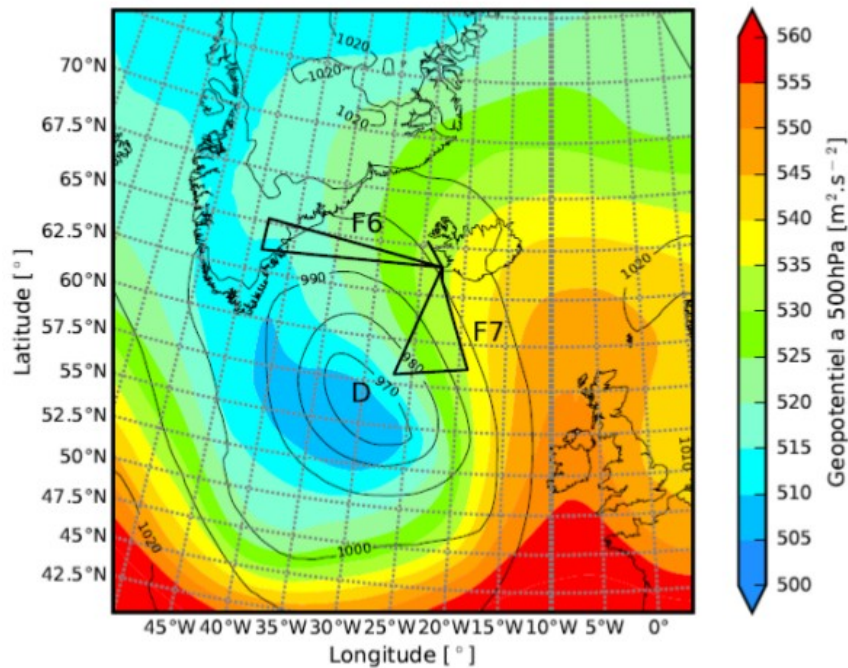


Joos and Wernli (QJRM, 2012)

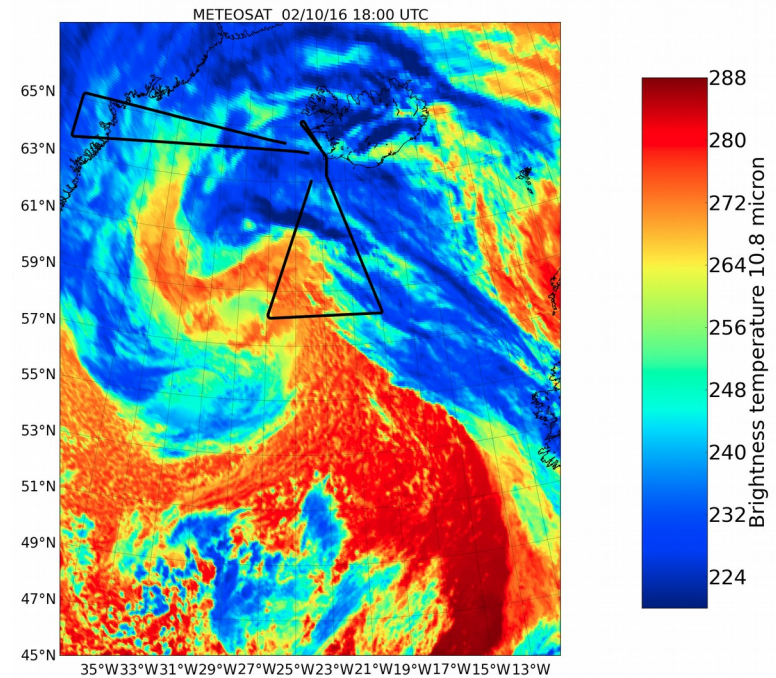
➡ Objectives : Comparison of two distinct microphysical schemes
(→ **mixed and ice phase and distribution's moments**)
and their impact on a WCB representation

02/10/2016 - Stalactite cyclone - Nawdex campaign

➡ Blocking situation over scandinavia during ~ 10 days
Apogee 02/10/2016 21:00 UTC (960 hPa)



Arpege analysis 02/10/2016 12:00



*Meteosat brightness temperature at
10.8 μm 02/10/2016 18:00*

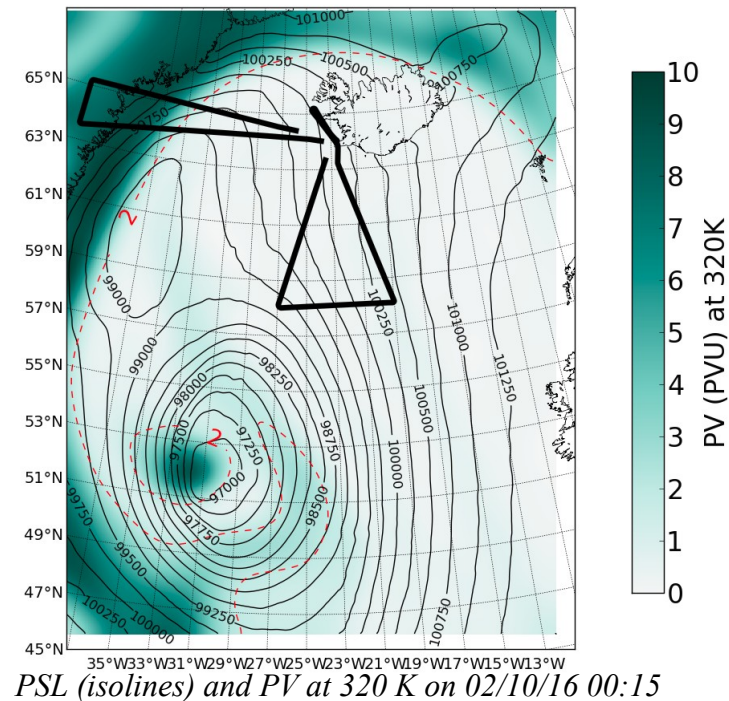
Falcon 20 flights : F6 - 09:30 → 11:30

F7 – 13:00 → 16:00

➡ RASTA embedded (Reflectivity, Ice Water Content, Wind)
+ F6 : Dropsondes (P,T,U,RH)

02/10/2016 - Experiences

- **Méso-NH** (Lac et al, 2018)
- 02/10/16 00h → 03/10/16 00h Output : every 15min
- CI and forcing : Arpege
- Domain : XY → **2.5 km*2.5 km** 800*1000 pts (Convection resolved)
Z → 56 verticals levels ~[0-20]km



Obs 'model' :

- +Flights collocated et cotemporalized
- +Radar simulator
- + 'Satellite' simulator (RTTOV Saunders et al, 2002)

- **Microphysics** : ICE3 (Pinty and Jabouille, 1998)
LIMA (Vié et al, 2016)

Microphysical schemes

ICE3

Droplet, rain, graupel, snow, ice

One moment (r)

➡ **Concentration evolution with LIMA**

- Simulation LIMA : aerosols loading homogeneous in space and configuration similar to ICE3
- Differences on the transition from a specy to an other
- Differences in precipitation velocities
- Radiation interaction

Ice supersaturation

Equilibrium humidity Ice/
Liquid : Linear
temperature function
between 0°C and -20°C

No equilibrium imposed for
cold species
Explicit calculation of vapor
deposition

- ### ➡ **Different vapor deposition for $T < 0^\circ\text{C}$ (on cold species and on droplets)**
- Latent heat release may differ (more important vapor ↔ liquid)**

LIMA

Two moments (r,N)

-Droplet, rain, ice

Influence of diabatic processes in warm conveyor belt of the Stalactite Cyclone in Meso-NH simulations and comparison between two microphysical schemes

00 – Case and simulation presentation

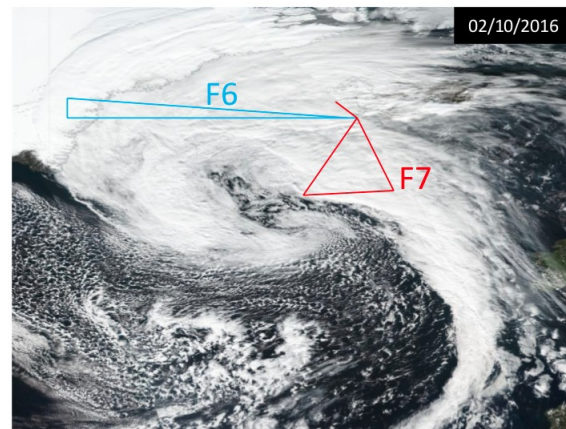
0 – Differences between two schemes ICE3/LIMA

1 – Comparisons with observations

2 – Comparisons of hydrometeors mixing ratio on the whole domain

3 – Impact on altitude anticyclone (lagrangian view point)

4 - Conclusions



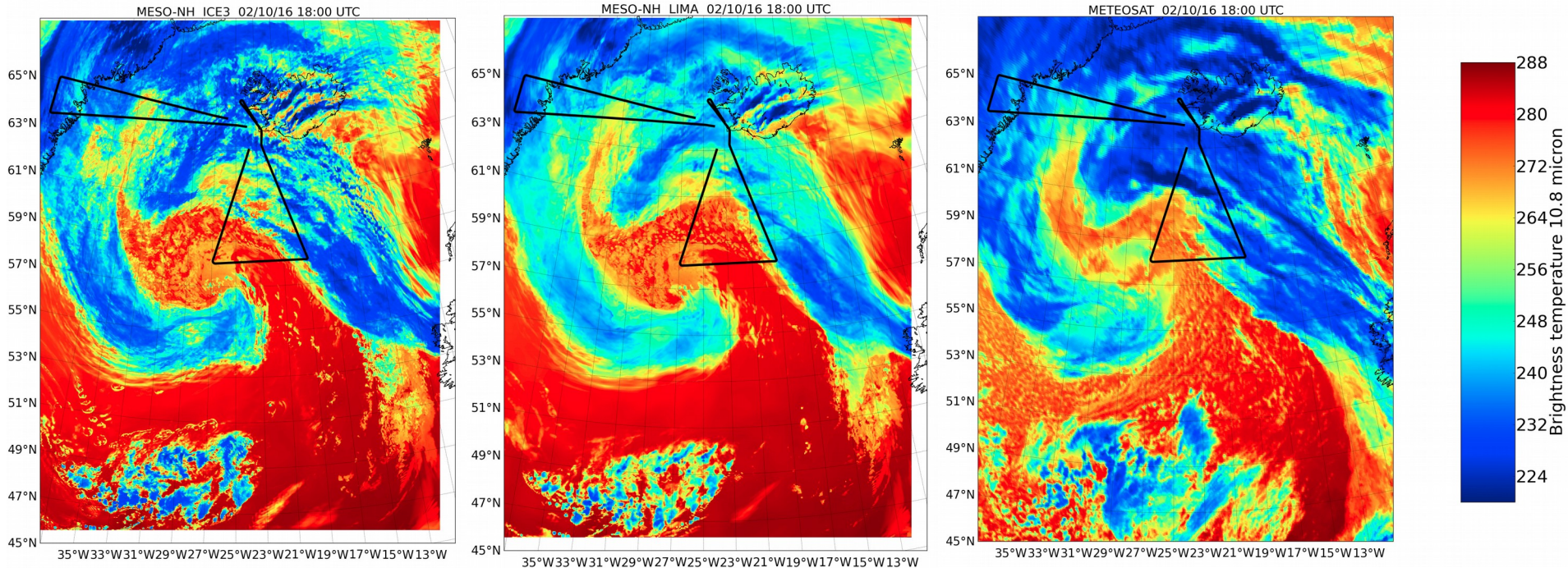
Clouds top and structure

Meteosat, brightness temperature 10.8 μm , 18:00

ICE3

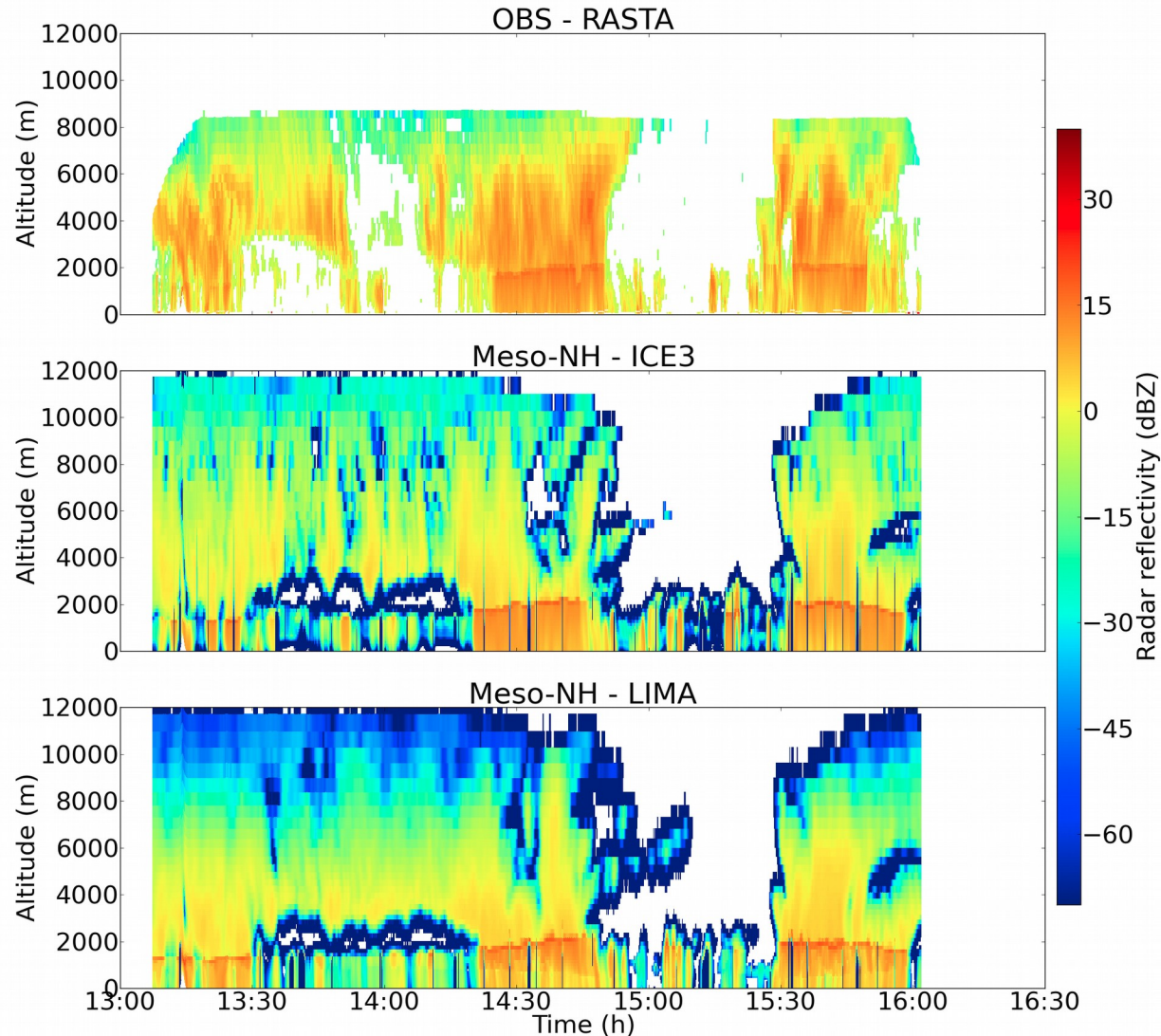
LIMA

OBS



➔ **Clouds structure very similar**
WCB (LIMA/ICE3) more NE than obs
LIMA clouds temperature higher

Radar reflectivity - F7



- Very close structure** (ICE3 less diffuse, subgrid condensation scheme)
- Melting area similar
- Strong Z underestimation
- LIMA top clouds have lower reflectivities** (as satellite data)

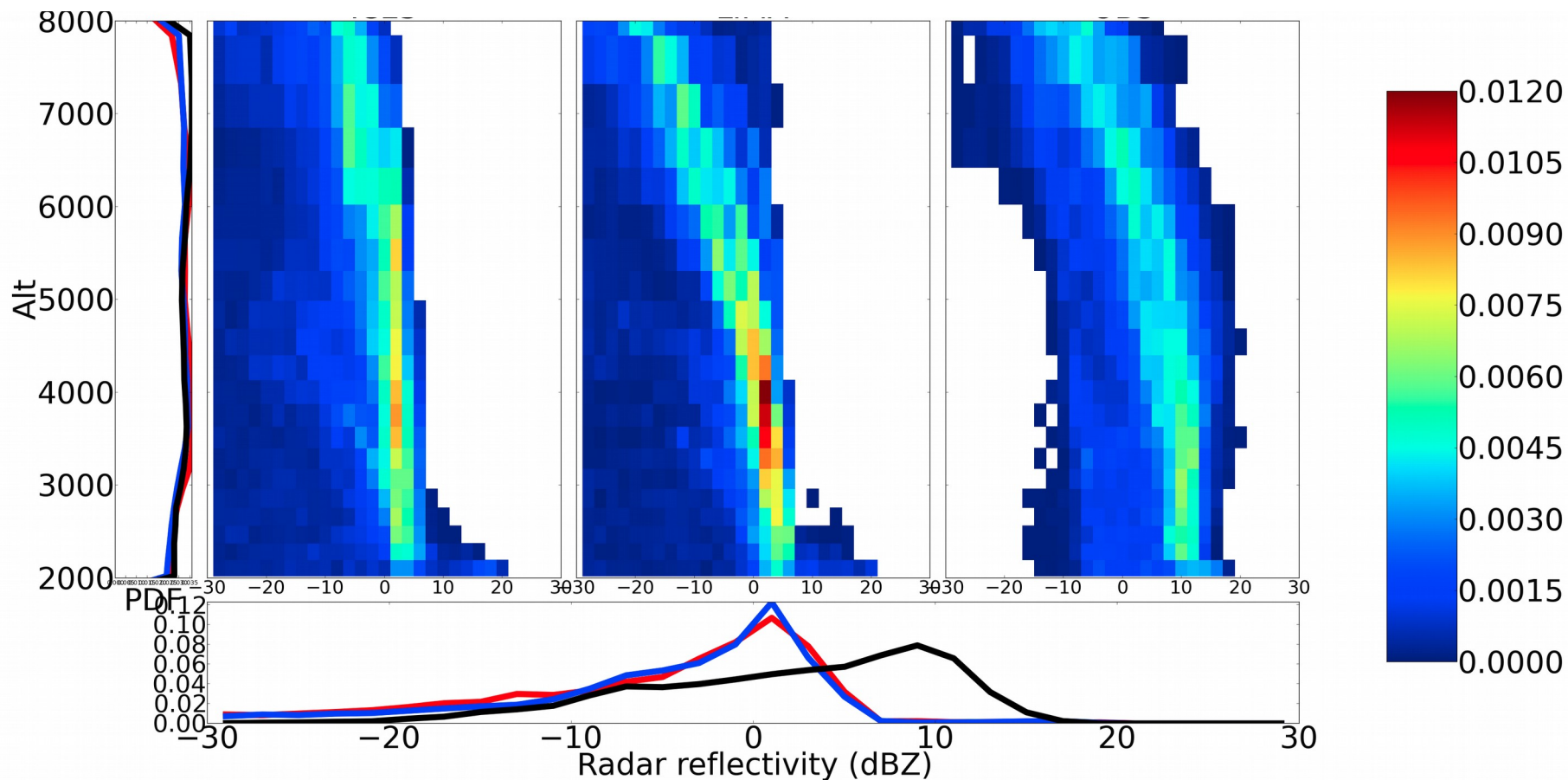


Radar reflectivity – F7 – Bivariate PDF

ICE3

LIMA

OBS



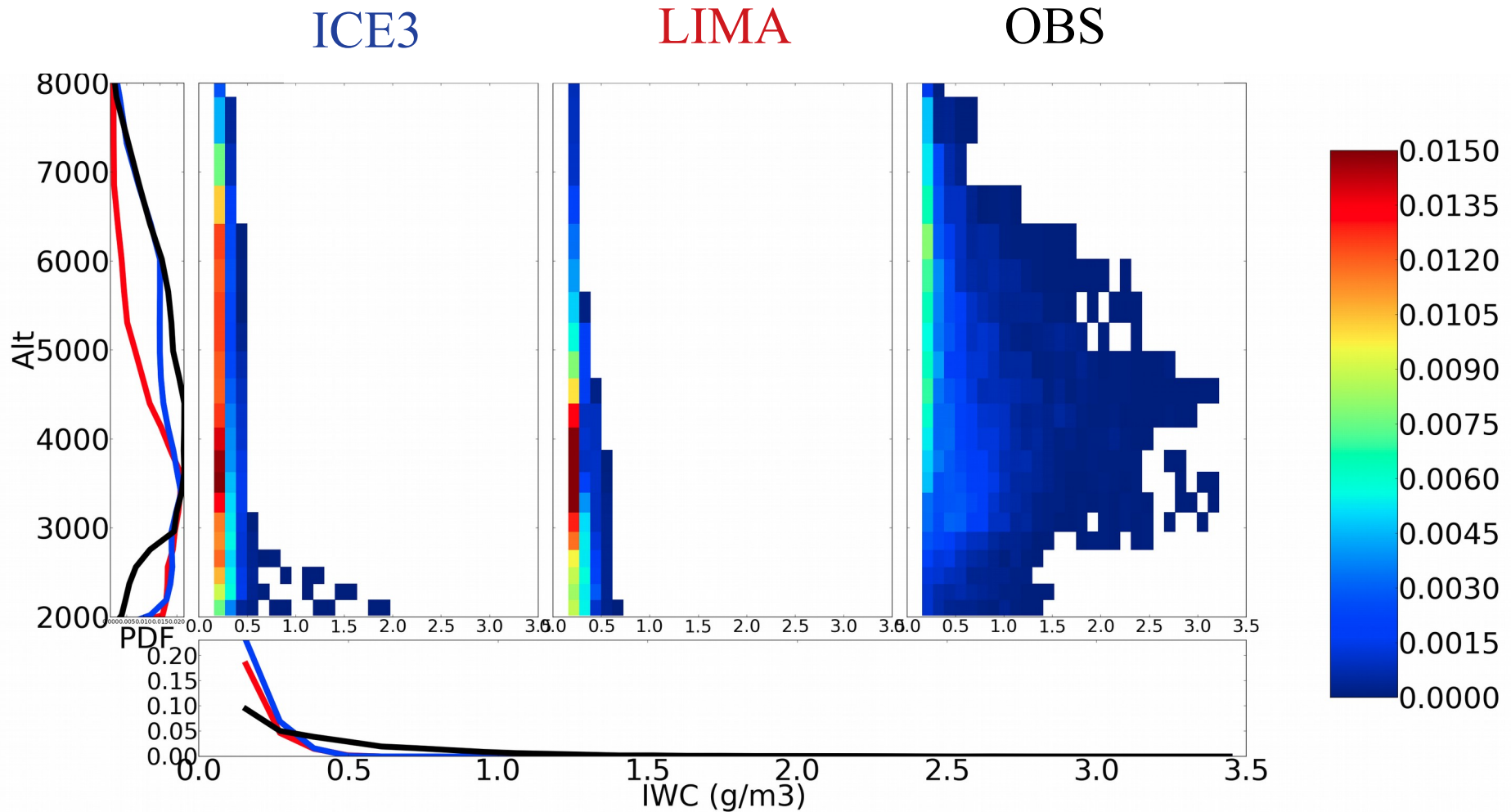
-Shift between models and obs

-PDF radar reflectivity : **Different repartition in intensity**

➔ -ICE3 high intensity in altitude closer to the obs

-LIMA shape slightly closer to the obs (same on HyMex campaign)

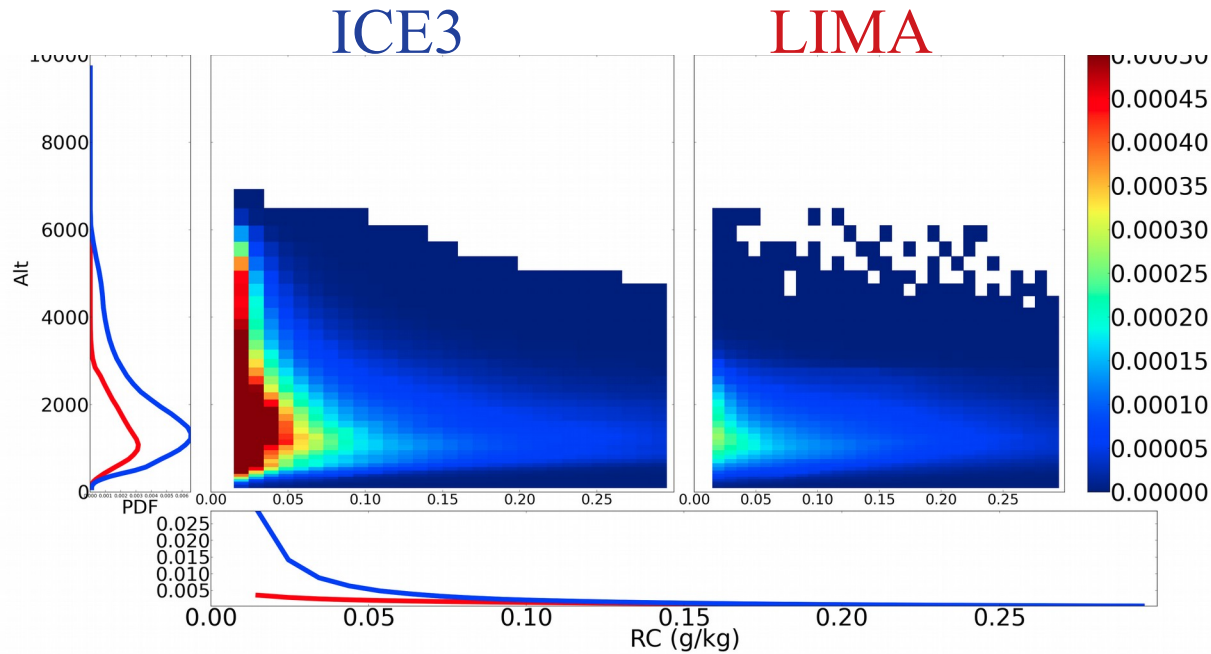
Ice Water Content – F7 – Bivariate PDF



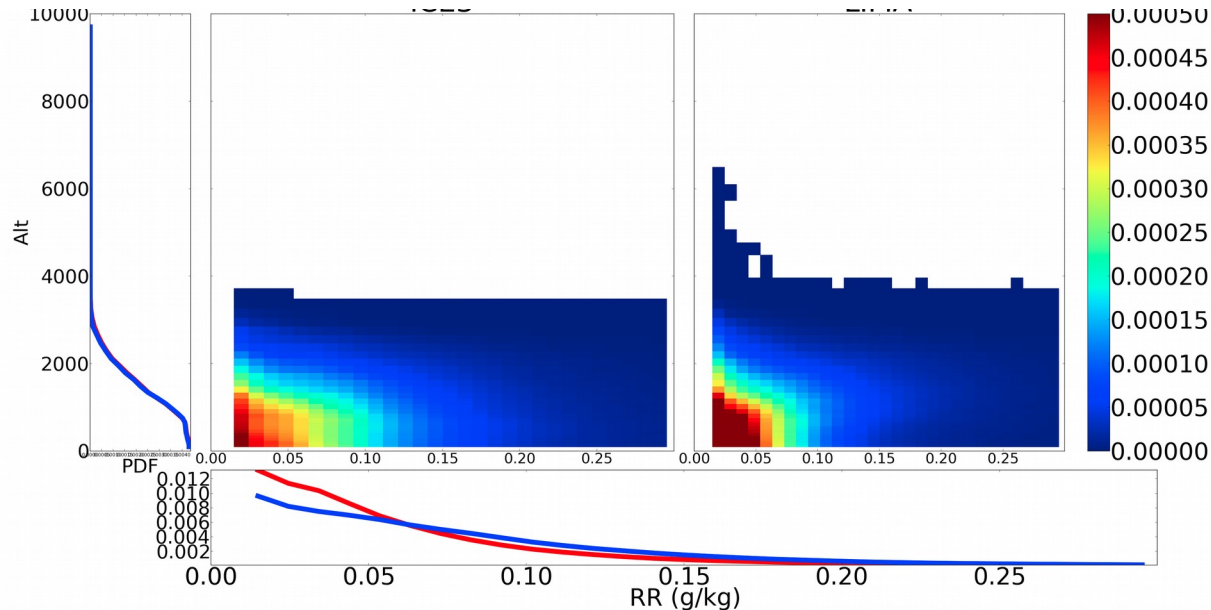
- ➡
- Lower IWC with LIMA and ICE3 than OBS
 - Different repartition with altitude
 - LIMA less ice in altitude

Hot microphysics mixing ratio whole domain (00h, 06h, 12h, 18h, 24h)

Droplets

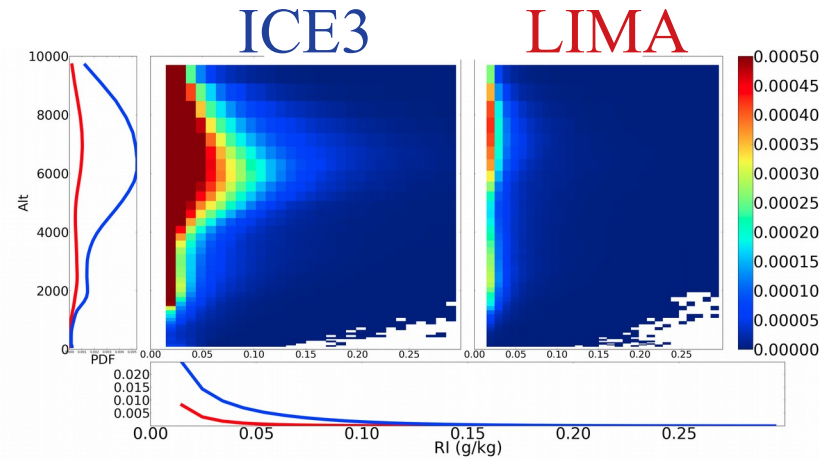


Rain

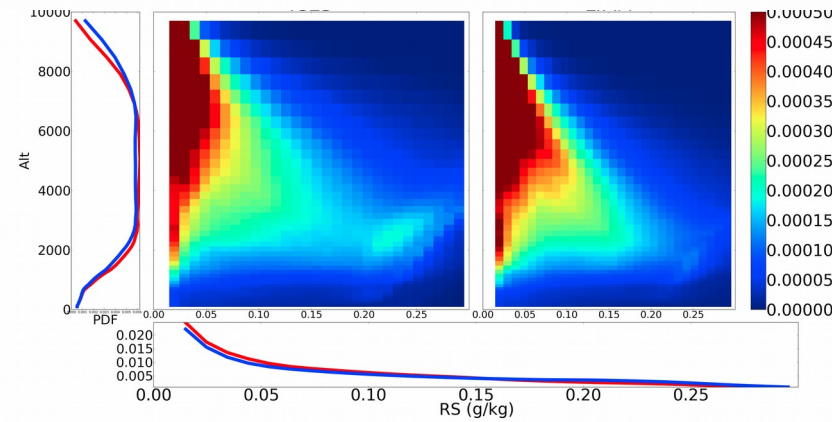


Cold microphysics mixing ratio whole domain (00h, 06h, 12h, 18h, 24h)

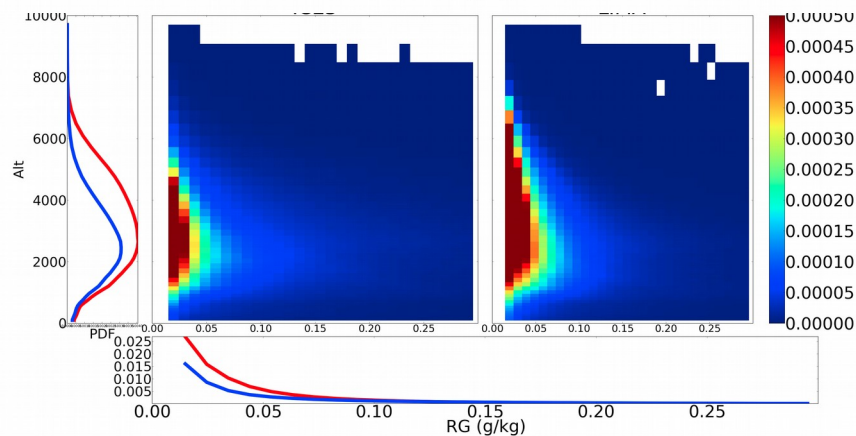
Ice



Snow



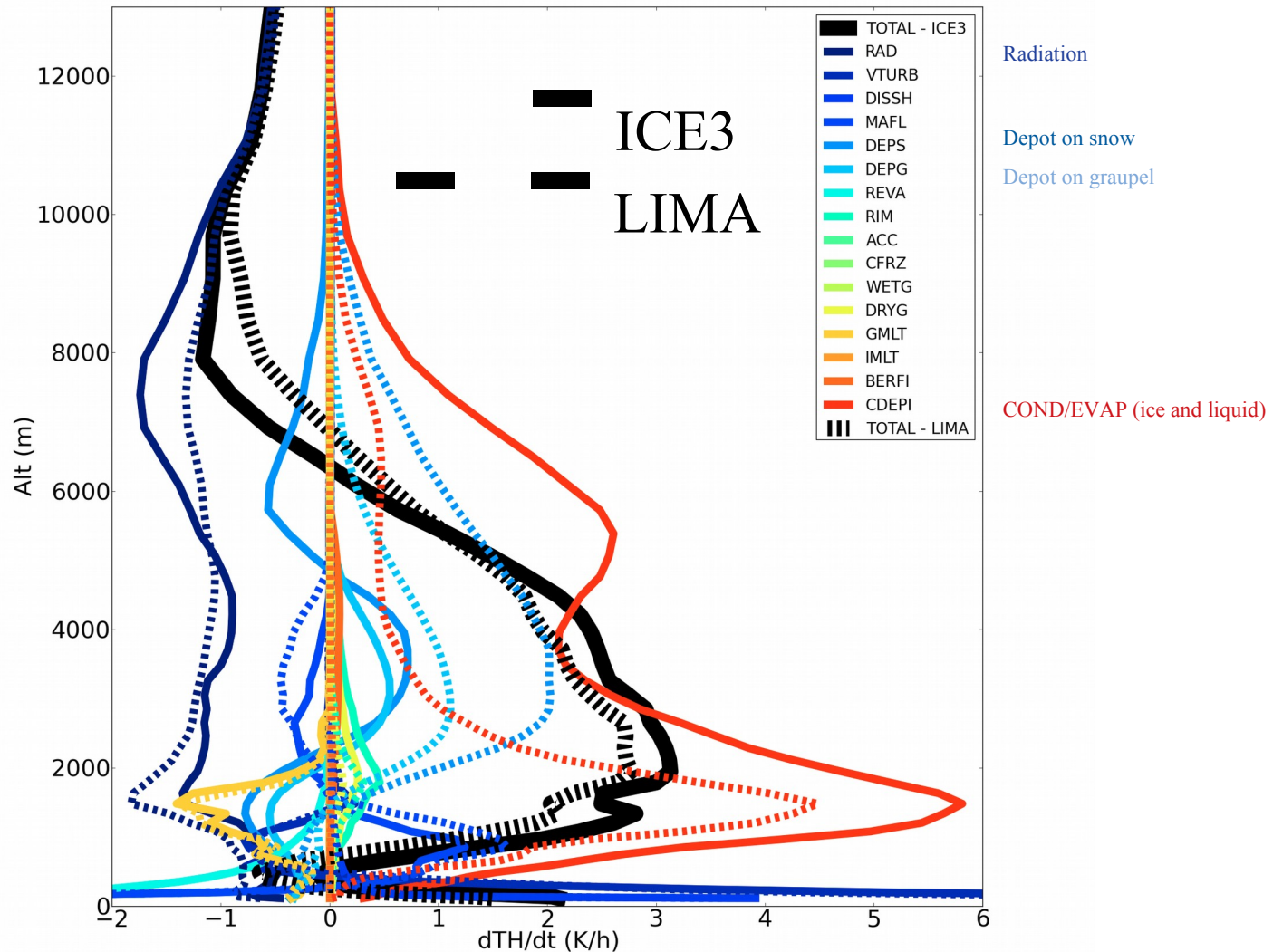
Graupel



Impact on diabatic heating ??

TH budget on the whole domain

03/10/16 00:00



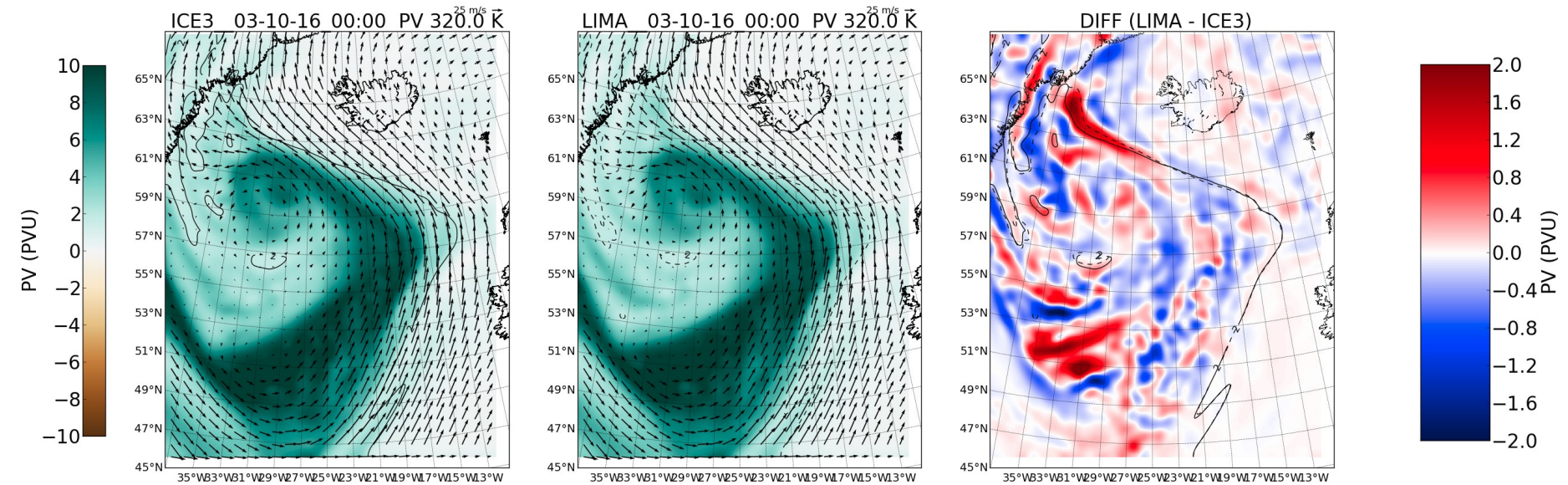
➡ TH dipole less pronounced with LIMA (and associated with different processes)¹⁴

Impact on altitude anticyclone – PV at isentrope 320 K – 03/10/16 - 00:00

ICE3

LIMA

LIMA-ICE3



— 2PVU ICE3
- - 2PVU LIMA

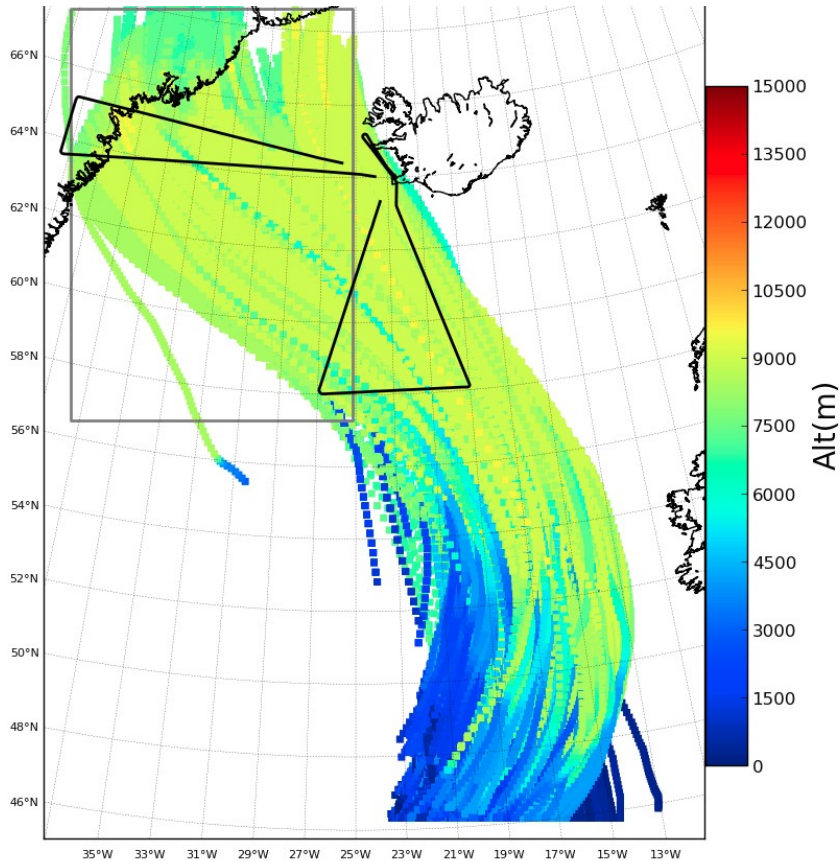
➡ LIMA 2PVU NE shift (2°)

WCBs (300hPa elevation in 24h, Jood and Wernli, 2012)

BACKWARD TRAJECTORIES : 168 960

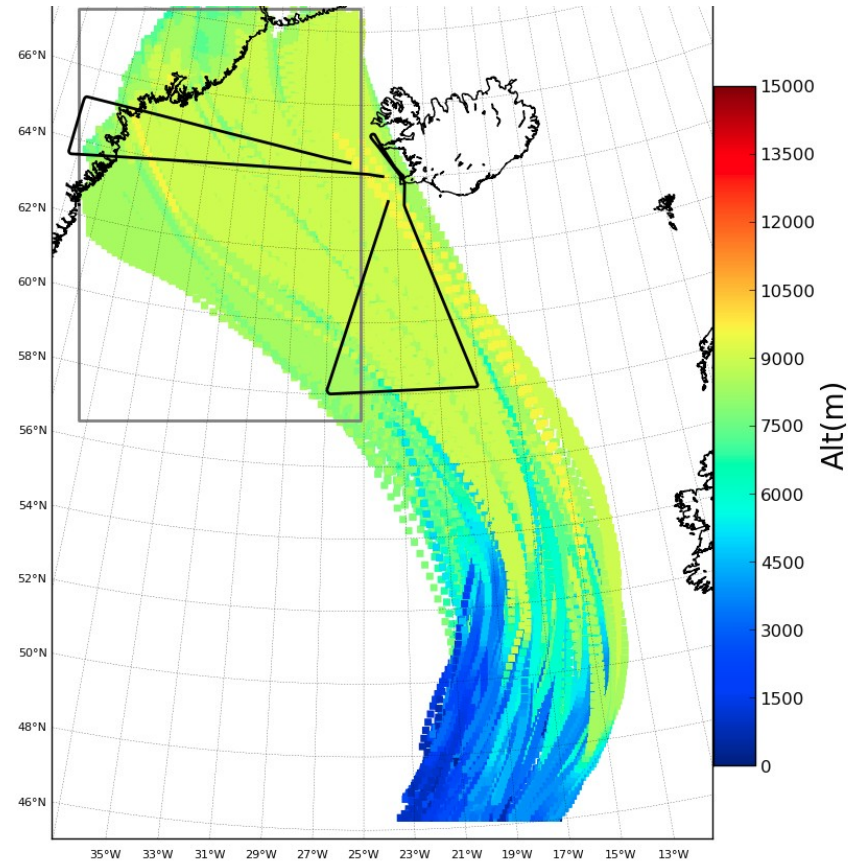
➔ Trajectories intersecting isentrope 320 (+-5) K on 03/10/16 00:00

ICE3



WCB : 3 060

LIMA



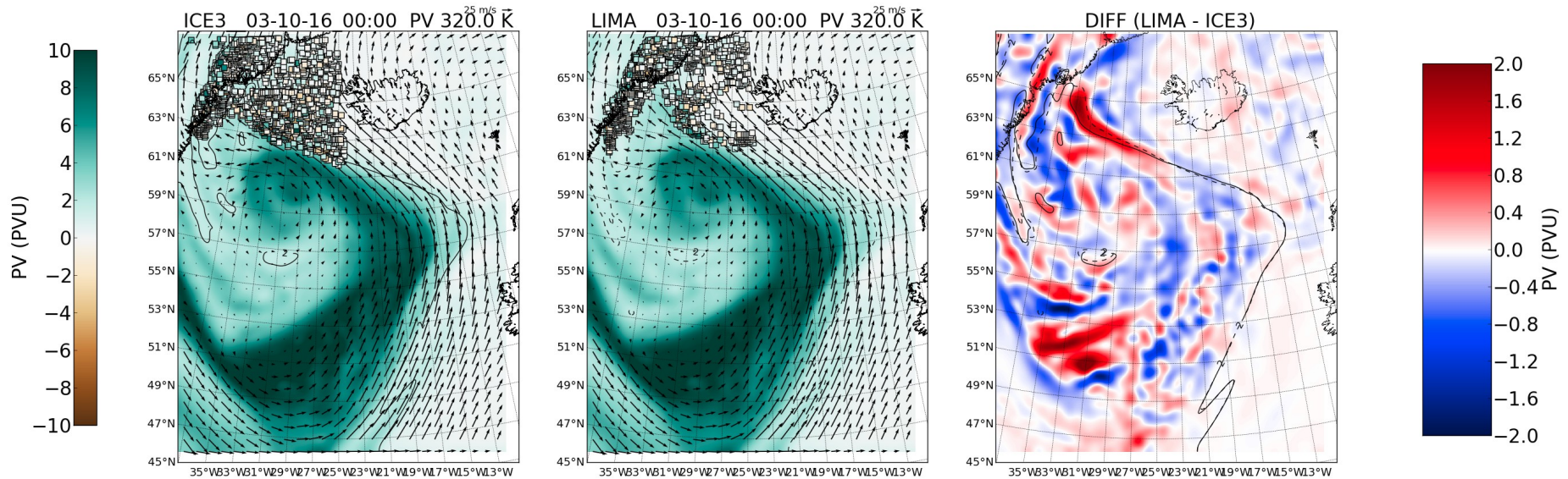
WCB : 1 719

Impact on altitude anticyclone – PV at isentrobe 320 K – 03/10/16 - 00:00

ICE3

LIMA

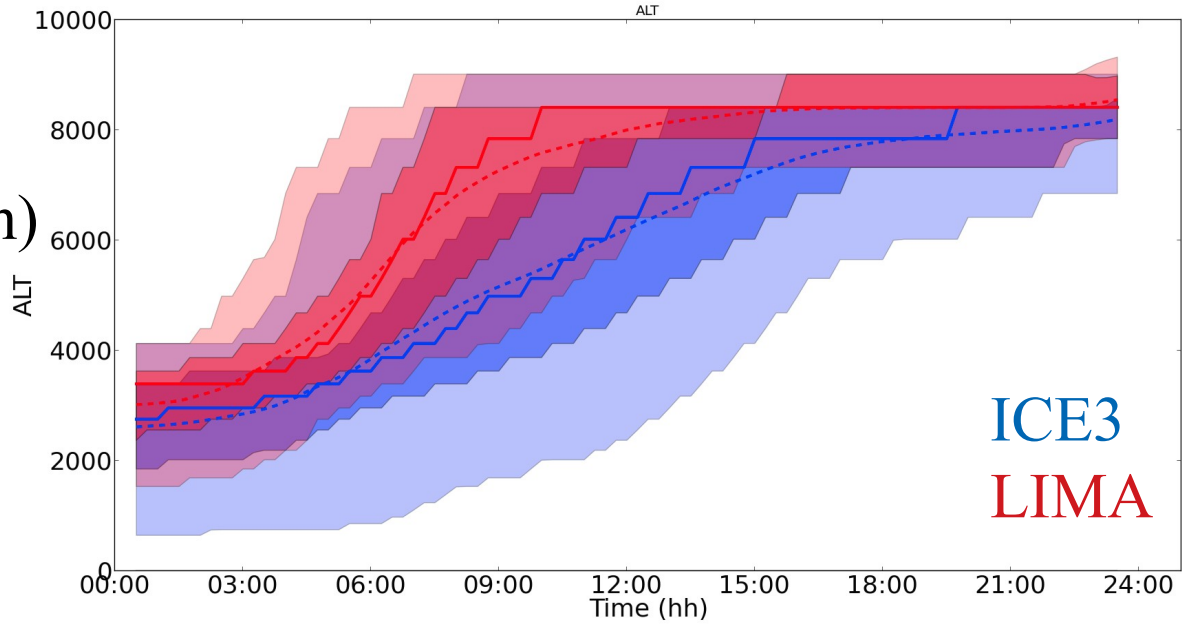
LIMA-ICE3



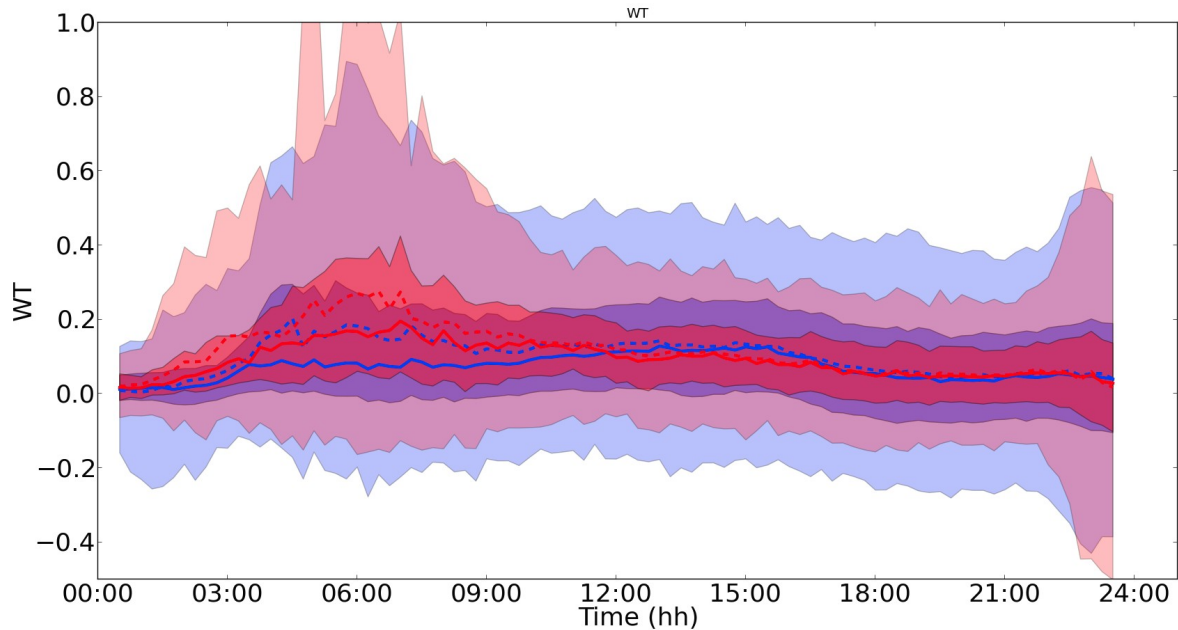
➔ **LIMA** less intersection with isentrobe 320K, **WCB** less active
Less reinforcement of the altitude anticyclone

Temporal evolution – Altitude and vertical velocities

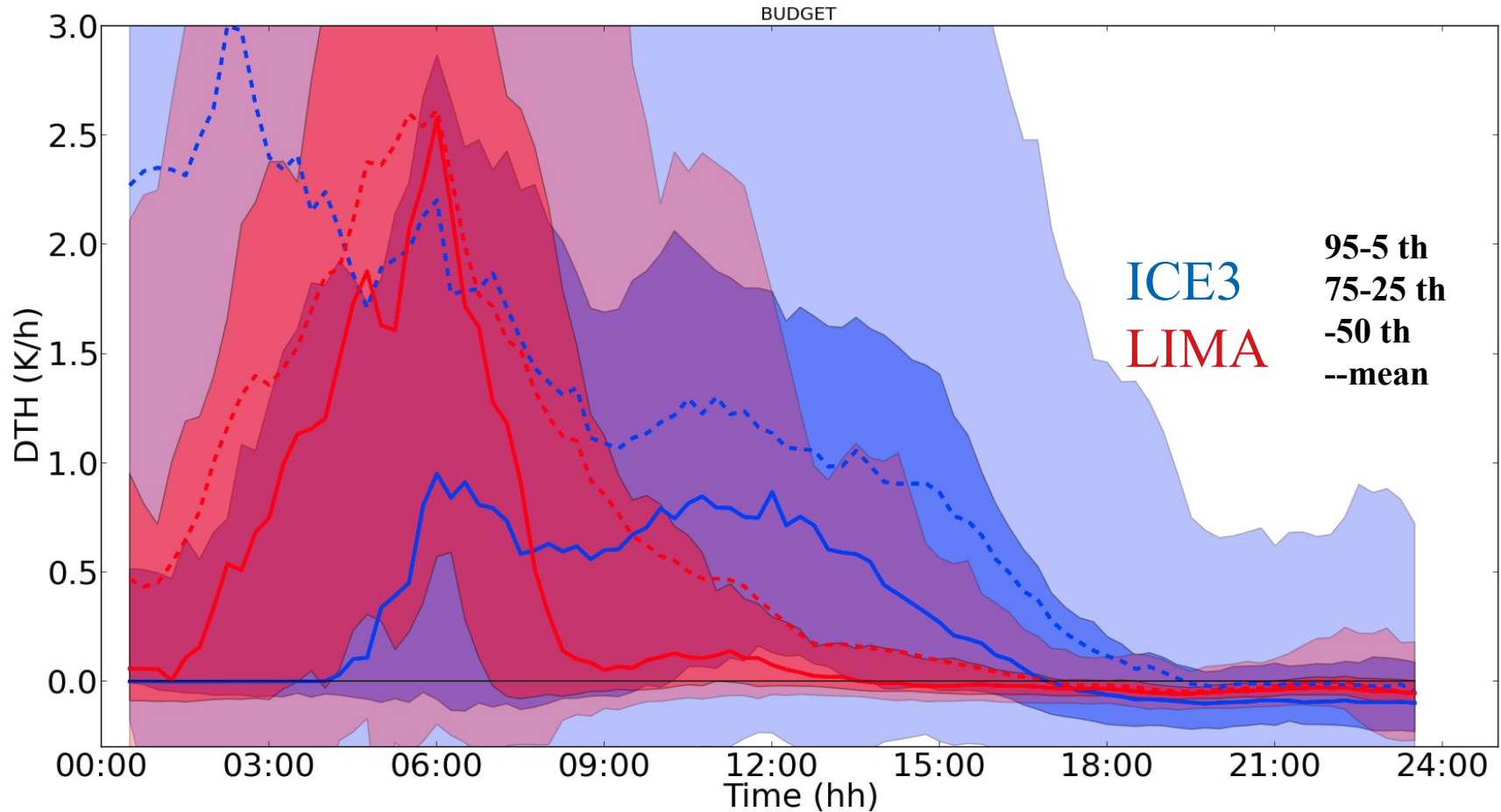
Altitude (m)



W (m/s)



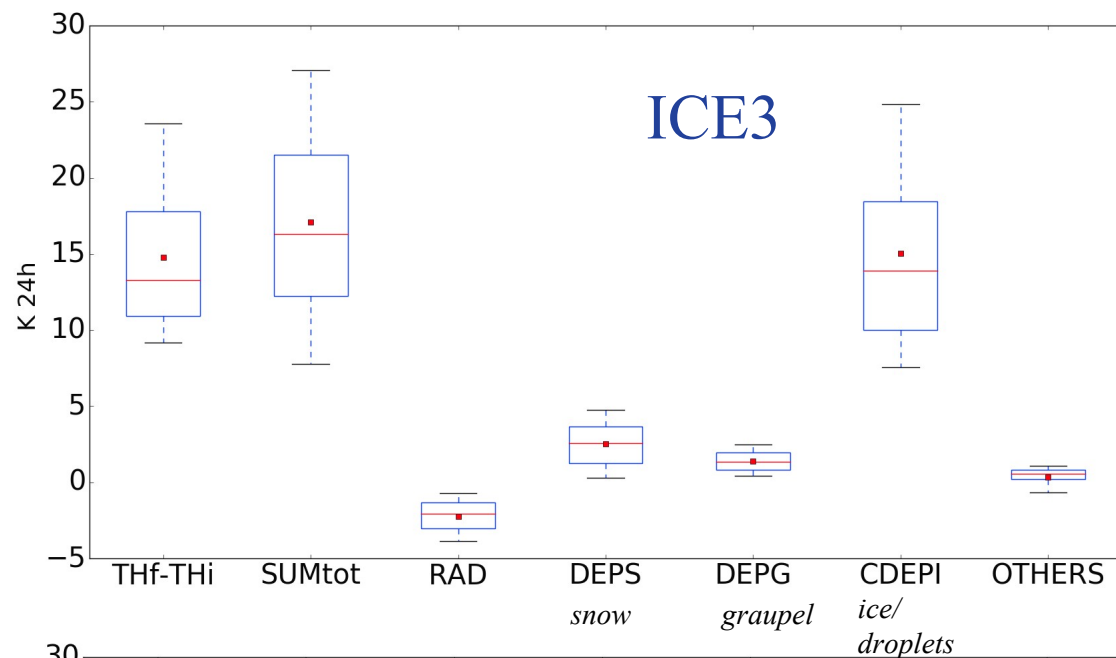
Temporal evolution – Potential temperature budget



Two peaks of latent heat release

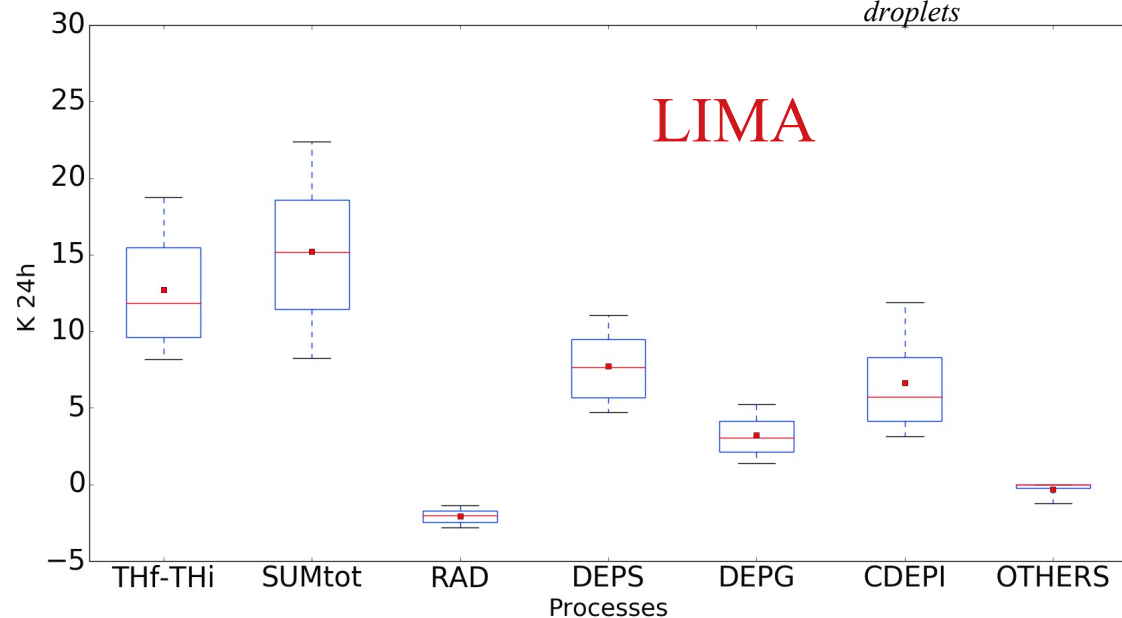
Second peak which may help to ascent more trajectories **less intense with LIMA**

Main diabatic processes (4/16), DTH/Dt : 00-24h

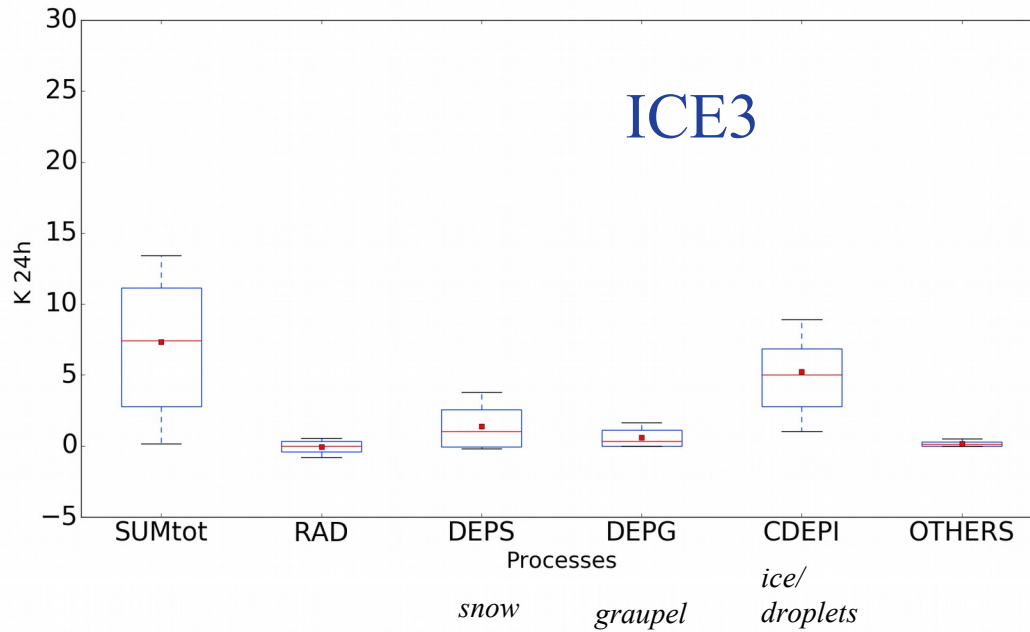


→ **Same main processes but different importances :**

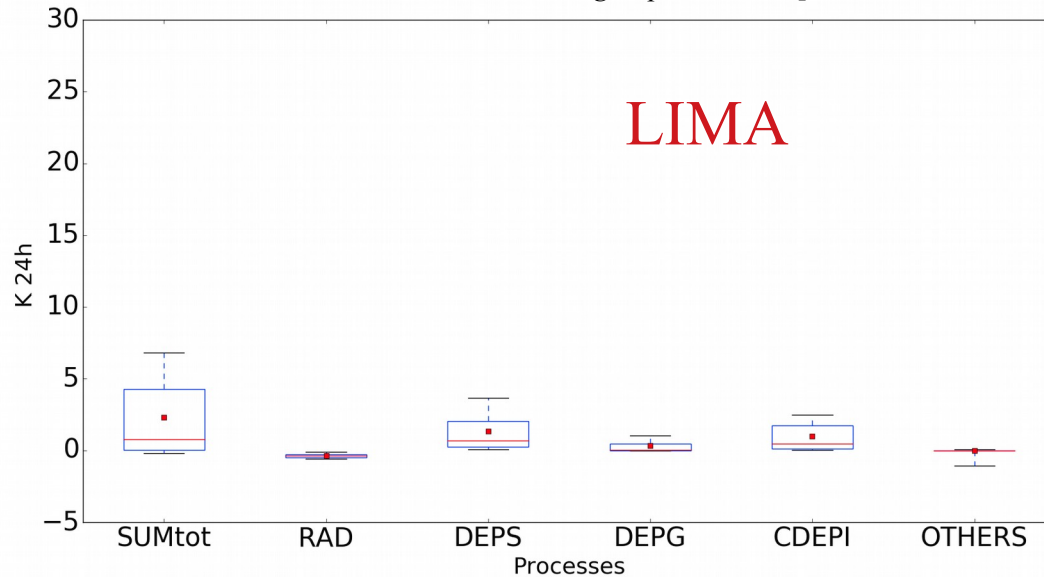
- Radiation
- Deposition on snow
- Deposition on graupel
- **Deposition on ice and droplets**



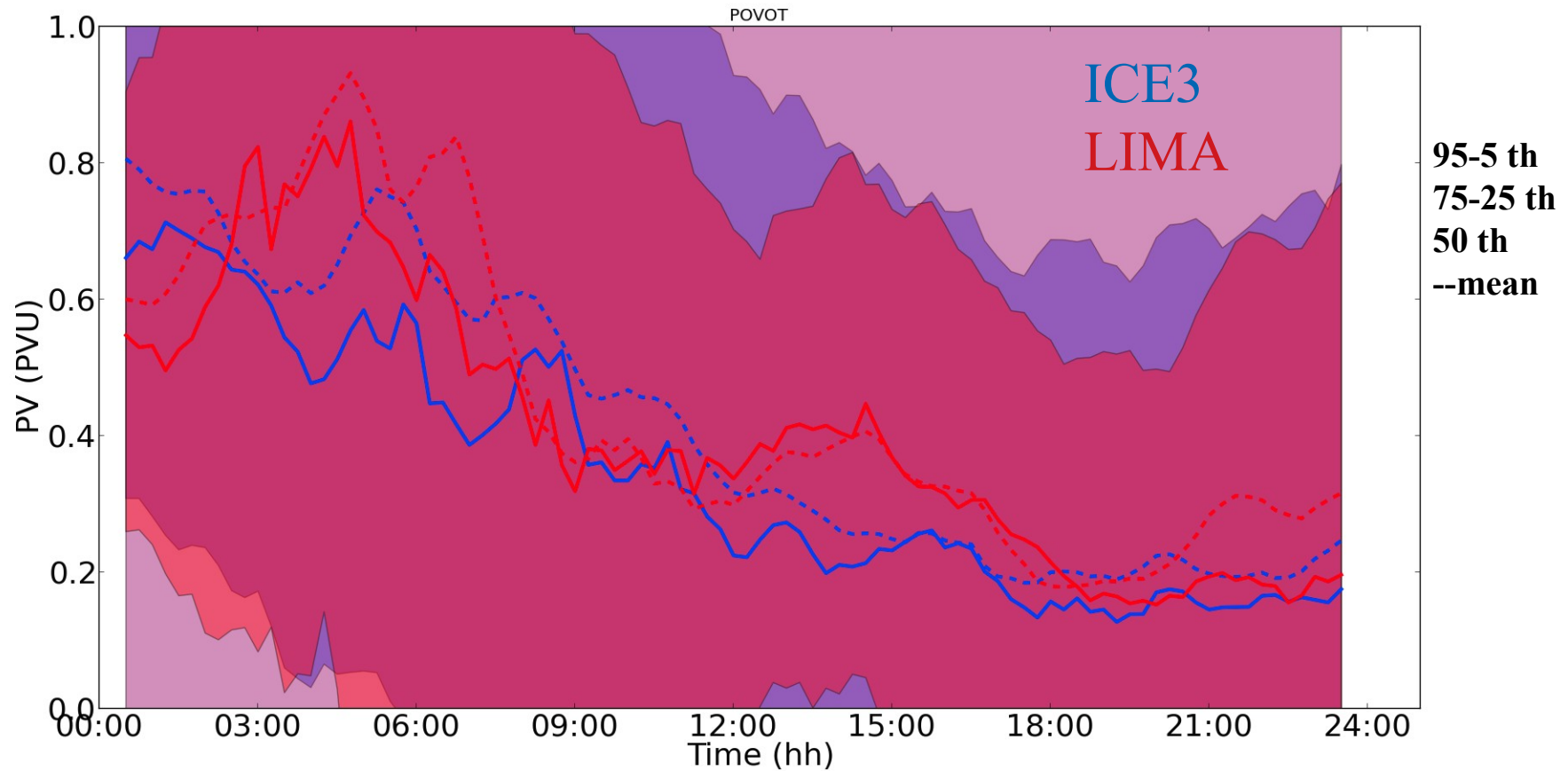
Main diabatic processes (4/16), DTH/Dt : **Second peak – 09:16h**



➔ Second peak with ICE3
due to **deposition on ice**



Temporal evolution – Potential vorticity



➔ LIMA higher PV after 24h (0.1 PVU)

CONCLUSION

➔ **Comparisons of two simulations with two distinct microphysical schemes of the Stalactite Cyclone**

CONCLUSION

- ➔ **Comparisons of two simulations with two distincts microphysical schemes of the Stalactite Cyclone**
- ➔ **→ ICE3/LIMA : Good dynamical representation (Satellite, Radar)**
 - After 24h, **LIMA 2° NE shift** compare to ICE3 for the 2PVU isoline at 320 K
 - **LIMA's WCB less active in number**
 - **LIMA seems to produce less ascent due to deposition on ice at higher altitude**
Same dominant processes for both : Ice and droplets deposition, snow deposition, graupel deposition and radiation with different repartition
 - **LIMA produces less ice than ICE3** on the whole domain and on F7
 - Comparison with reflectivities observed :
 - Both shift on intensity
 - LIMA shape closer to the obs
 - ICE3 intensity in altitude closer to the obs

CONCLUSION

- ➡ **Comparisons of two simulations with two distincts microphysical schemes of the Stalactite Cyclone**
- ➡ → **ICE3/LIMA : Good dynamical representation** (Satellite, Radar)
 - After 24h, **LIMA 2° NE shift** compare to ICE3 for the 2PVU isoline at 320 K
 - **LIMA's WCB less active** in **number**
 - **LIMA seems to produce less ascent due to deposition on ice at higher altitude**
Same dominant processes for both : Ice and droplets deposition, snow deposition, graupel deposition and radiation with different repartition
 - **LIMA produces less ice than ICE3** on the whole domain and on F7
 - Comparison with reflectivities observed :
 - Both shift on intensity
 - LIMA shape closer to the obs
 - ICE3 intensity in altitude closer to the obs
- ➡ **Intensities of the WCB seems to differ due to the ice production and subsequent vapor deposition**
But small impact on upper level after 24h

PERSPECTIVES

→ Equivalent study on a case with a less intense warm conveyor belt
To have weaker impact of synoptic flow

Any case suggestions ?

→ Test on physics (including terminal fall speed of snow should impact reflectivities) and aerosols concentration with LIMA on this other case (negligeable impact on Stalactite cyclone)

Any other tests suggestions ?

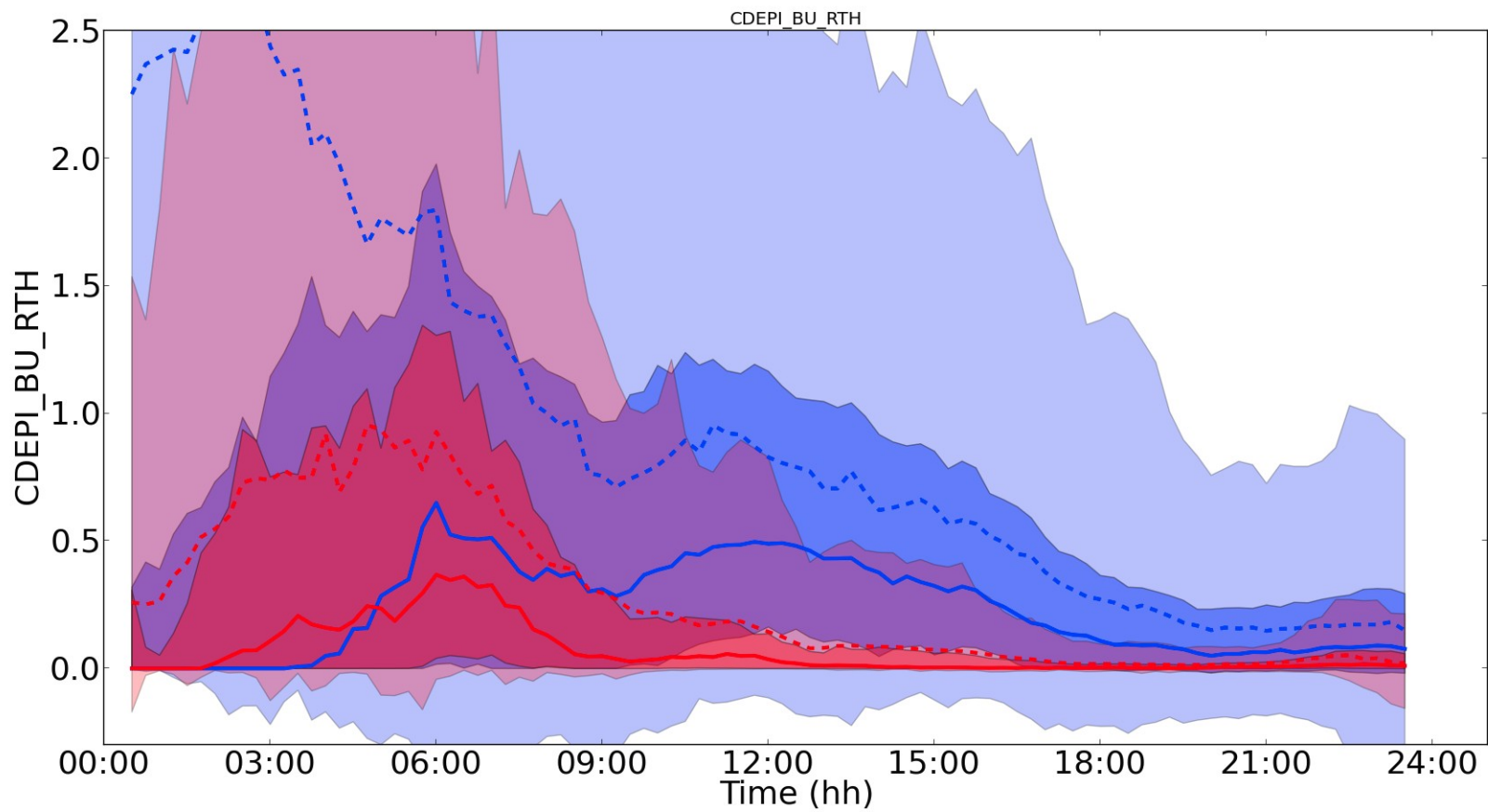
02/10/2016

F6

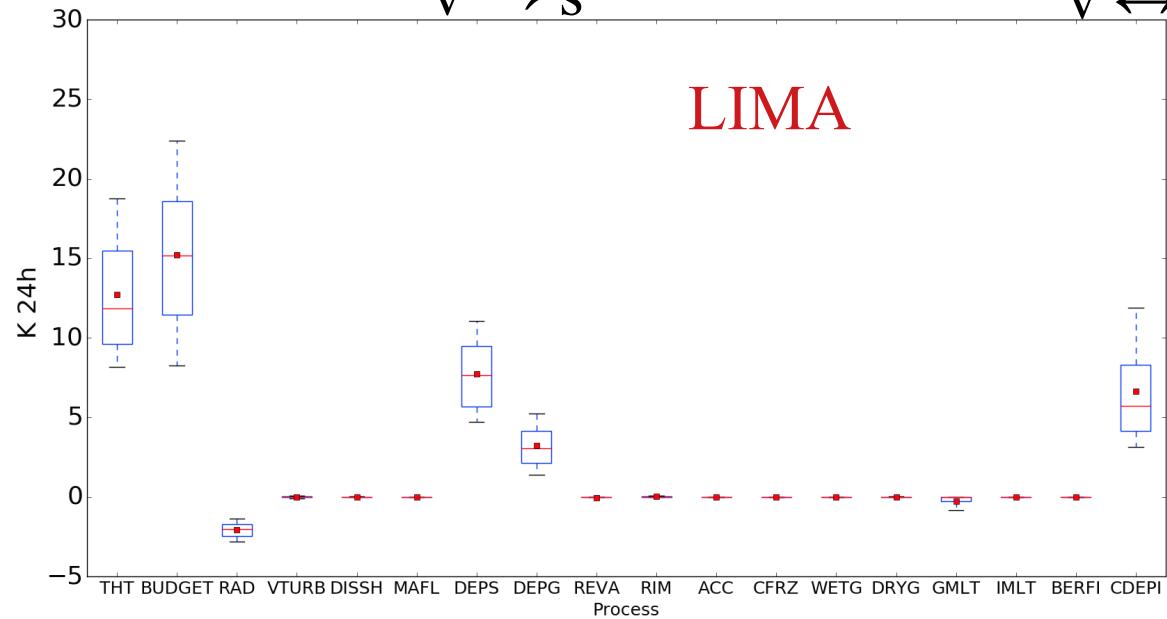
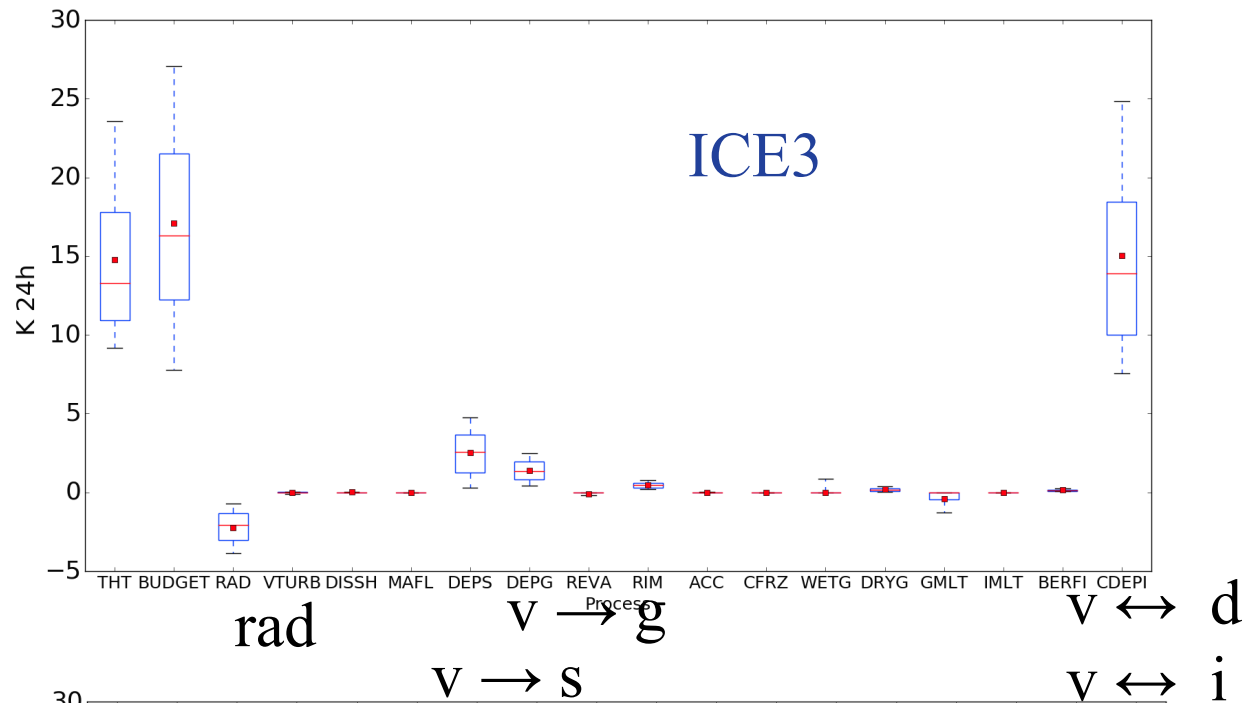
F7

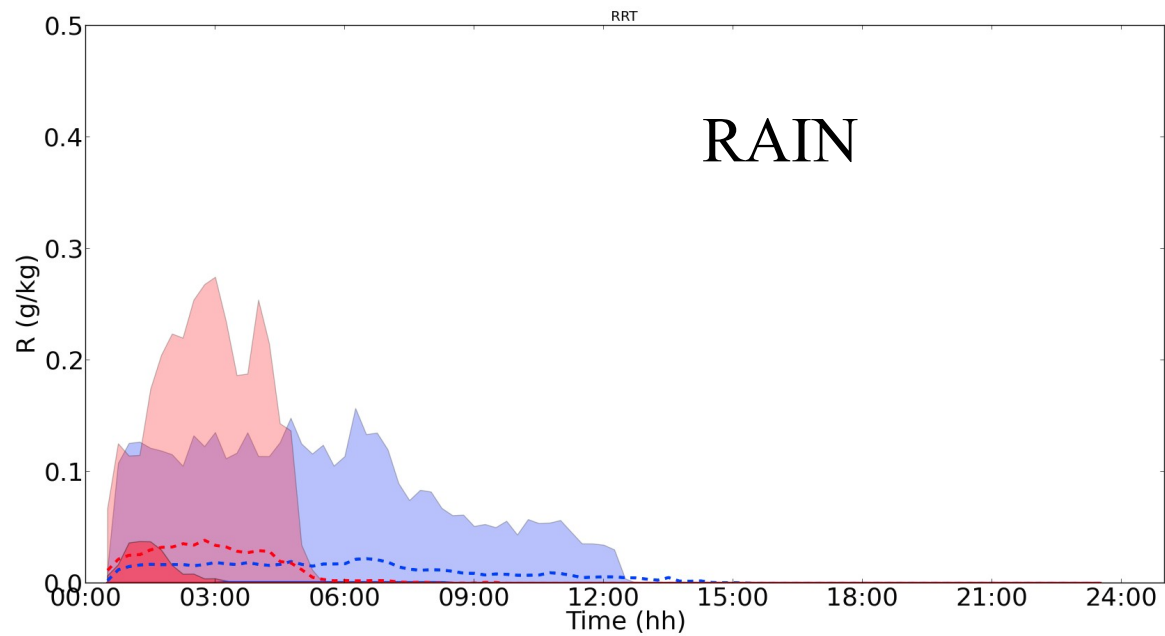
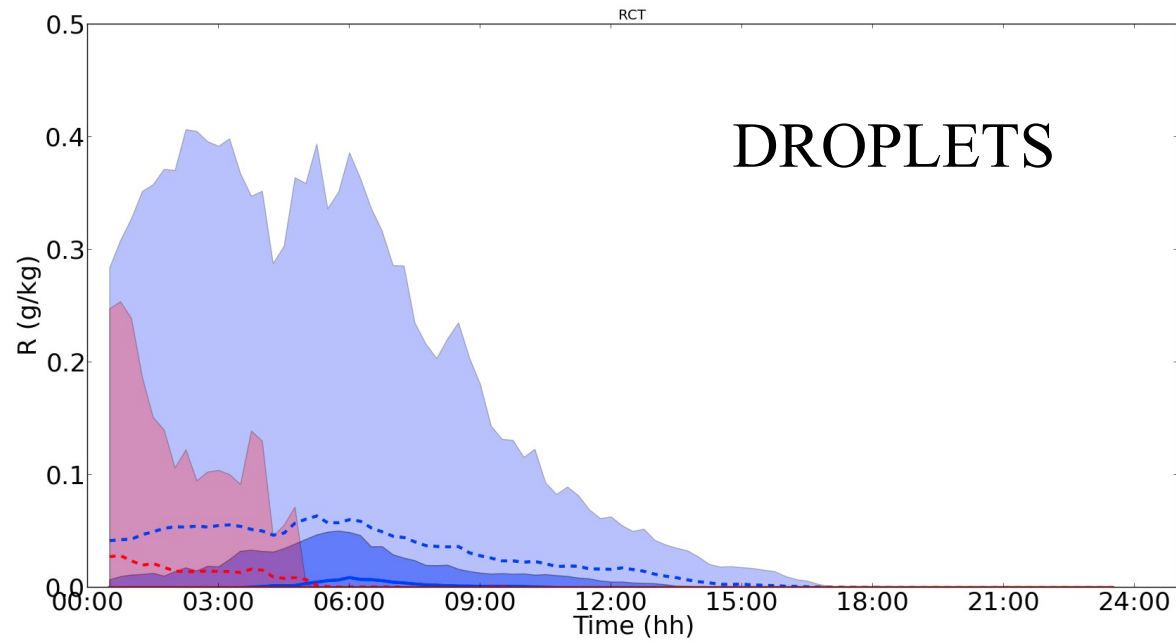
Thank you for your attention

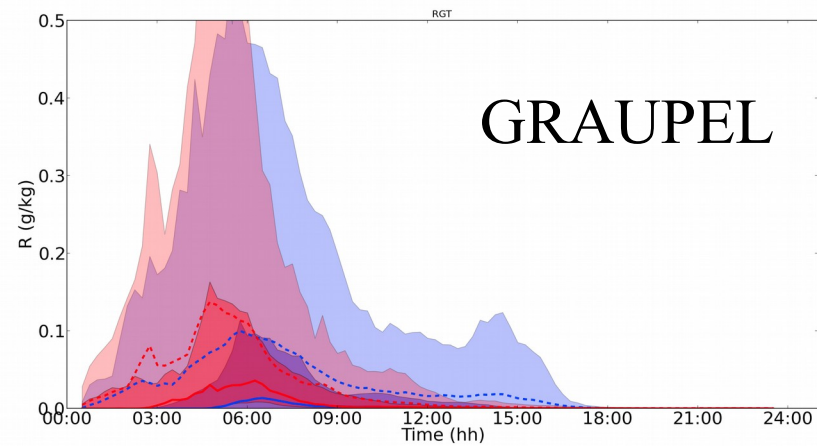
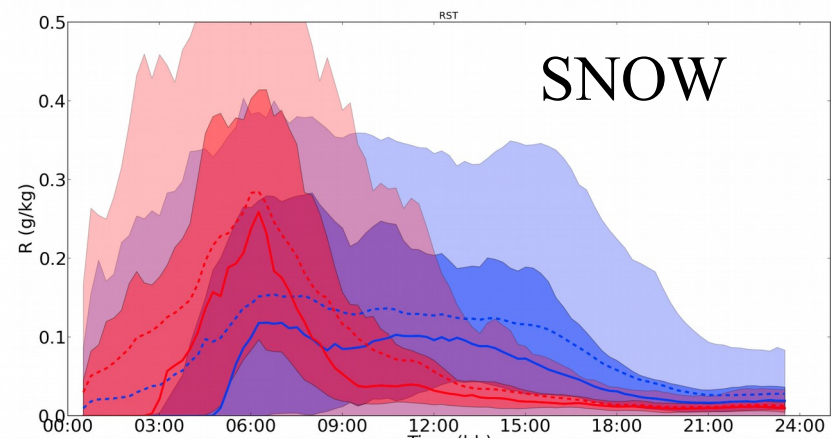
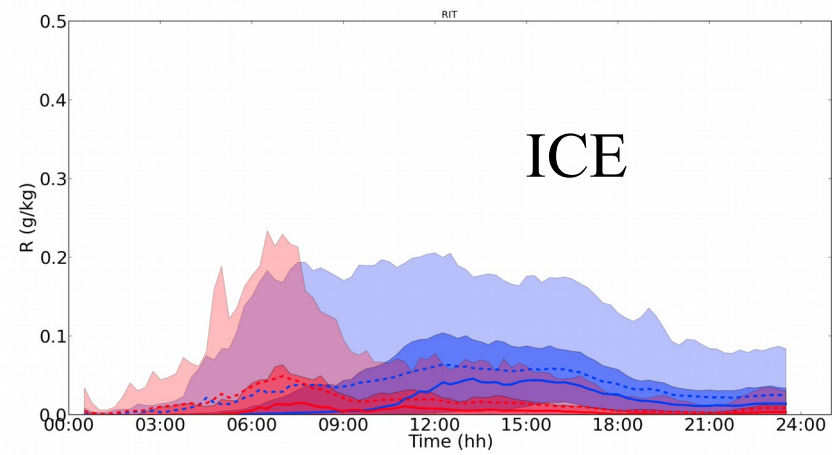
Questions ?



Diabatic processes, DTH/Dt : 00-24h





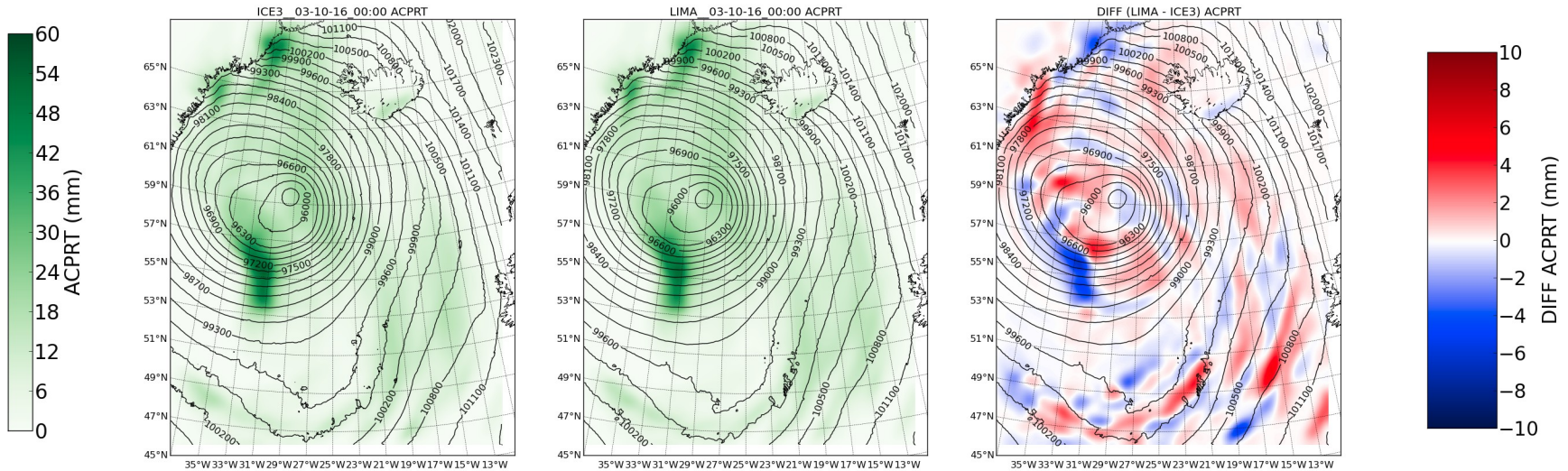


PRECIPITATIONS 24H

ICE3

LIMA

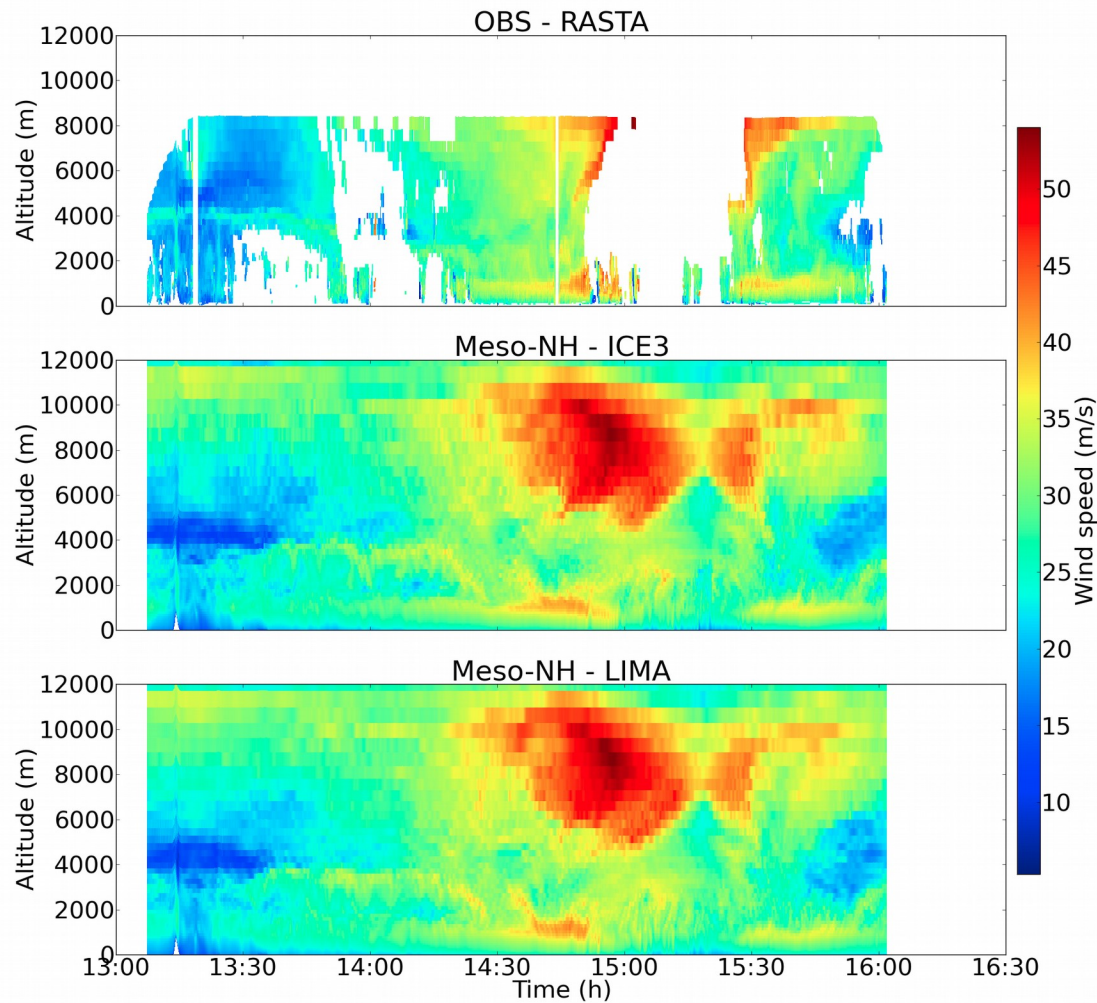
LIMA-ICE3



Isolines Psea (03/10/16 00h)

➔ Digging and structure similar
LIMA more precipitation (+1.8%) (more rain species)

Wind speed - F7



➔ **Slight shift with observation (also on satellite data)**
But correct dynamical representation

ICE3/LIMA, very **similar** horizontal wind speed and structure

Test effectués sur LIMA :

-Tests aux aérosols :

-CCN 100 (CCN 300)

-CCN 1000 (CCN 300)

-IFN 1000 (IFN 10 000)

-IFN 100000 (IFN 10 000)

-Concentration de CCN constante en altitude

—→Plus de glace ?

-Threshold diameter entre la glace et la neige pour LIMA, fixé à 125 μm , mais pourrait atteindre 250 μm selon Zhang, 2014

-Production secondaire de cristaux de glace

-Deff constant pour le rayonnement